

CMOS Monolithic Active Pixel Sensors for Ionising Radiation

P.P.Allport², G.Casse², A.Evans^{2,3*}, A.R.Faruqi⁴, B.Gallop^{1,3}, R.Henderson⁴, M.Prydderch³, R.Turchetta³, M.Tyndel³, J.Velthuis², G.Villani³, N.Waltham³

1 University of Birmingham

2 University of Liverpool

* Speaker

3 Rutherford Appleton Laboratory

4 MRC Laboratory of Molecular Biology

- Overview
 - Introduction
 - UK HEP test structures
 - Star Tracker, 525x525 Space Science Prototype
 - Summary

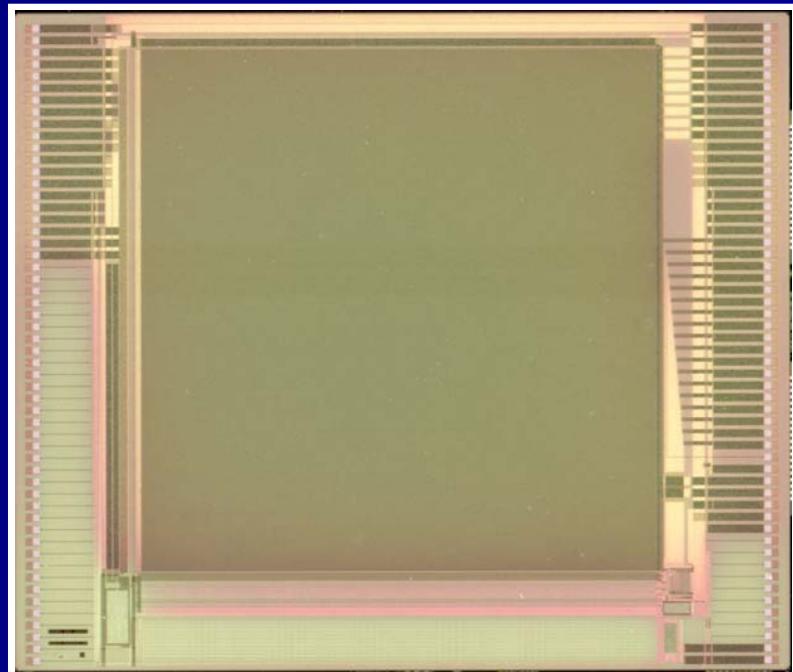


THE UNIVERSITY
of LIVERPOOL

Introduction

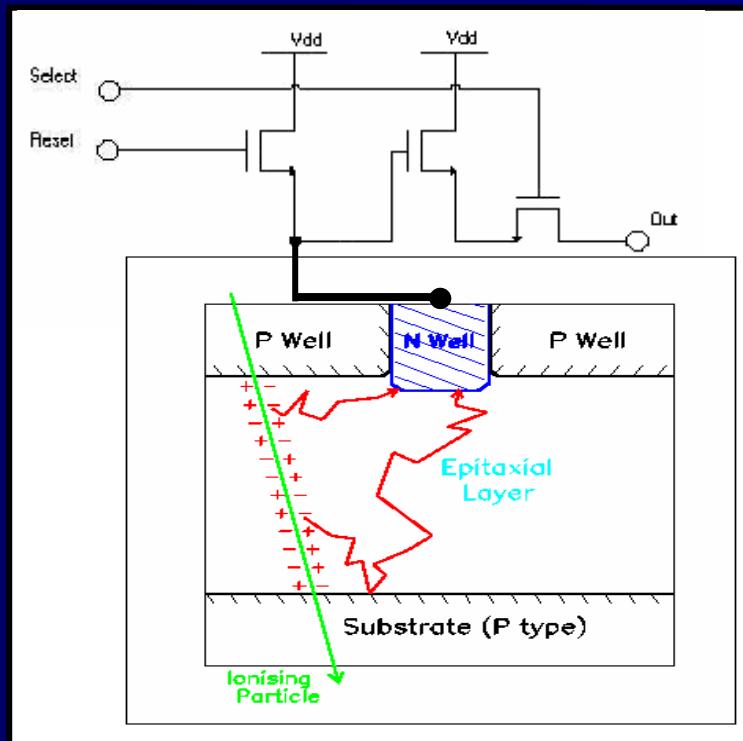
- Pixelated Detectors $\sim 3 - 100 \mu\text{m}$
- Standard VLSI CMOS technology
 - Low Power Consumption & Low Cost
- Industrial drive \rightarrow Visual Light Imaging
 - Digital Cameras e.g. Mobile-phone, Video Cameras...
- Scientific Application
 - High Energy Physics, Space Science...
- Advantages
 - Radiation hardness
 - Readout speed
 - On chip & in pixel intelligence
 - Room temperature operation

Dimensions: 2cm*2cm

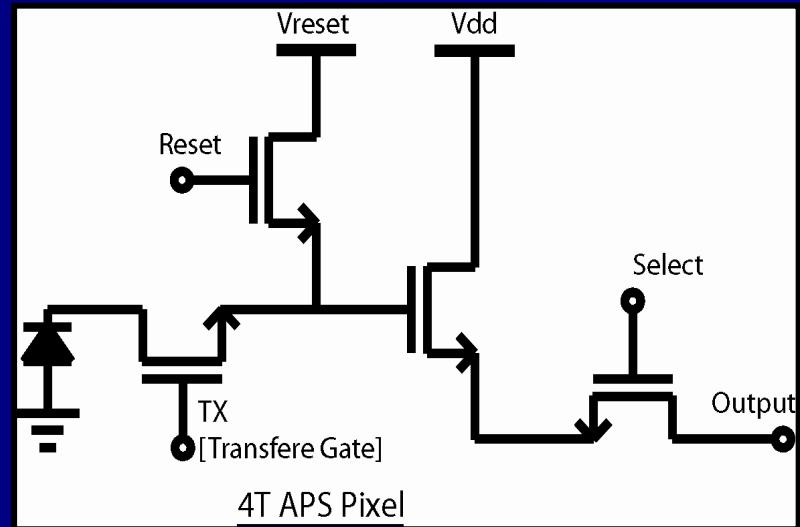


Star Tracker (525x525 $\rightarrow 25\mu\text{m}^2$ pitch)

Operation Principle



Basic 3Transistor (3T) pixel



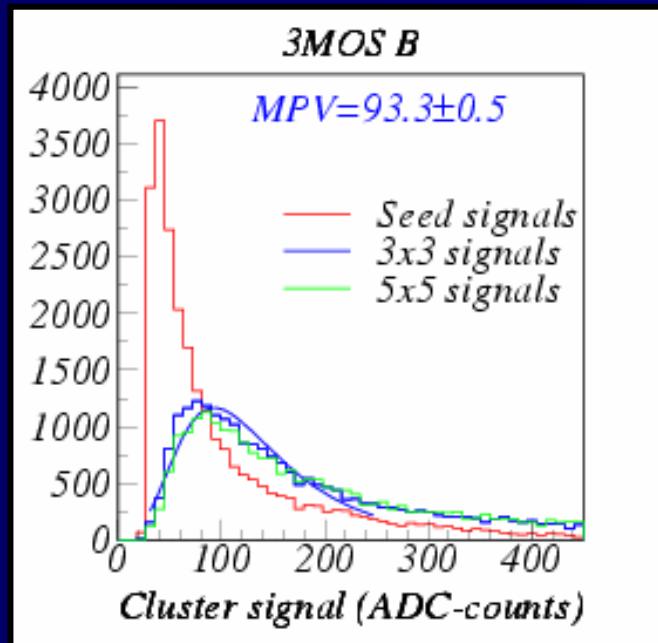
Basic 4Transistor (4T) pixel

UK HEP Test Structures

- APS1 (2001), IBM 0.25μm CMOS (2μm epi' layer)
 - Standard 3T & 4T pixels
 - Parametric Analysis
 - Radiation Hard, - 10^{12} protons/cm² [x10 requirement Linear Collider]
- APS2 (2003), TSMC 0.25μm CMOS (CIS) (8μm epi' layer)
 - Standard 3T & 4T pixels
 - FAPS, 10 memory cells, - burst readout 100ns
- APS3 (2004), TSMC 0.25μm CMOS (MS)
 - New pixel design – Deep N-well Diode

APS2

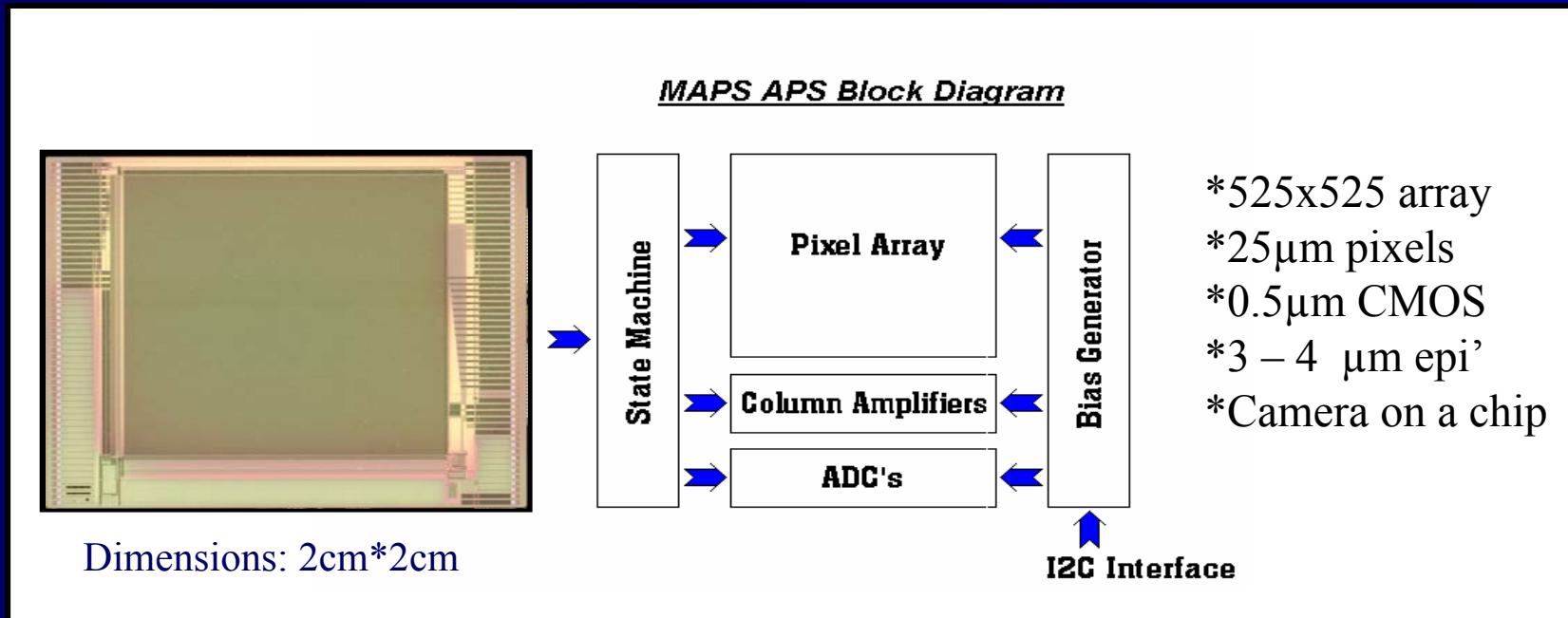
Source: Ru106 (Ruthenium)



Cluster Algorithm - Seed Signal $\geq 6\sigma$
- Area Signal $\geq 2\sigma$

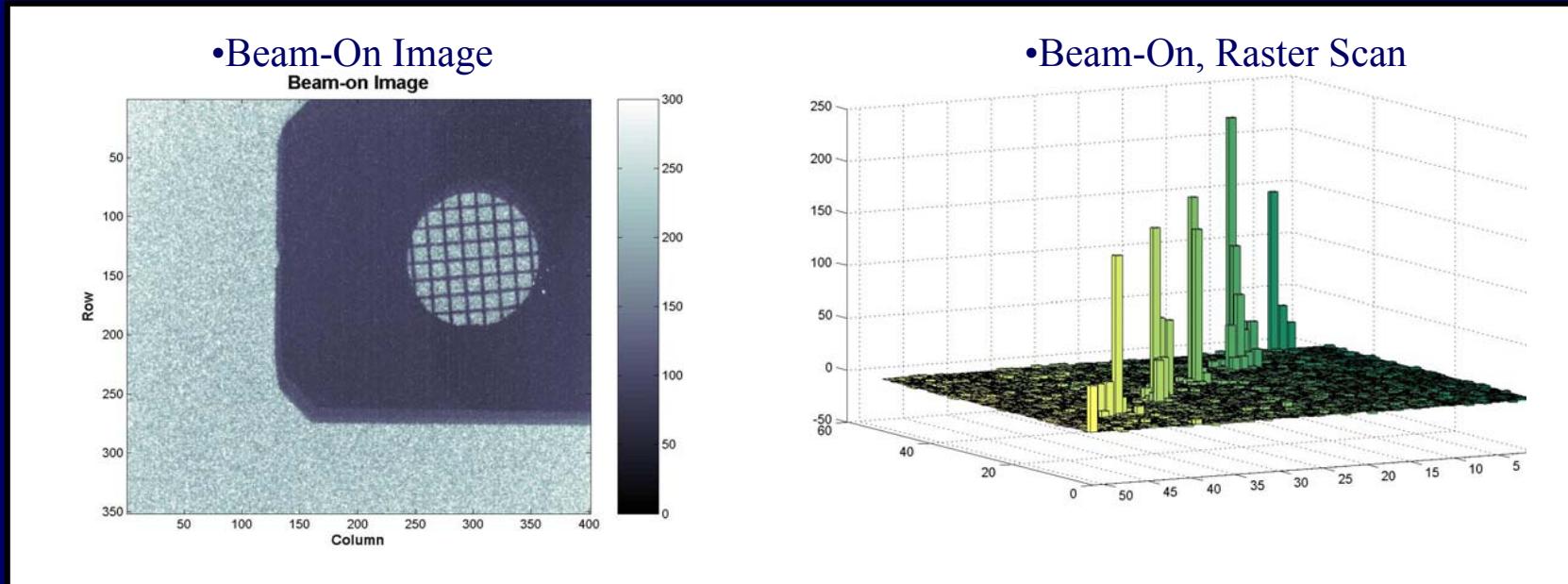
Type	Signal (3x3)	Noise (3x3)	S/N
3T B	93.3 +/- 0.5	3.828 +/- 0.006	24.4 +/- 0.1
3T A	66.1 +/- 0.8	3.298 +/- 0.007	20.0 +/- 0.2
4T C	97.8 +/- 0.9	4.14 +/- 0.01	23.6 +/- 0.2
4T B	107.7 +/- 0.9	4.17 +/- 0.01	22.8 +/- 0.2
4T A	105 +/- 1	4.5 +/- 0.01	23.3 +/- 0.2

Star Tracker [Space Science Prototype]



- Visual, Integrating Sphere ($\lambda = 510\text{nm}$)
- UV, Front etched and back-thinned
- Low Energy Electrons, 10keV \rightarrow 120keV

120 keV Electron Microscope [LMB Cambridge]

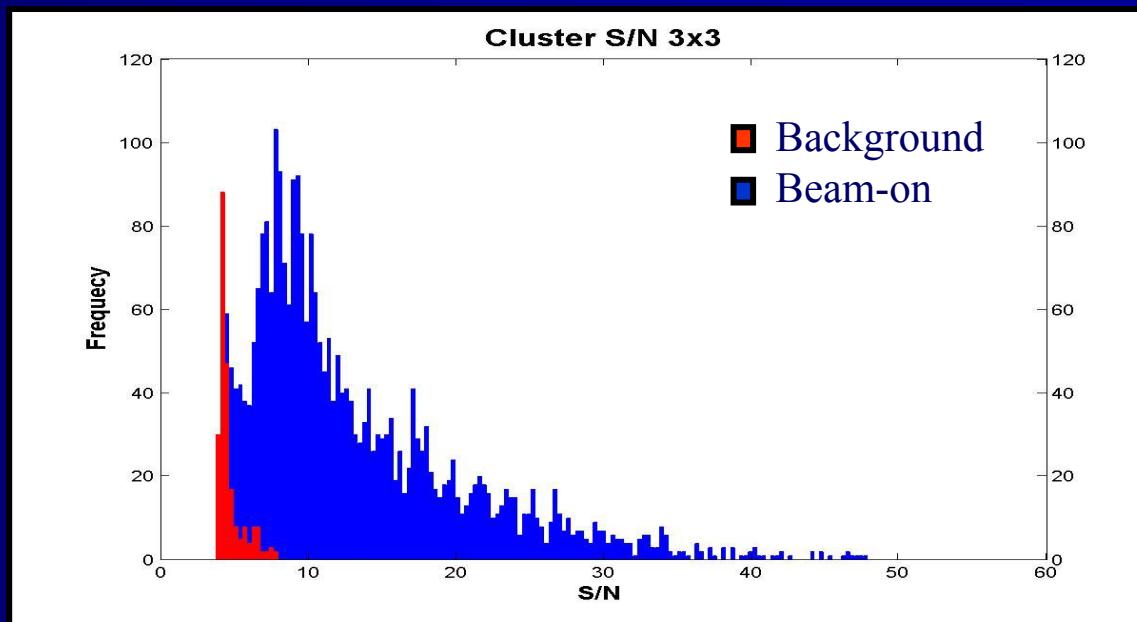


- Calibration
 - PTC, # of ADC units per e/h pair = 0.010 ± 0.001
 - Noise floor = 1.2 ± 0.1 ADC units
 - ETC, # of ADC's per incident 120keV electron = 23 ± 2
 - S/N ~ 20

Single Electron Sensitivity

[120 keV E.M.]

- Low electron flat field illumination of ~ 6 electrons per 10 pixels
- Increased noise
 - Noise floor = 3.2 ± 0.3 ADC units
 - Expected Signal Peak = 23 ± 2 ADC units
 - Expected S/N Ratio ~ 7
- Cluster Algorithm
 - [Seed $> 4\sigma$, neighbours $> 2\sigma$]
- S/N Ratio = 8 ± 2



Summary

- MAPS are a new and promising technology
- Detector for HEP experiments
 - MIP S/N = 24.4 ± 0.1
- Star Tracker, demonstrated direct detection of 120 keV electrons
 - 120 keV S/N = 8 ± 2
- Results expected soon for
 - Flexible Active Pixel Sensors (10 memory cells per pixel)
 - Deep N_Well

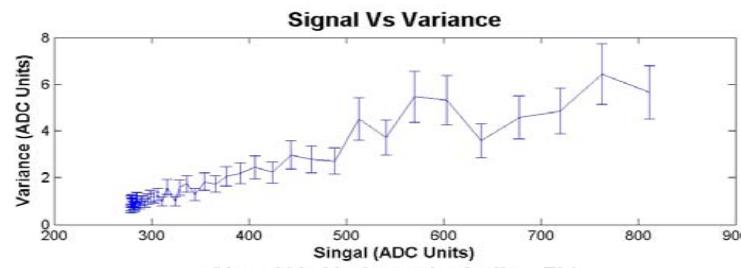
MAPS

• Star_Tracker, Characterisation:

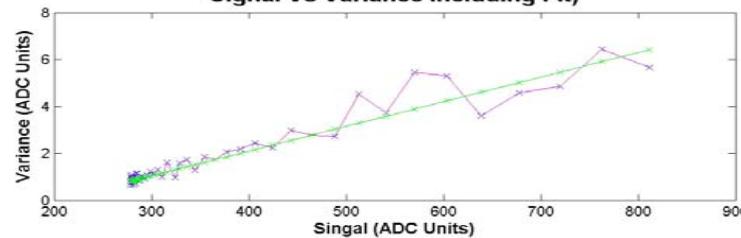
- Uniform Light Source → Integrating sphere → PTC
510nm

PTC

Variance (ADC Units)



Signal (ADC Units)

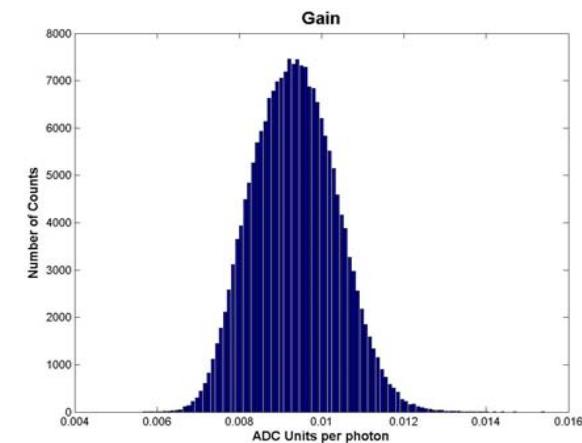


→ Full Well Capacity = 400k elec'

→ Signal Uniformity

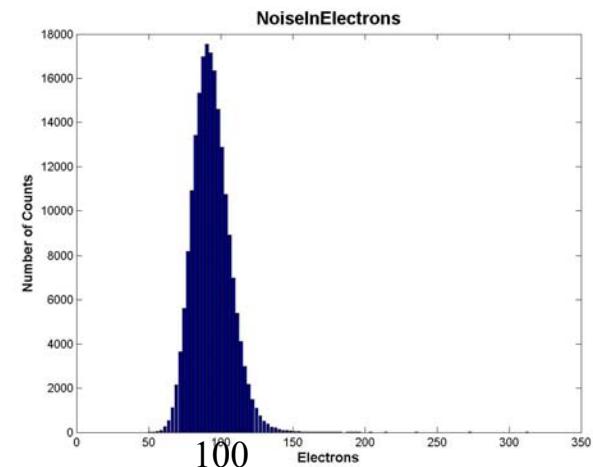
International Workshop on
Radiation Imaging Detectors
Arwel Evans

Gain Distribution Photons



Gain = 0.01 ± 0.0009 ADC/photon

Noise Distribution Electrons



Electrons

MAPS

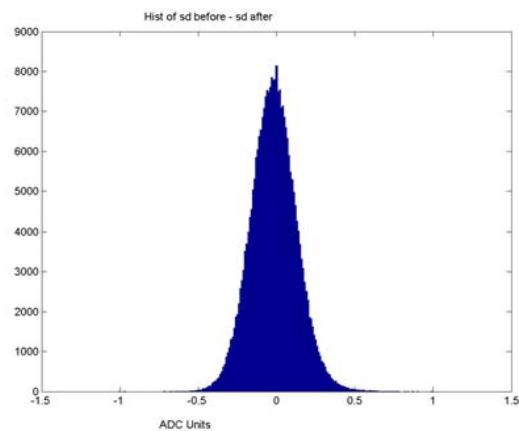
• Radiation Damage

- Dose: ~ 18krads
- Pedestal shift of ~ 21 ADC units
- Noise unchanged

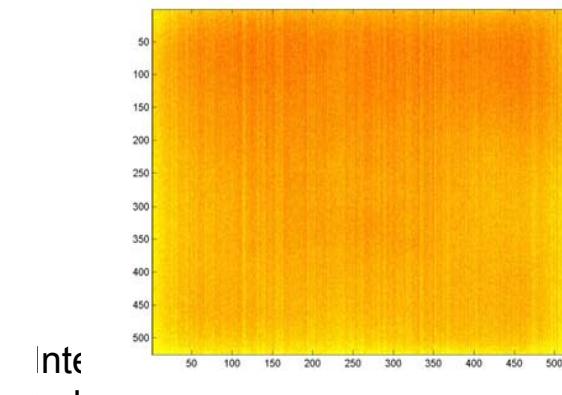
Main sources of pedestal change

- Increase in leakage current NO
- Transistor threshold voltage shift (V_{th}) YES, but can be corrected by DS or *design in deep submicron

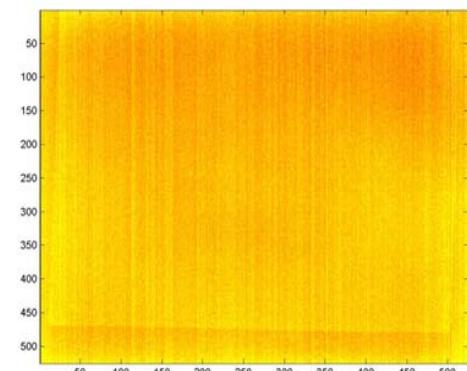
*Noise Histogram: Difference after-before



*Background Before Experiment

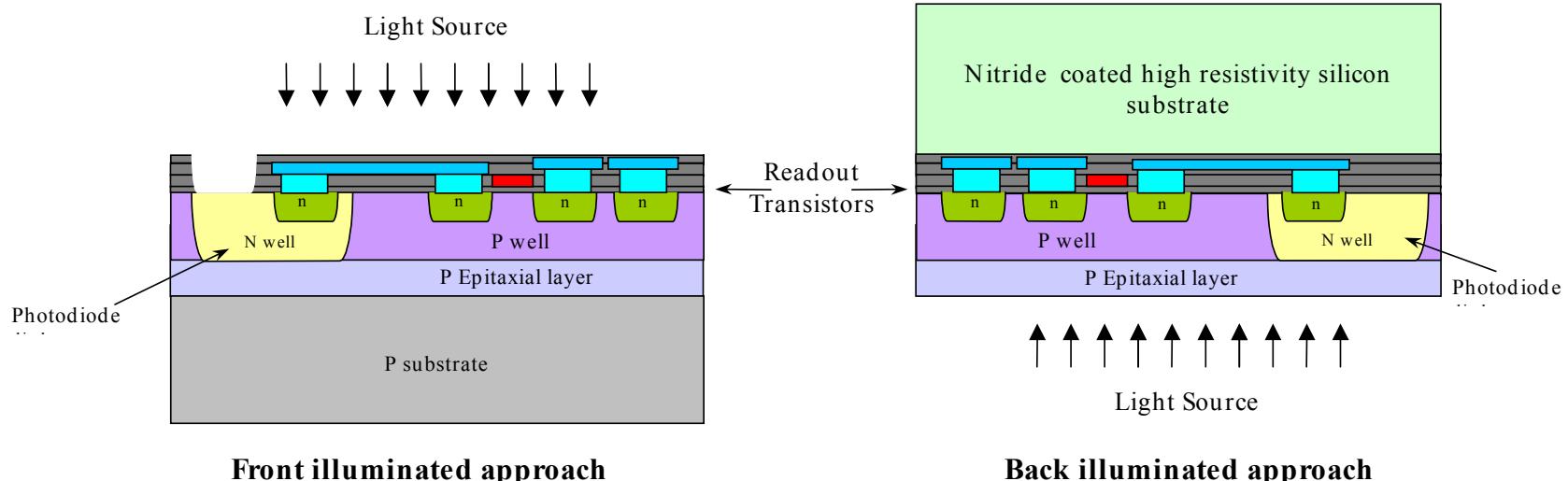


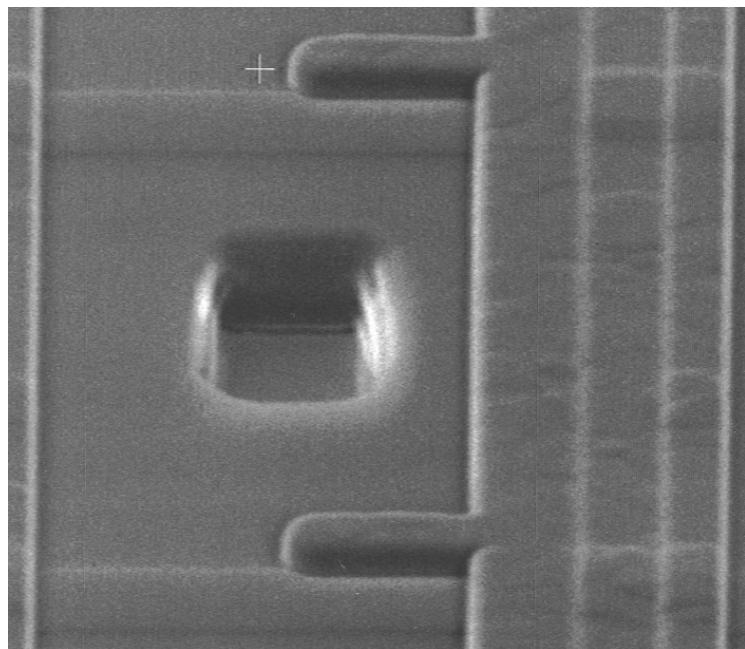
*Background After Experiment



MAPS

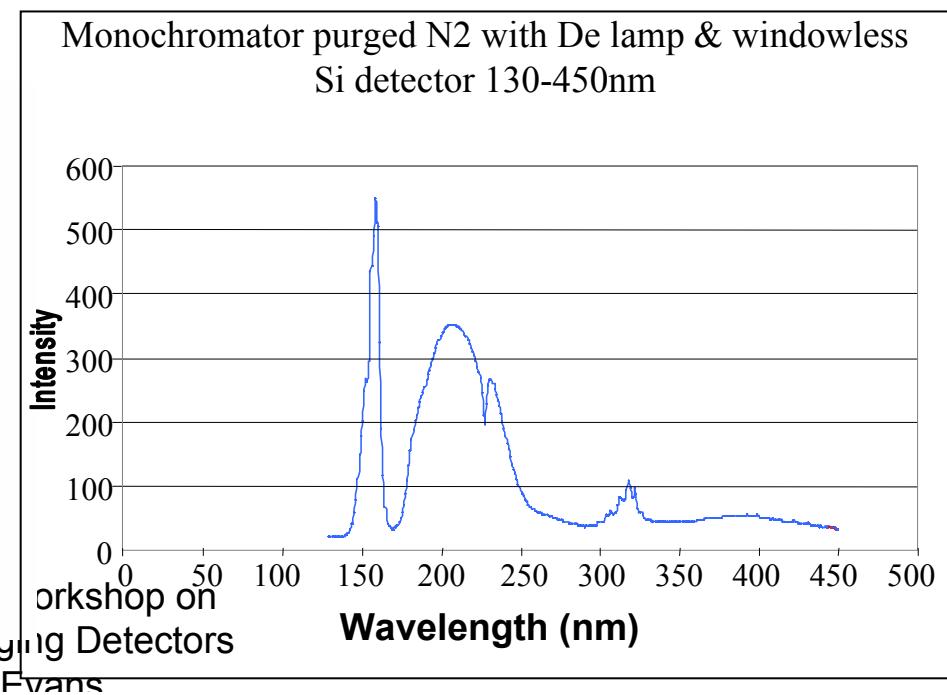
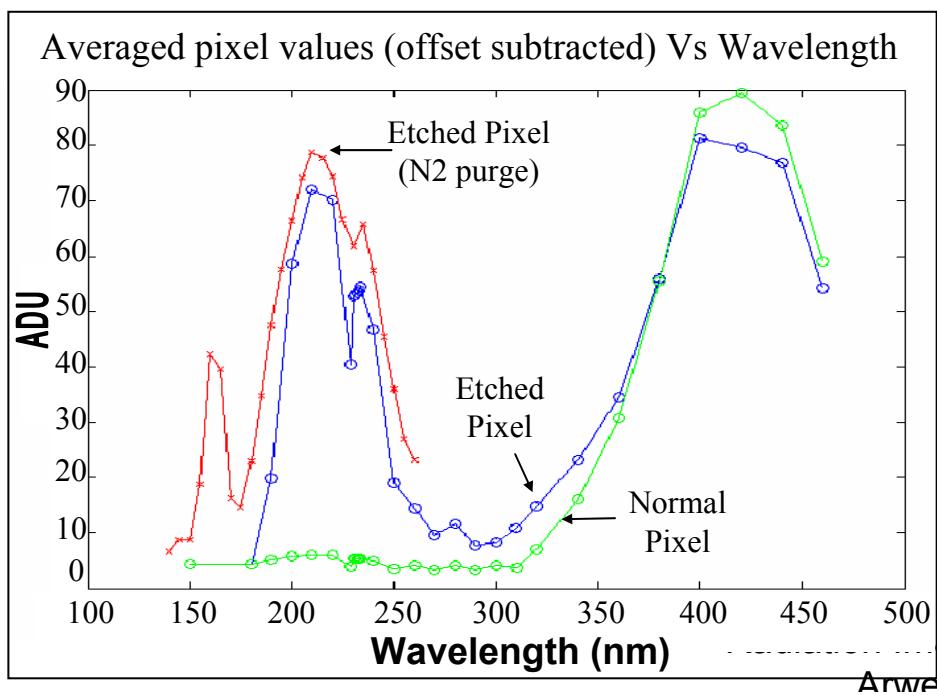
- UV program





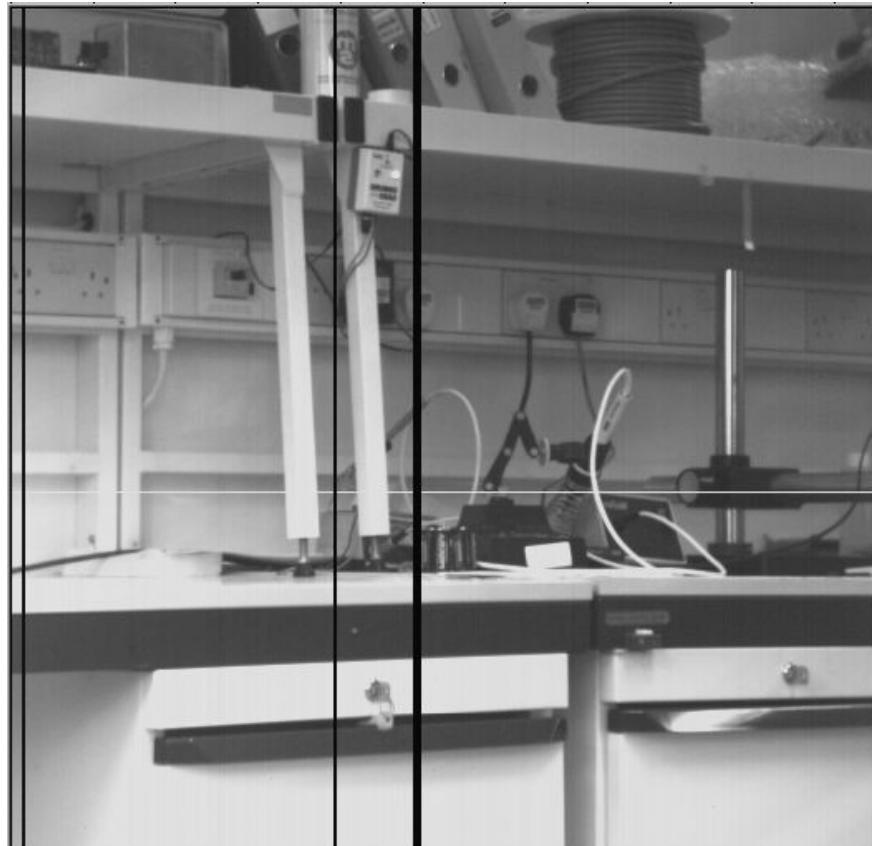
- Front Etched UV Response

*Focused Ion Beam



MAPS

- Back Thinning
- Image from two back-thinned Star_Tracker sensors



Radiation Imaging Detectors
Arwel Evans

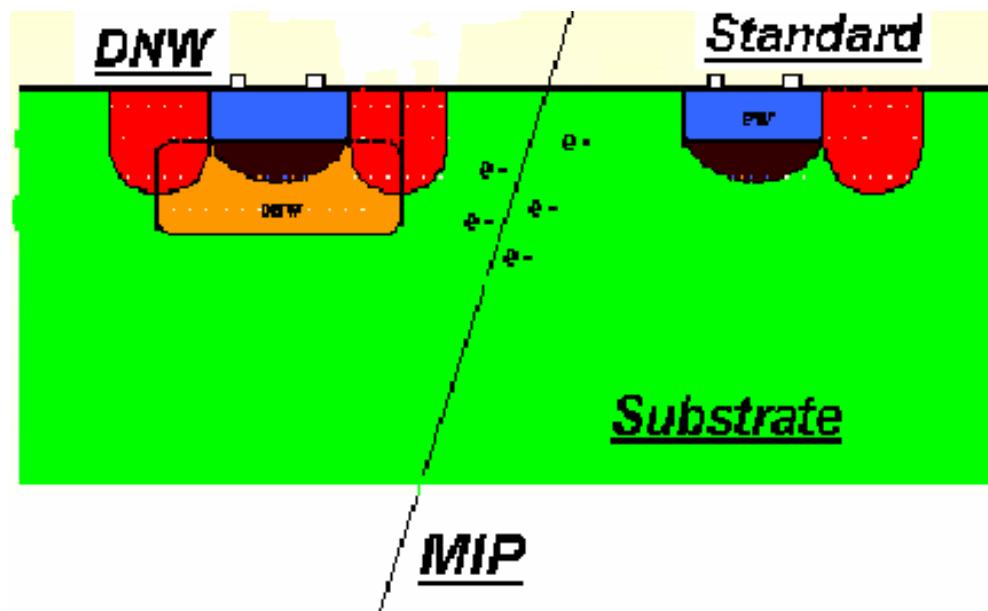
MAPS

•RalHepAPS 3

Deep N-Well Diode

→ Increased E-field - Charge Collection efficiency, Amount of Charge Collected & Charge Collection Speed

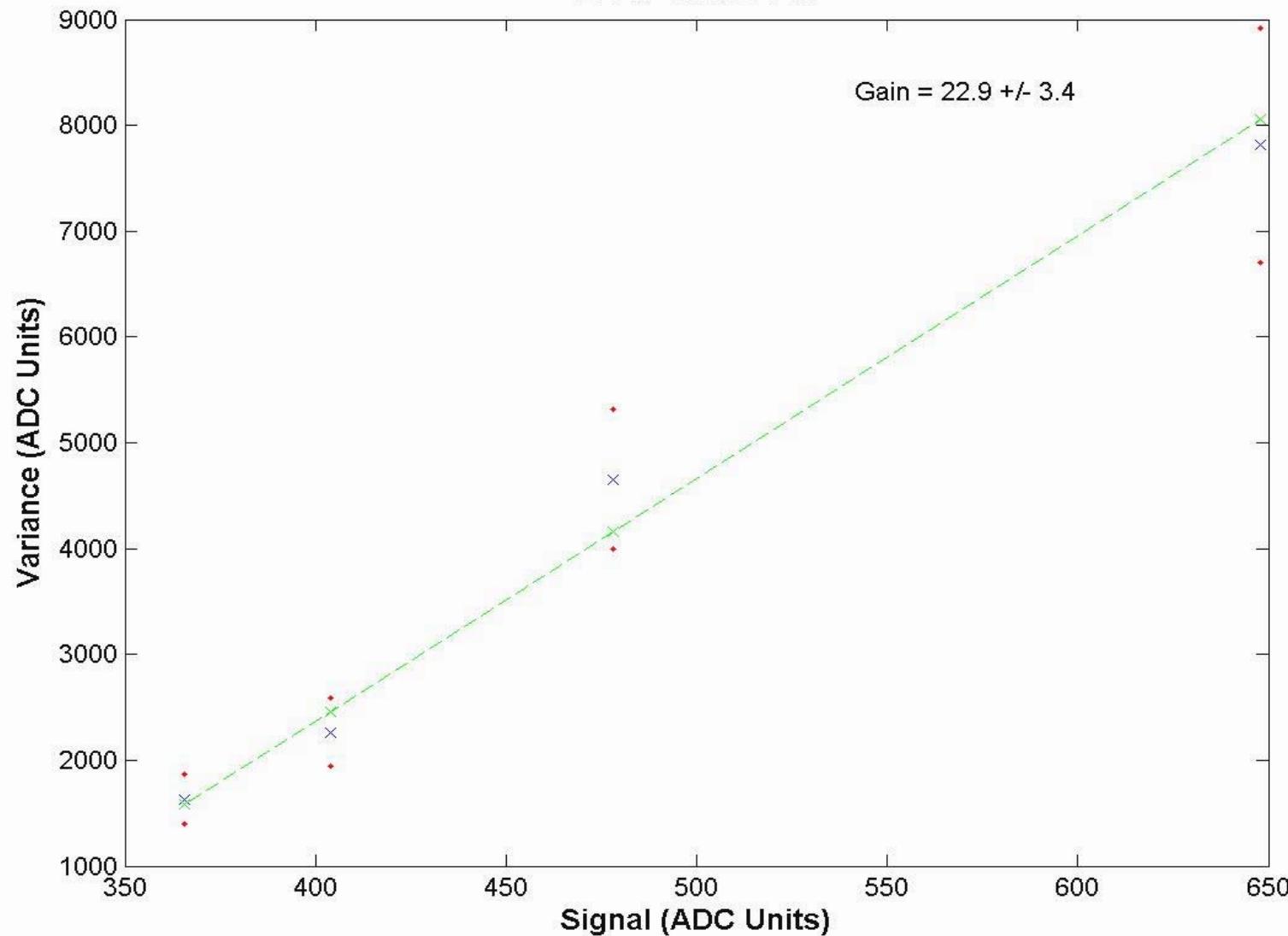
- Increased Radiation Hardness
- Negative Substrate Contact



*Designed by: Arwel Evans, Bruce Gallop,
Renato Turchetta

International Workshop on
Radiation Imaging Detectors
Arwel Evans

ETC With Fit



PTC

