Thermalization in quantum fields

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Outline

Motivation 2PI effective action formalism Lorentz structure of the fermionic two-point functions Discretization Numerical results

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Applications of nonequilibrium field theory

- 1. Dynamics of early universe fields
 - Reheating dynamics (mainly classical, Hartree)
 - Formation of topological structures (classical)
 - Development of the radiation dominated universe
- 2. Heavy ion collisions



- Thermalization time scale
- Rethermalization after critical behaviour
- 3. Statistical field theory
 - Equilibration of a many-body system
 - Emergence of the Bose-Einsten and Fermi-Dirac statistics
 - Renormalization ...

What is thermalization?

Thermal equilibrium:

$$\hat{\rho} = e^{-\beta \hat{H}} / \mathrm{Tr} e^{-\beta \hat{H}} \qquad \left\langle \hat{X} \right\rangle = \mathrm{Tr} \hat{X} \hat{\rho}$$

Is thermalization possible in closed nonlinear system?

- Ch. Wetterich: Equilibrium is a fixed point of the evolution
- $\rho \not\rightarrow e^{-\beta \hat{H}} / \text{Tr} e^{-\beta \hat{H}}$ Unitarity!
- $\langle \hat{H} \rangle$ =const. uniquely determines the equilibrium ensemble.

But: $\langle \hat{H}^2 \rangle$, $\langle \hat{H}^3 \rangle$, ... conserved *(initial conditions)*

- The quantum ensemble cannot converge to equilibrium!
- Still, the quantum average of some selected observables may converge to the equilibrium value:

 $\langle \Phi(x)\Phi(y)\rangle_{\text{noneq}} \longrightarrow \langle \Phi(x)\Phi(y)\rangle_{\text{thermal}}, \text{ as } x_0, y_0 \to \infty$

Toy model: Chiral quark model, symmetric phase

Quantum initial value problem

 $\langle T_c \Phi(y) \Phi(x) \rangle =?$ $\langle T_c \bar{\Psi}(y) \Psi(x) \rangle =?$

The higher connected n-point functions are dropped

A convenient assumption: The density operator at t = 0 is quadratic

- Switching on the interaction at t = 0
- Problems with coupling renormalization:

either the final equilibrium or the initial state is singular

• Considerably simplifies our equations

Task: – Set up explicite equations of motion for the scalar and fermionic two-point functions

- Solve them numerically in the two-time-variable plane (x_0, y_0)

Approximation schemes

Boltzmann equation

- 1. Small occupation numbers, delute gas
- 2. On-shell processes only

Classical field theory

- 1. High occupation numbers (cosmological applications)
- 2. Nonperturbative dynamics (particle production, domain formation)
- 3. No quantum effects, failure of the UV description
- 4. Classical equilibrium \neq quantum equilibrium

Naive strategy for solving the quantum dynamics

- 1. Equations for the n-point functions built of Heisenberg operators
- 2a Collisionless approximation:

Connected 4-point functions are neglected \rightarrow no scattering, no thermalization

2b Inclusion of the 4-point functions, truncation at the 6-point functions:

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\rightarrow solution blows up
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Way out: The truncation of the hierachy should be carried out on the level of the
two-particle irreducible (2PI) effective action.Propagator resummation

 \rightarrow stable dynamics

2PI effective action for the fermionic fields

Classical action:

$$S = \int d^4x \left(ar{\psi}_i(x) [i\partial - m_f] \psi_i(x) + V(ar{\psi},\psi)
ight)$$

Effective action:

$$\Gamma[D] = -i\operatorname{Tr} \ln D^{-1} - i\operatorname{Tr} D_0^{-1} D + \Gamma_2[D] + const$$
$$iD_{0,ij}^{-1}(x,y) = (i\partial - m_f)\delta_c^4(x-y)\delta_{ij}$$

Equation of motion:

$$\frac{\delta\Gamma[D]}{\delta D_{ij}}(x,y) = 0$$

Self energy and the propagators (Schwinger–Dyson):

$$D_{ij}^{-1}(x,y) = D_{0,ij}(x,y)^{-1} - \Sigma_{ij}(x,y;D)$$
 $\Sigma_{ij}(x,y;D) = -i rac{\delta\Gamma_2[D]}{\delta D_{ji}(y,x)}$

EOM for the propagator:

$$(i\partial_x - m_f)D_{ij}(x,y) - i\int_z \Sigma_{ik}(x,z;D)D_{kj}(z,y) = i\delta_c^4(x-y)\delta_{ij}$$

 Σ contains the infinite power series of the propagators \rightarrow truncation

Time explicite equations

We consider the real and imaginary part of the propagator:

$$D_{ij}(x,y) = F_{ij}(x,y) - \frac{i}{2}\rho_{ij}(x,y) \operatorname{sgn}_{c}(x^{0},y^{0})$$

For fermionic fields:

$$(i\partial - m - \Sigma_0) F(x, y) = \int_{y_0}^{x_0} dz \Sigma^{\rho}(x, z) F(z, y) - \int_{y_0}^{y_0} dz \Sigma^{F}(x, z) \rho(z, y)$$
$$(i\partial - m - \Sigma_0) \rho(x, y) = \int_{y_0}^{x_0} dz \Sigma^{\rho}(x, z) \rho(z, y)$$

For scalar fields:

$$\left(\partial_x^2 + m^2 + \Sigma_{0,i}(x)\right) F_{ij}(x,y) = \int_{x_0}^{y_0} dz \Sigma_{ik}^F(x,z) \rho_{kj}(z,y) - \int_{x_0}^{x_0} dz \Sigma_{ik}^\rho(x,z) F_{kj}(z,y) \\ \left(\partial_x^2 + m^2 + \Sigma_{0,i}(x)\right) \rho_{ij}(x,y) = \int_{x_0}^{y_0} dz \Sigma_{ik}^\rho(x,z) \rho_{kj}(z,y)$$

The collision terms appear in the forms of memory kernels. The equations are time-reversal symmetric *(Energy is conserved)*

Lorentz structure

$$ho=
ho_S+i\gamma_5
ho_P+\gamma_\mu
ho_V^\mu+\gamma_\mu\gamma_5
ho_A^\mu+rac{1}{2}\sigma_{\mu
u}
ho_T^{\mu
u}$$
 and similarly for

Symmetry requirements imposed on the initial conditions:

- Space reflection and rotation: $\rightarrow \rho_A^0 = 0, \rho_A = 0$
- CP symmetry

$$egin{aligned} &
ho_V^0(x^0,y^0;p) =
ho_V^0(y^0,x^0;p), &
ho_V(x^0,y^0;p) = -
ho_V(y^0,x^0;p), \ &F_V^0(x^0,y^0;p) = -F_V^0(y^0,x^0;p), &F_V(x^0,y^0;p) = F_V(y^0,x^0;p), \end{aligned}$$

 $\rightarrow \rho_V, F_V$ are real, ρ_V^0, F_V^0 are imaginary

Discretiation:

- First order space derivative can be transformed out
 - \rightarrow no spatial fermion doubling occurs
- The time-like lattice spacing is much less than the spatial one
 → no time-like doubling either
- The discretized equations for ρ_V, ρ⁰_V and F⁰_V, F_V are reminiscent of the standard Leap-frog prescription for the canonica co-ordinate and momentum.
 → stable numerics

The formulation of the equations by means of the two-point functions gives a way to avoid the fermion doubling problem.

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The order of approximation

Self energies form a coupling constant expansion:

$$\Sigma^{\rho}_{\phi}(x^{0}, y^{0}; \vec{p}) = -8g^{2}N_{f} \int \frac{\mathrm{d}^{3}q}{(2\pi)^{3}} \rho^{\mu}_{V}(x^{0}, y^{0}; \vec{q}) F_{V,\mu}(x^{0}, y^{0}; \vec{p} - \vec{q}) ,$$

$$\Sigma^{F}_{\phi}(x^{0}, y^{0}; \vec{p}) = -4g^{2}N_{f} \int \frac{\mathrm{d}^{3}q}{(2\pi)^{3}} \left[F^{\mu}_{V}(x^{0}, y^{0}; \vec{q}) F_{V,\mu}(x^{0}, y^{0}; \vec{p} - \vec{q}) - \frac{1}{4}\rho^{\mu}_{V}(x^{0}, y^{0}; \vec{q}) \rho_{V,\mu}(x^{0}, y^{0}; \vec{p} - \vec{q}) \right] ,$$

$$\begin{split} \Sigma_{V}^{\rho,\mu}(x^{0},y^{0};\vec{p}) &= -g^{2}N_{s}\int \frac{\mathrm{d}^{3}q}{(2\pi)^{3}} \left[F_{V}^{\mu}(x^{0},y^{0};\vec{q}) \,\rho_{\phi}(x^{0},y^{0};\vec{p}-\vec{q}) \right. \\ & \left. + \rho_{V}^{\mu}(x^{0},y^{0};\vec{q}) \,F_{\phi}(x^{0},y^{0};\vec{p}-\vec{q}) \right], \\ \Sigma_{V}^{F,\mu}(x^{0},y^{0};\vec{p}) &= -g^{2}N_{s}\int \frac{\mathrm{d}^{3}q}{(2\pi)^{3}} \left[F_{V}^{\mu}(x^{0},y^{0};\vec{q}) \,F_{\phi}(x^{0},y^{0};\vec{p}-\vec{q}) \right. \\ & \left. - \frac{1}{4}\rho_{V}^{\mu}(x^{0},y^{0};\vec{q}) \,\rho_{\phi}(x^{0},y^{0};\vec{p}-\vec{q}) \right]. \end{split}$$

Infinite ladder diagrams are summed

Loss of initial information

- Different initial fermion number distribution
- Equal (conserved) energy density (uniquely determines the final temperature)
- The equations obey the time-reflection symmetry



Evolution of the particle distribution

• Particle number distribution from the quasiparticle picture:

$$F_{V}(t, t'; p) = \left(\frac{1}{2} - n_{qp}^{f}(p)\right) \cos[p(t - t')];$$

$$F_{\phi}(t, t'; p)|_{t=t'=\text{now}} = \frac{1}{\epsilon_{0}}(p) \left[n_{qp}^{s}(p) + \frac{1}{2}\right],$$

$$\partial_{t}\partial_{t'} F_{\phi}(t, t'; p)|_{t=t'=\text{now}} = \epsilon_{0}(p) \left[n_{qp}^{s}(p) + \frac{1}{2}\right],$$



• The emerging distribution is thermal,

its β parameter defines the quasiparticle temperature

Final equilibrium

Exact equilibrium relation:



- 1. First numerical evindence for thermalization in 3 + 1 dimensions
- 2. First observation of the formation of the Fermi-Dirac statistics
- 3. Estimate for thermalization time (RHIC) $\tau^{thermalization} \lesssim 1 \text{fm}$
- 4. Fermionic preheating ...