

Optimisation of fast likelihood functions for dark matter and rare event searches



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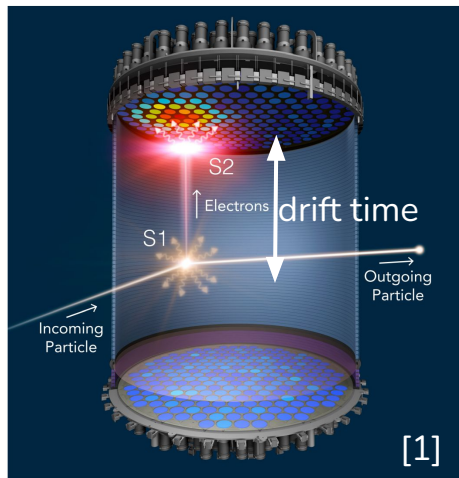
Particular thanks: Robert James
for introduction to this topic and continued help

How does LUX ZEPLIN (LZ) search for WIMPs?

Observables:

S2/S1 sizes - Electron vs **Nuclear** Recoil(ER vs NR)
Radius/drift-time - some further discrimination

Use Models **Monte Carlo (MC)** to produce probability functions **usually just in S1/S2**



Calibration data

WS data

NEST
Noble
Element
Simulation
Technique

Modelling detector
Detector specific package

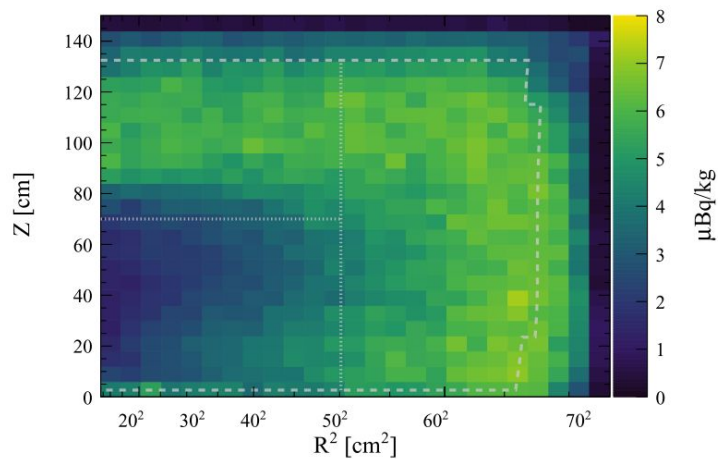
Tuning

Fit to data with likelihood model and Hypothesis test (**inference**)

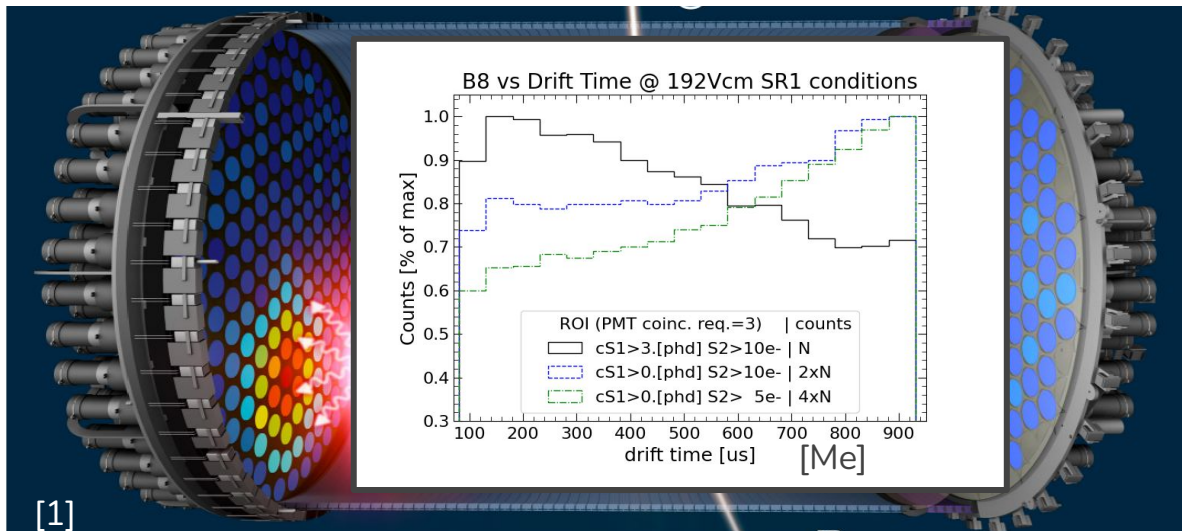
Tune by comparing MC to data

Why have a multidimensional model?

- Backgrounds: Inferred spatial distribution of dominant background of lead-214, tagged by its progenitor polonium.
- Detector effects: Low energy NRs like 8B solar neutrinos coherent nuclear scatters have drift time dependence from light collection efficiencies



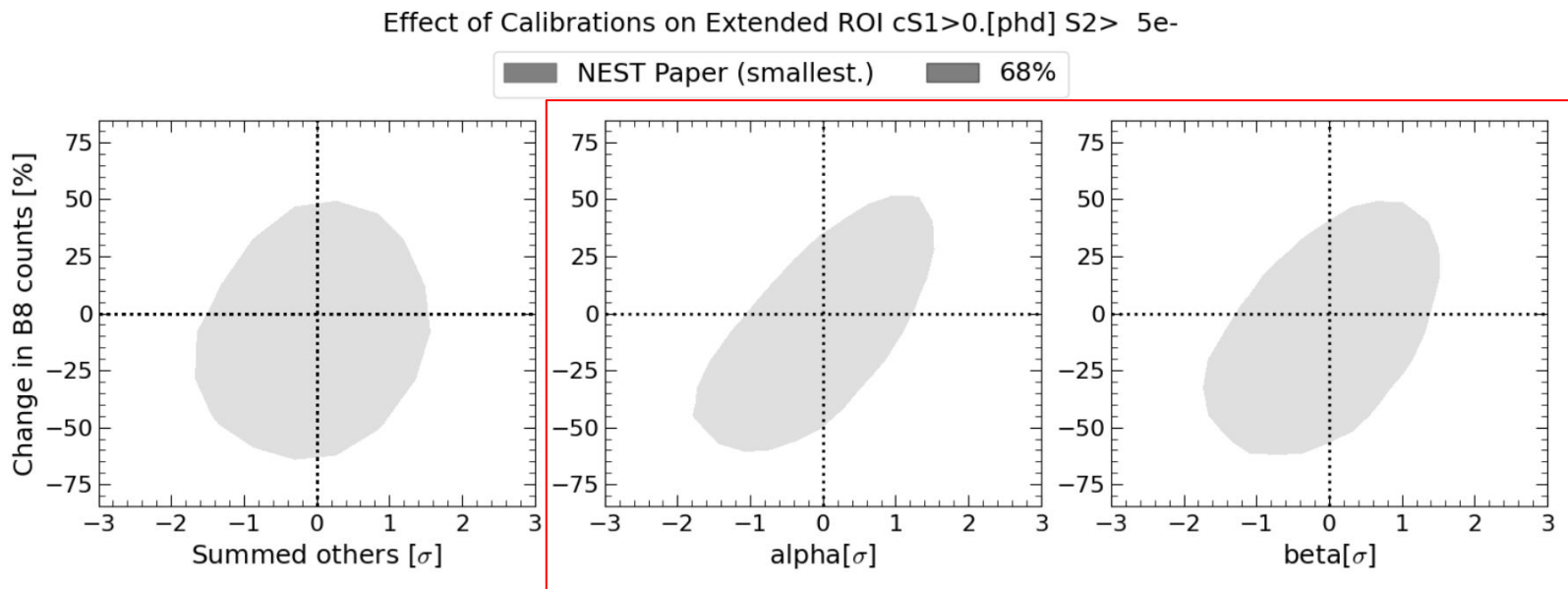
[2] (a) Observed ^{218}Po Distribution



[1]

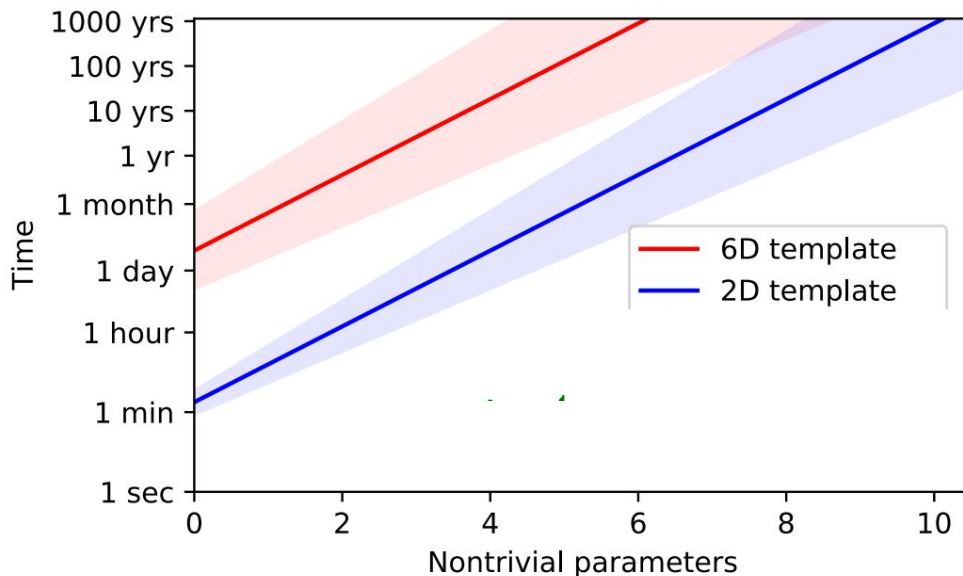
Why have a model with shape varying parameters?

- Acceptance driven by shape variation around boundary cuts
- Significant shape uncertainty gives rate uncertainty
- Default NEST parameters' uncertainty are significance
 - Calibrations tell us more than this!



Why only S1/S2?

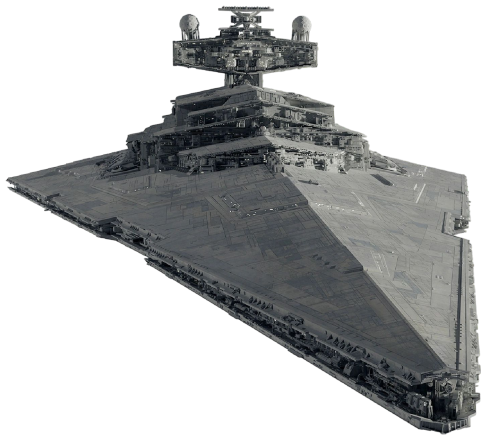
If we want full multidimensional fits w or w/out shape varying nuisance parameters templates won't cut it but flamedisx will



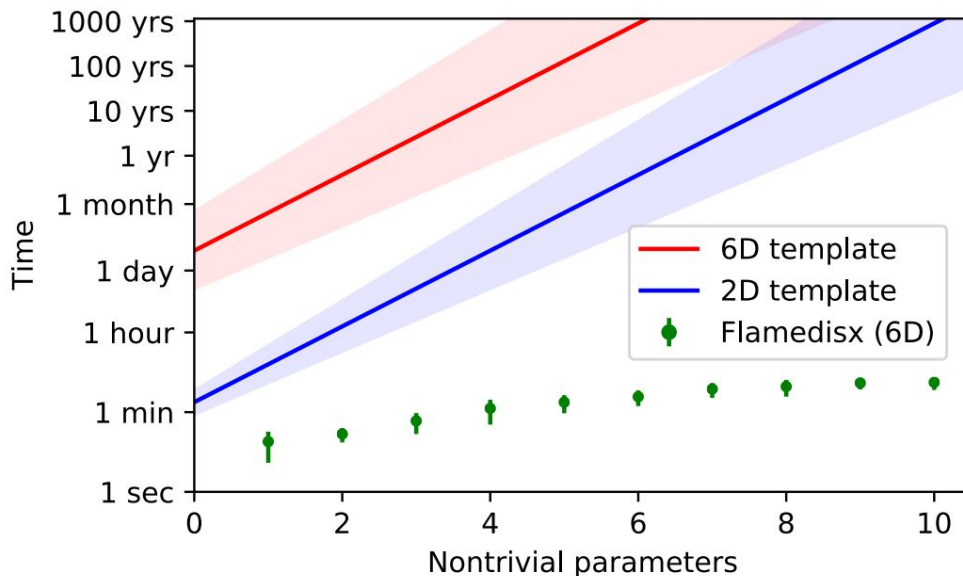
[3] Fitting to $O(1000)$ ER data set.
7 templates for each parameter (anchor points)

Why only S1/S2?

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[Imperial I class star destroyer](#)



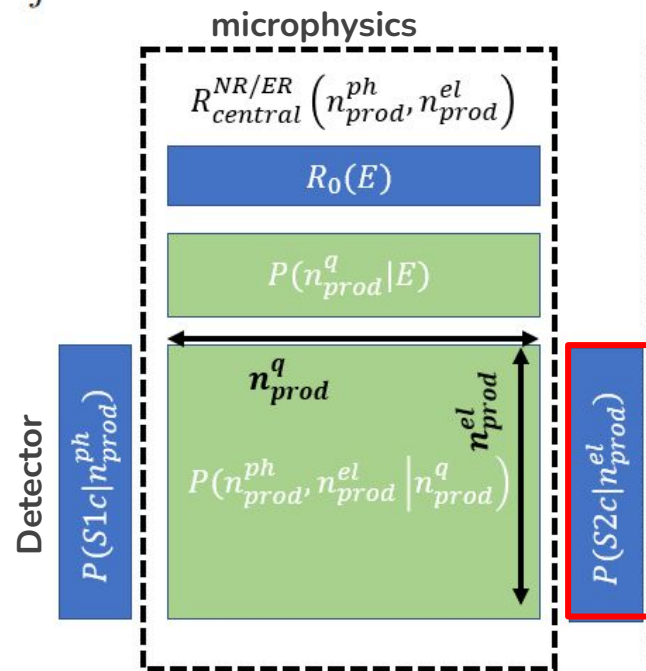
[3] Fitting to $O(1000)$ ER data set.

7 templates for each parameter (anchor points)

Using flamedisx

$$\ln(L(\vec{\theta}, \{\vec{s}_i\})) = -\mu(\vec{\theta}) + \sum_i^{\text{events}} \ln\left(\sum_j^{\text{sources}} R^j(\vec{\theta}, \vec{s}_i)\right) + \text{const.}$$

Guess all underlying parameters that significantly contribute to data

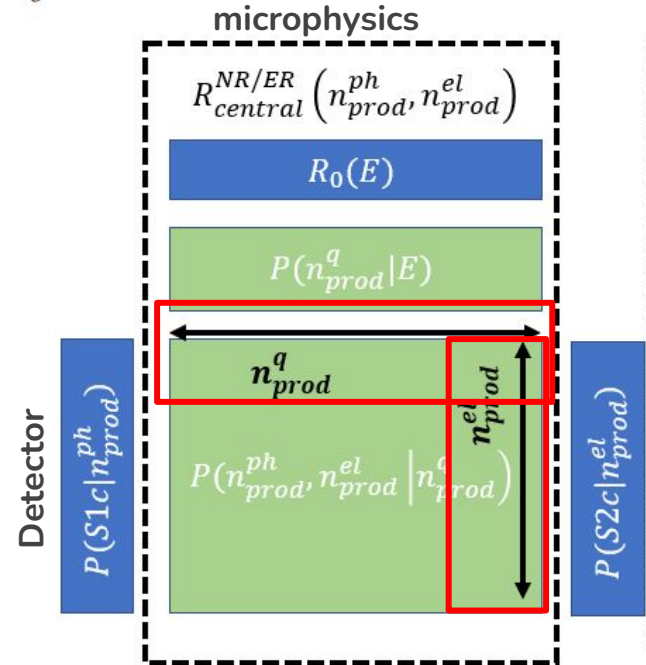


Using flamedisx

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Guess all underlying parameters that significantly contribute to data

Explicitly evaluate the **differential rate** on those parameters



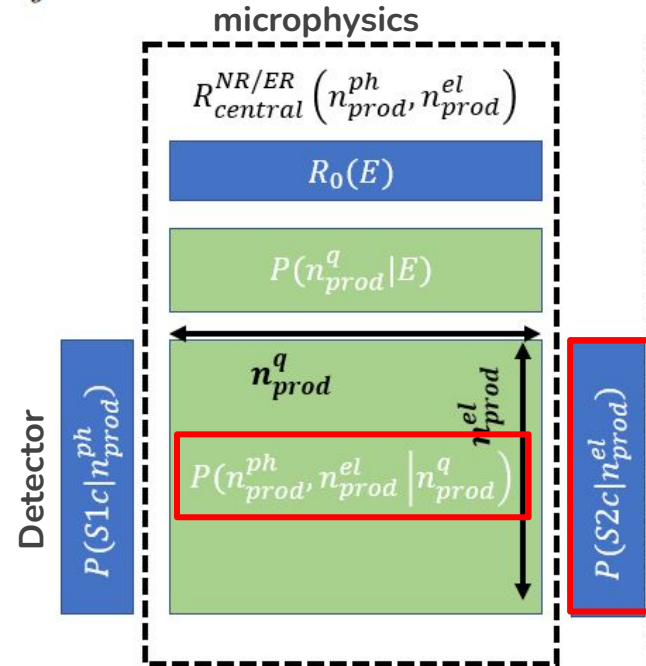
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Guess all underlying parameters that significantly contribute to data

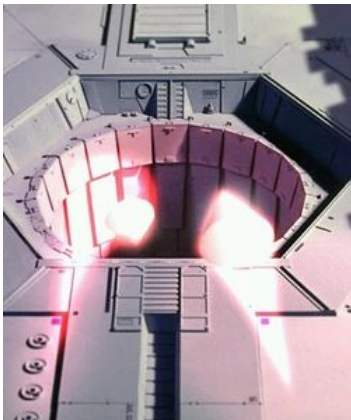
Explicitly evaluate the **differential rate** on those parameters

Treat as tensor operation and utilise differentiable programming



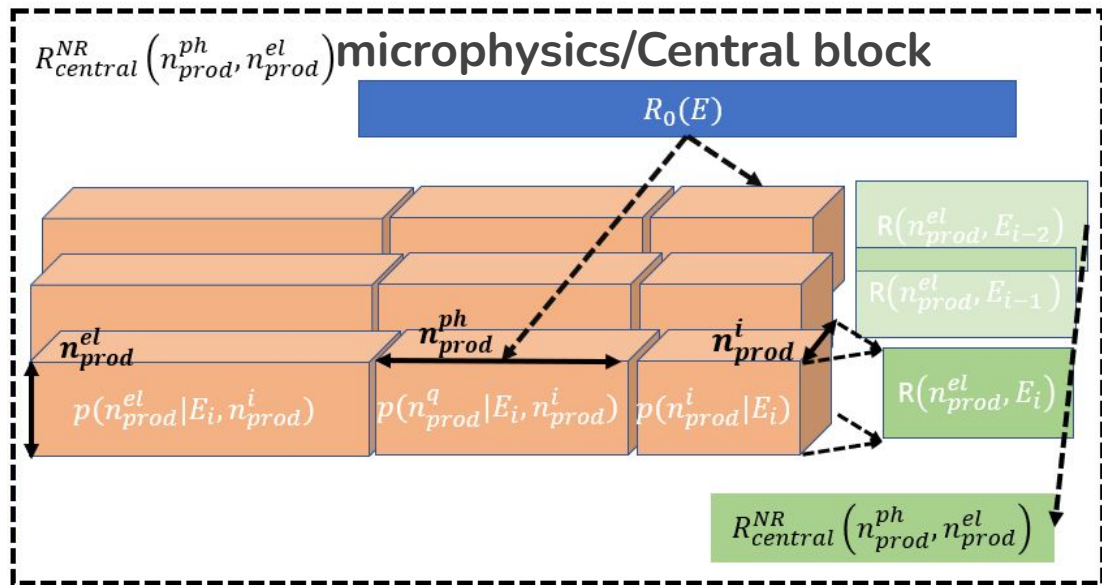
Implementing NEST -> FlameNEST [4]

Convolutd yield models



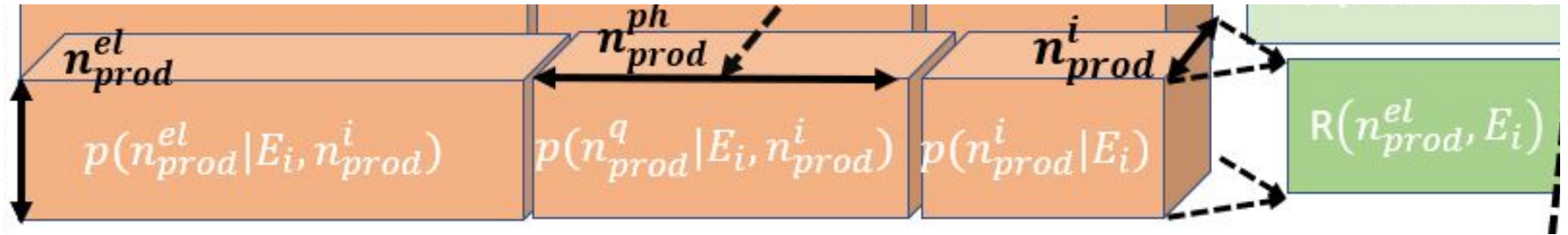
OOM memory failed to allocate

Developers implemented the NEST models and caused performance issues



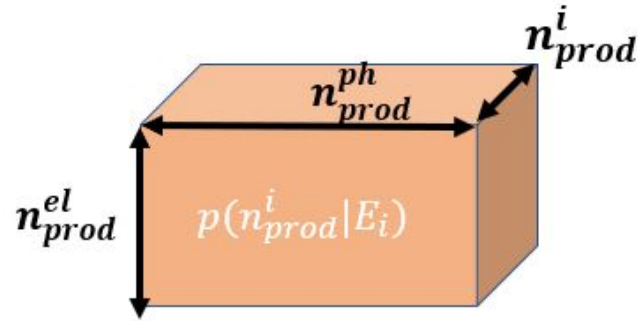


Fixing the problem



- Each block represents a tensor
- Each dimension of the block is the range of underlying parameters
- Each function is evaluated for every element of that block

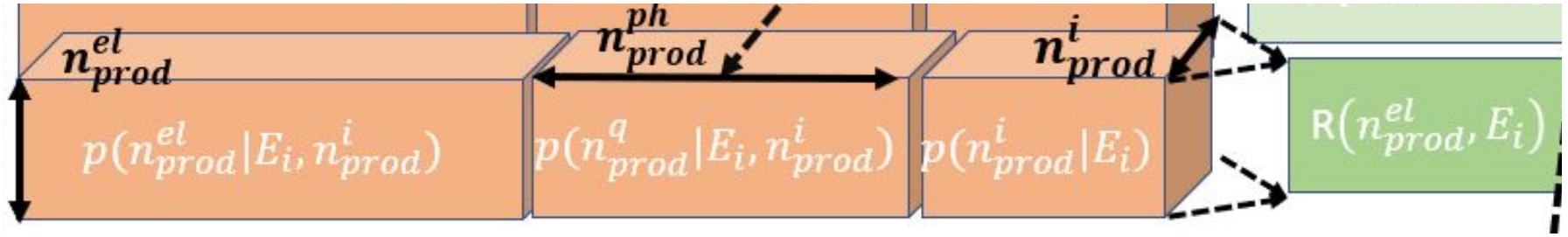
Degenerate dimensions



- The model function that represents recombination only depends on ions produced
- It is being evaluated on a tensor of ions, photons, and electrons
- Many degenerate evaluations of the model

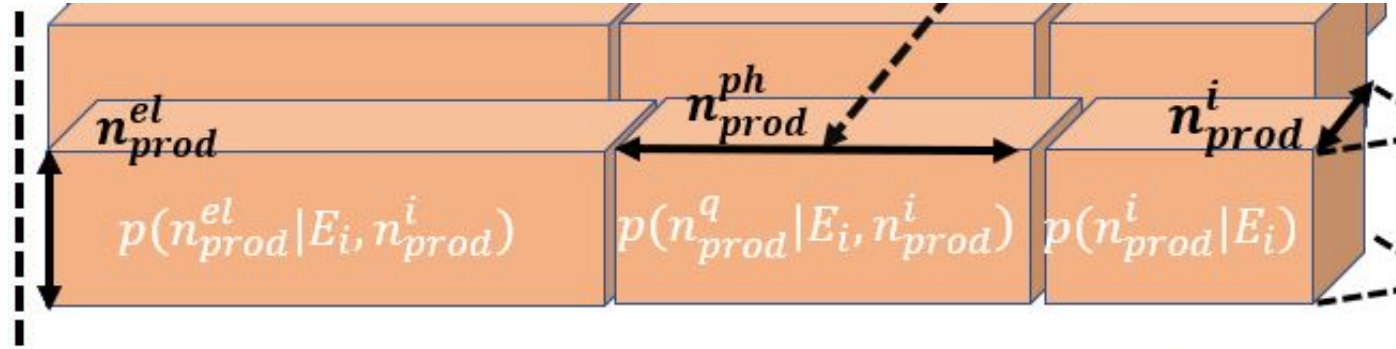


Degenerate dimensions



- Every model here has degenerate dimensions in this way
- Each evaluation of a function in differentiable programming represents a graph of primitive functions
- Consumes a lot more memory than just the value of the function

Fixing this problem



- To fix this problem I **carefully** implemented **unique and gather** to calculate the model functions.
 - Careful as these functions can cause performance issues
 - **Only use when significantly reduces degeneracy**
- Photons not explicitly in the model but quanta=photons+electrons

Explicit profiling results

Differential Rates	Before				After			
	Peak Memory (GiBs)	Trace time (MM:SS)	ex.time (SS)	Top Operations	Peak Memory (GiBs)	Trace time (MM:SS)	ex.time (SS)	Top Operations
det.param g1 batch size 5	12.0	02:05	02	yield tfp functions tensor ops	2.0	08:16	02	gather/tensordot tensor ops
yield.param α batch size 1	28.0	04:13	02	yield tfp functions and gradients of them	1.0	04:25	01	gather/tensordot tensor ops

Reduction of 6/28x of memory usage for detector/yield parameters.

- =6/28x **speed up** as can processes more events simultaneously

Memory dominated by tensor manipulation instead of model functions

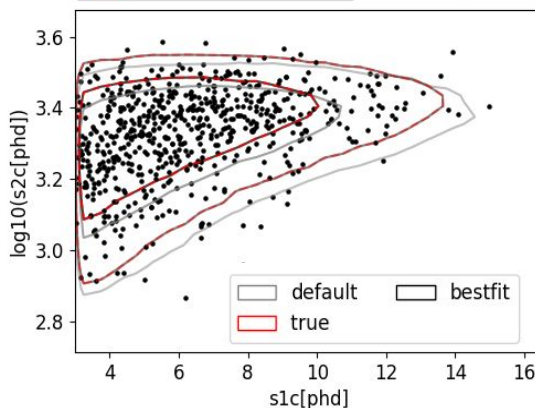
- Weaker scaling of memory with parameters= can float many more parameters

Tracing time does increase but execution time same/smaller

- Negligible as long as batch size \ll data size.

Testing with simulated detector (public LZ information)

- Using a test low energy flat nuclear recoil source:
- Time: 11-14mins to fit
 - 30mins to generate total rate estimator.

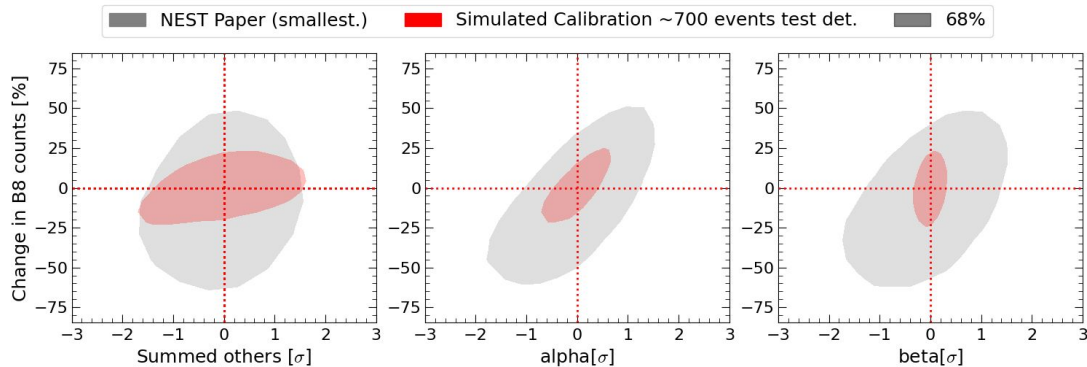


	predicted	True	bestfit	scale
$\mu_{r.m}$	n/a	2.00	2.00 ± 0.09	10^0
α	1.10 ± 0.05	1.11	1.11 ± 0.02	10^1
β	1.10 ± 0.05	1.08	1.08 ± 0.01	10^0
ϵ	1.26 ± 0.29	1.07	1.04 ± 0.03	10^1

- Accurately finds the distribution!**

- Auto-differentiation gives covariance and uncertainties at bestfit
- Significant constraints** with just few number of points

Effect of Calibrations on Extended WS ROI





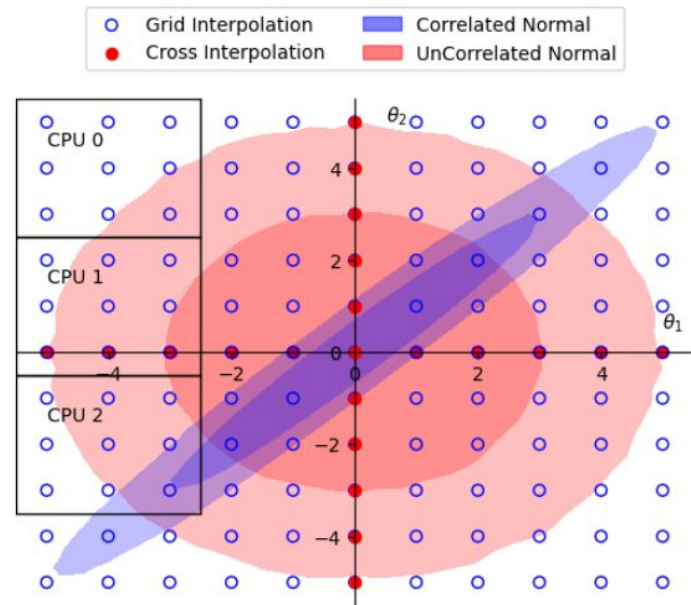
Why is it incomplete?

$$\ln(L(\vec{\theta}, \{\vec{s}_i\})) = -\mu(\vec{\theta}) + \sum_i^{\text{events}} \ln\left(\sum_j^{\text{sources}} R^j(\vec{\theta}, \vec{s}_i)\right)$$

My work focused on the **differential rate** term

Evaluate the **total rate** using simulations of fixed points and interpolate

Still only need total counts so better than full templates



$$\mathcal{O}(n_{\text{parameters}} \times n_{\text{anchors}})$$

$$\mathcal{O}(n_{\text{anchors}}^{n_{\text{parameters}}})$$



Solution: for now

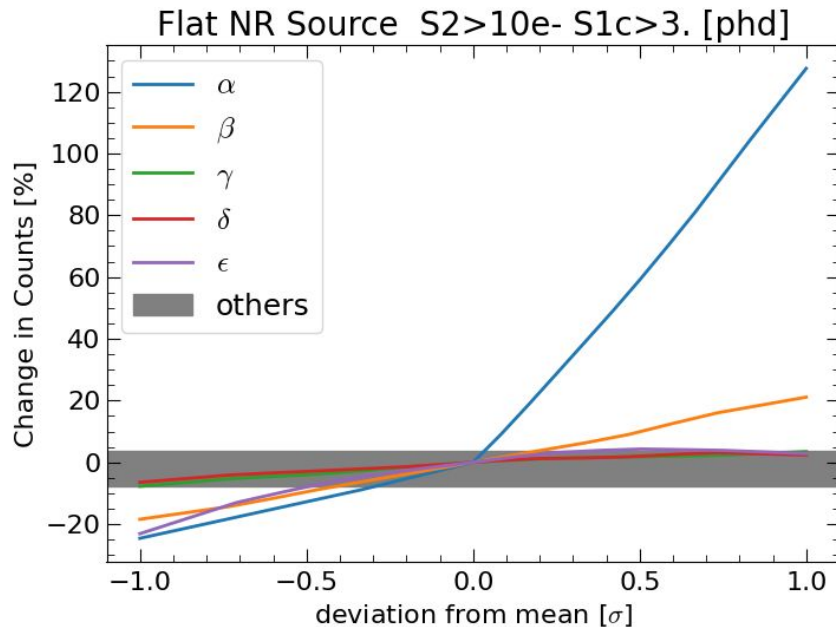
Pick the three biggest impacts

Lots of solutions to explore:

Yield functions are easy to evaluate so:


- Multi-level simulations
- Multi-fidelity simulations
- Creating a better grid
- Reparameterize the model

Or explicit integration with better tools



$$\mu(\vec{\theta}) = \sum_j \mu^j(\vec{\theta}) \quad \mu^j(\vec{\theta}) \propto \int R^j(\vec{\theta}, \vec{s}) d\vec{s}$$

Conclusion

- 
1. Hopefully my plot gore at the start convinced you that
 - a. Position distributions are important
 - b. Shape varying parameters are important
 2. Flamedisx explicitly evaluates differential rate and allows for:
 - a. Multidimensionality
 - b. Shape varying parameters
 3. Implementing NEST caused performance issues:
 - a. **My work fixed those performance issues**
 - b. Shown Multi-dimensional and parameter inside central block possible
 4. Total rate estimators are the next challenge

LZ (LUX-ZEPLIN) Collaboration, 38 Institutions

250 scientists, engineers, and technical staff

<https://lz.lbl.gov/>

- Black Hills State University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- King's College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Los Angeles
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- University of Oxford
- University of Rochester
- University of Sheffield
- University of Sydney
- University of Texas at Austin
- University of Wisconsin, Madison
- University of Zürich



LZ Collaboration Meeting at SURF, June 2023



Science and
Technology
Facilities Council

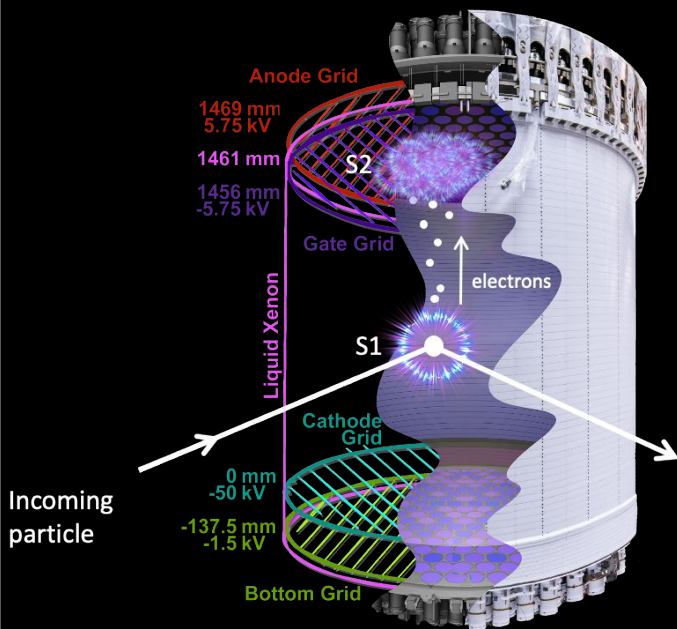


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U.S. Department of Energy
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Bibliography:

[1] LUX-ZEPLIN Technical Design report Arxiv:[1703.09144](https://arxiv.org/abs/1703.09144)

[2]- Background Determination for the LUX-ZEPLIN (LZ) Dark Matter Experiment :
[10.1103/PhysRevD.108.012010](https://arxiv.org/abs/10.1103/PhysRevD.108.012010)

[3] Finding Dark Matter Faster with Explicit Profile Likelihoods
[10.1103/PhysRevD.102.072010](https://arxiv.org/abs/10.1103/PhysRevD.102.072010)

[4] FlameNEST: explicit profile likelihoods with the Noble Element Simulation Technique
[10.1088/1748-0221/17/08/P08012](https://arxiv.org/abs/10.1088/1748-0221/17/08/P08012)

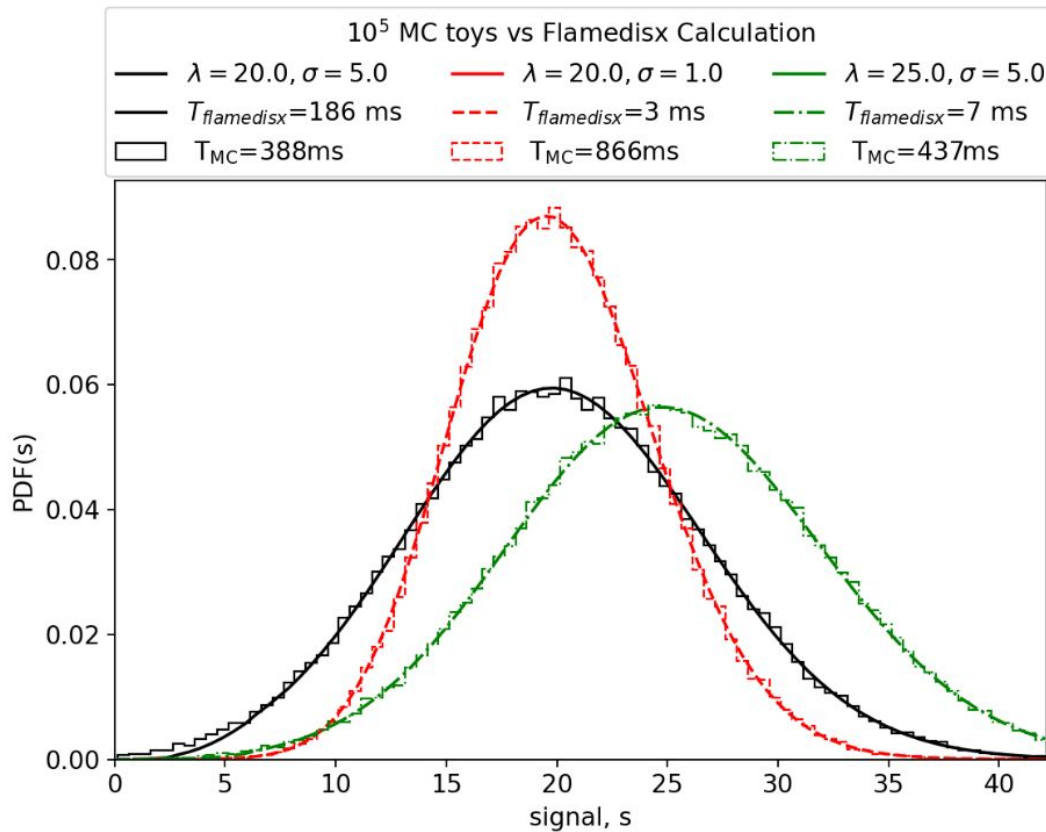
Miscellanea





Performance metric

$$P(s) = \sum_n \text{Gaus}(s|n, \sigma) \times \text{Pois}(n|\lambda)$$



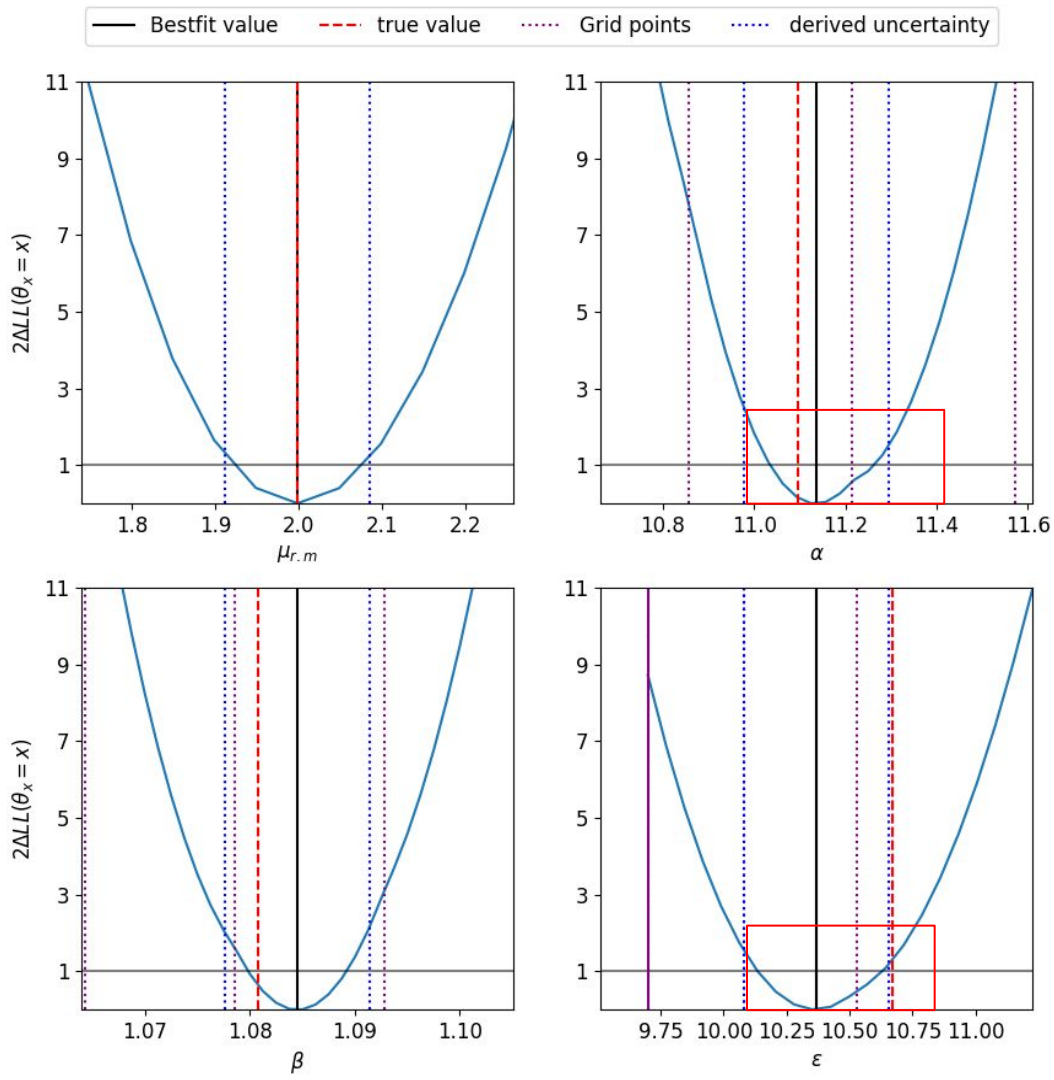
Parameters I'm talking about

Parameter	Description	Trace time
α	Linearly scale mean n_{prod}^q with energy	$11_{-0.5}^{+2.0} \text{keV}^{-\beta}$
β	Power law of mean n_{prod}^q with energy	1.1 ± 0.5
γ	Linear dependence of mean n_{prod}^{el} with density and electric field	$(4.8 \pm 0.2) \times 10^{-2}$
δ	power law of mean n_{prod}^{el} with electric field	$(4.8 \pm 0.2) \times 10^{-2}$
ϵ	Changes energy scale of mean n_{prod}^{el} energy dependence changes	$12.6_{-2.9}^{+3.4}$
ζ	Translates sigmoid of mean n_{prod}^{el} in energy	0.3 ± 0.1
η	Changes sigmoid shape of mean n_{prod}^{el} in energy	2 ± 1
θ	Translates sigmoid of mean n_{prod}^{ph} in energy	0.30 ± 0.05
l	Changes sigmoid shape of mean n_{prod}^{ph} in energy	2.0 ± 0.5



Some issues

Kinks in the likelihood between anchor points indicate that the differential rate term is showing correlation between parameters not captured in rate estimator.

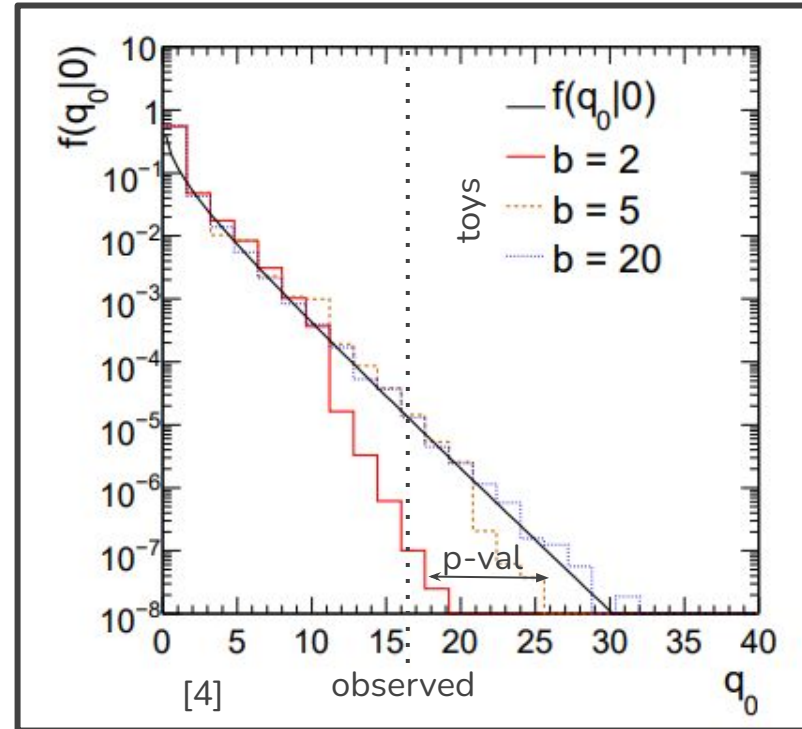


Why use *interpolated* rate estimators?

Markov Chain Monte Carlo could efficiently find the best fit with many parameters - simulate rate at every step.

Issues:

- Throw out all our diff programming benefits- too slow to evaluate rate estimator gradients+hessians
- Non-asymptotic** inference requires many many best fits $O(1000)$
- Asymptotic limit setting still requires $O(40)$.

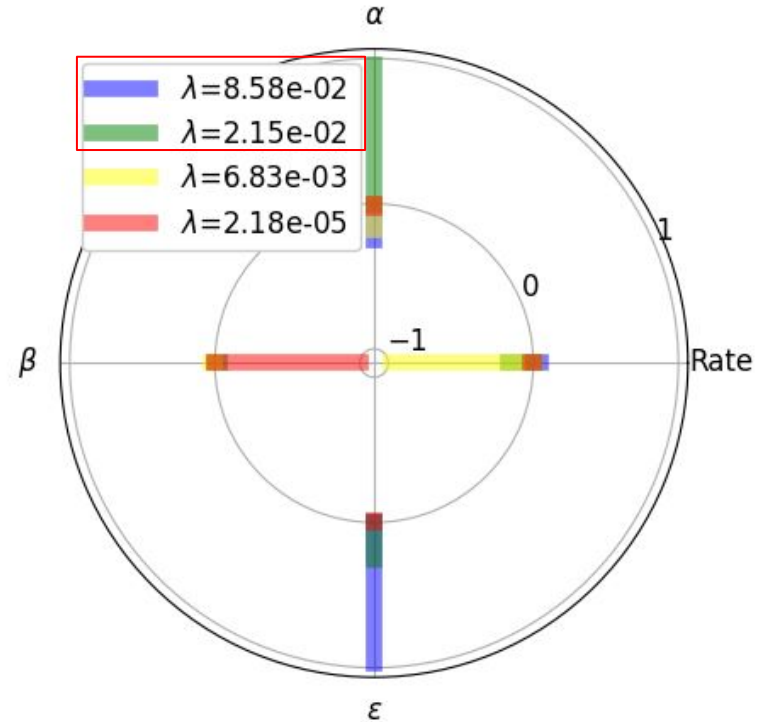




Fun Possible solution

1. Use MCMC or some other approximation to find the best fit to calibration data
2. Perform a **principal component analysis**
 - a. Largest eigenvalue eigenvectors of covariance matrix “**most information**”
 - b. Covariance \sim inverse hessian of likelihood
 - c. Tells us “**in which direction the likelihood/constraint is most flat**”
3. Use this to inform a reduced dimensionality

Here we would use epsilon-alpha and alpha-rate-beta.





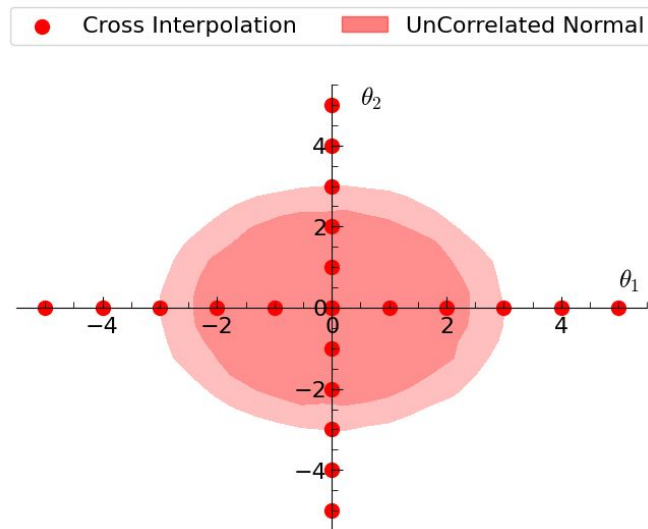
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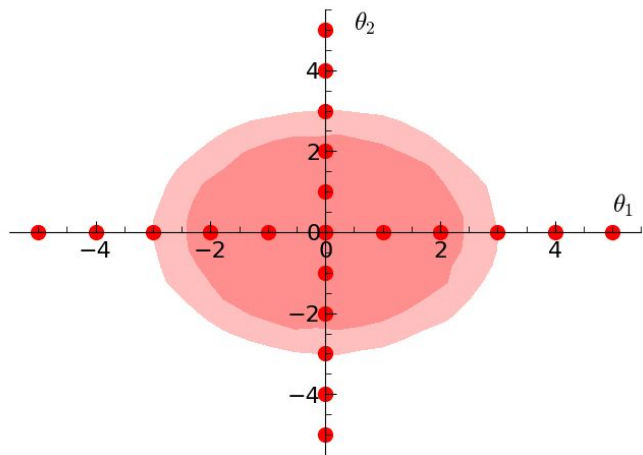
$$\mathcal{O}(n_{\text{parameters}} \times n_{\text{anchors}})$$

Rate estimator kerfuffle

Once you're correlated you need a grid to capture correlations and it gets out of hand

$$\mathcal{O}(n_{parameters} \times n_{anchors})$$

● Cross Interpolation ■ UnCorrelated Normal



$$\mathcal{O}(n_{anchors}^{n_{parameters}})$$

○ Grid Interpolation ■ Correlated Normal

