

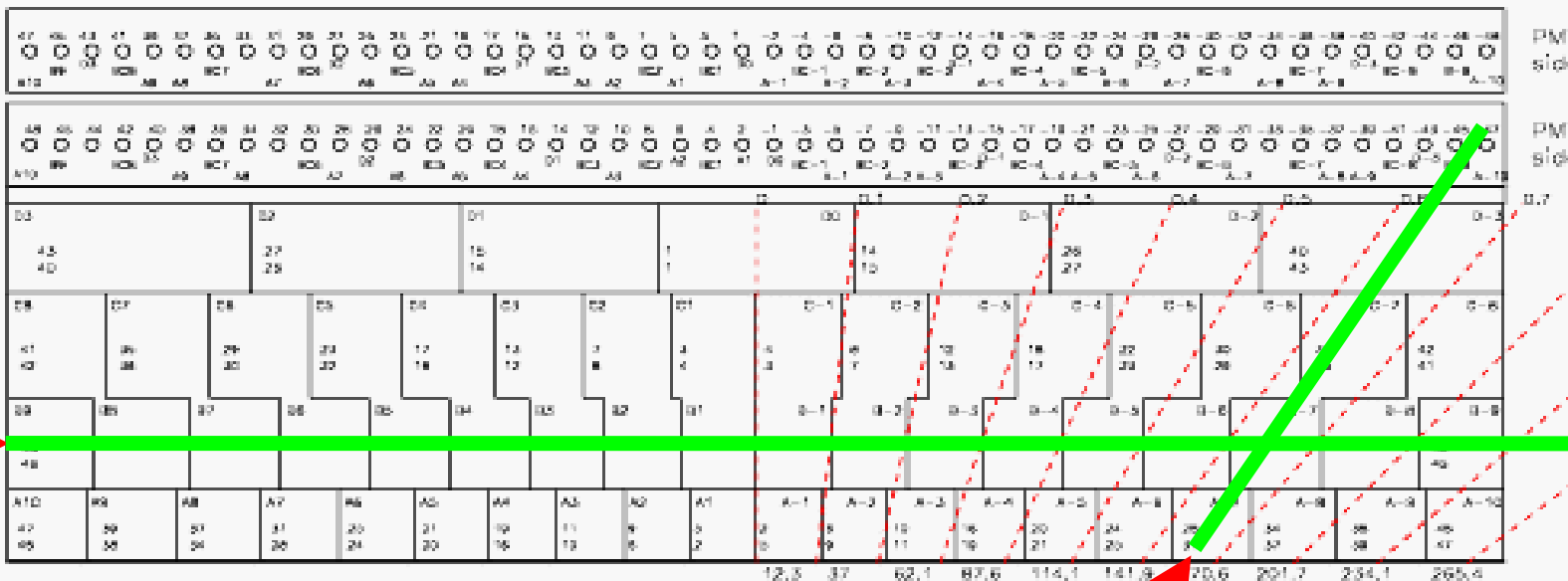
ATLAS Simulation with FADS/Goofy: electrons in TileCal

Andrea Dotti Anna Lupi
University and INFN Pisa

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Data sets

0 deg runs



$\eta = -0.65$
runs

$\theta = 90^\circ$

$\eta = -0.65$

energy (GeV)	run number
10	210720
20	210563
50	210487
100	210551
180	210054

energy (GeV)	run number
10	210731
20	210612
50	210479
100	210311
180	210121

- TestBeam data from July 2002
- Two different impinging directions have been analyzed for 5 different energies

MC Simulations

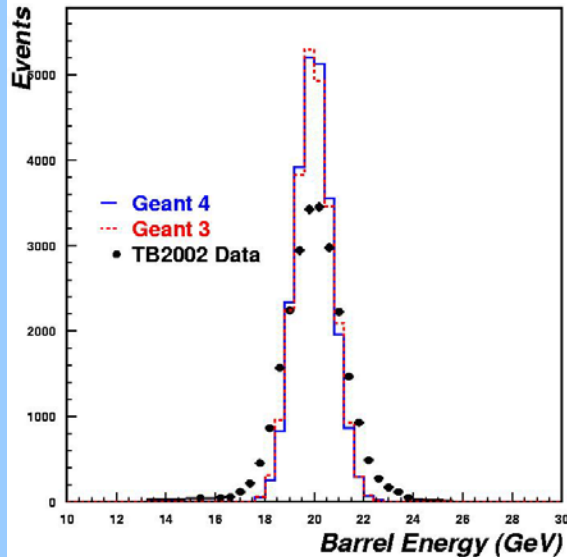
- A detailed study of signal shape showed that the contribution of electronic noise and photostatistic fluctuations are important (in particular at low energies) for a correct simulation of the calorimeter response
- These two effects have been added in the analysis steps
- TB data have been compared with G4 TileCal simulation (FADS/Goofy, V. Tsulaya) using the two physics lists QGSP 2.7 LHEP 3.6

Calibration and noise smearing

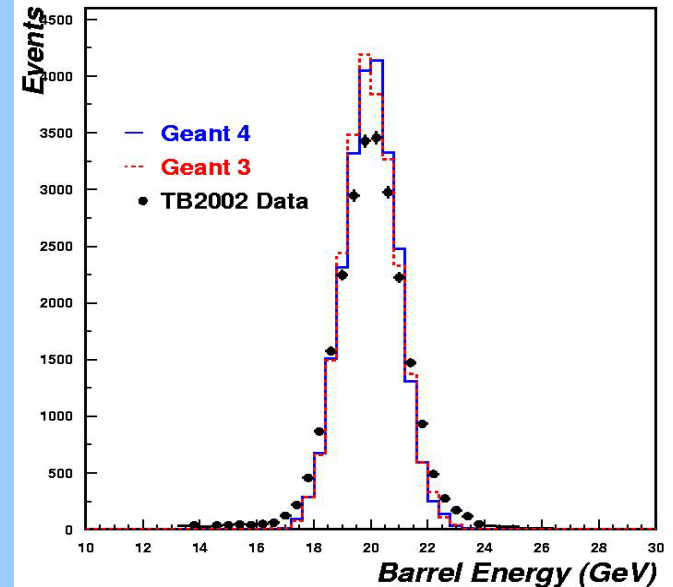
- Both Data and MC have been calibrated to the beam energy: **average visible energy normalized to beam energy**
- **Electronic noise** has been measured from data and added to simulation as a gaussian smearing
- **Photostatistic fluctuations** have been added as an additional source of noise in simulations (current value 53pe/GeV). Approximated as another gaussian smearing added to simulated data

Noise and photostatistic contribution (1/3): electrons $\theta=90^\circ$ and $\eta=-0.65$

20 GeV electrons, $\theta=90$ (Data Run r0210595)

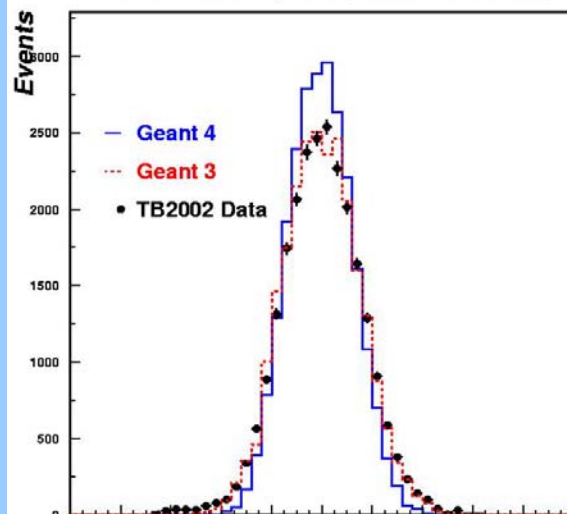


20 GeV electrons, $\theta=90$ (Data Run r0210595)

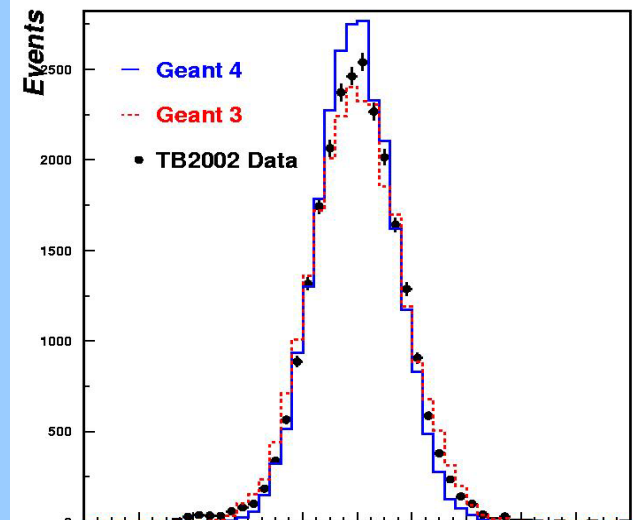


Only
electronic
noise
added:
agreement
improves in
both
geometries

20 GeV electrons, $\eta=0.65$ (Data Run r0210612)

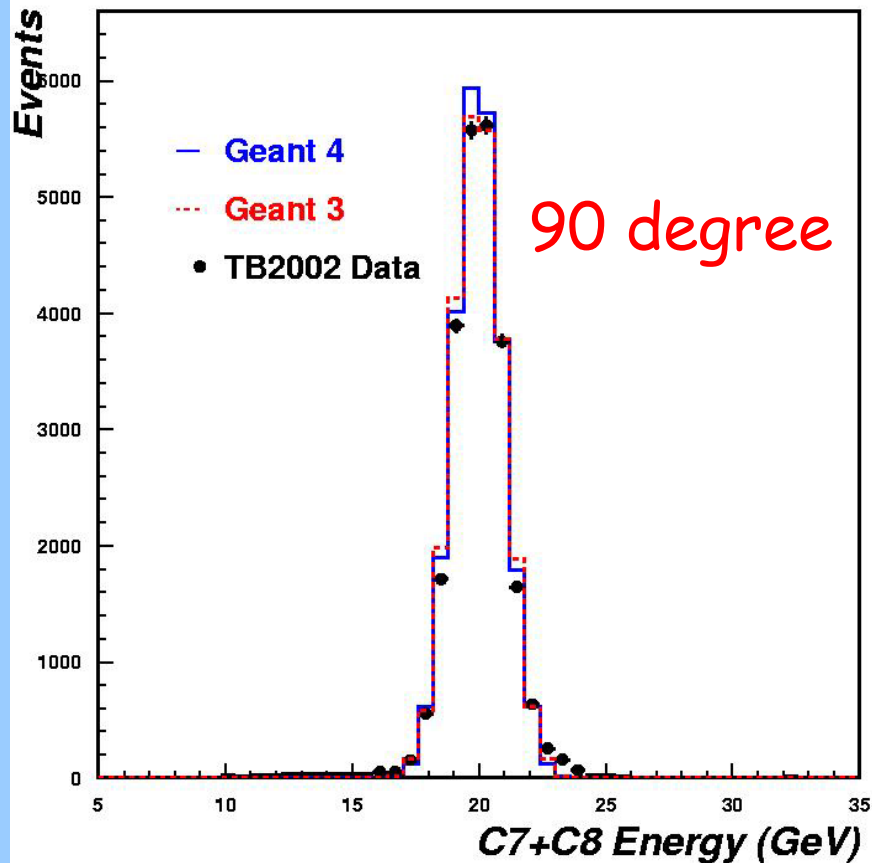


20 GeV electrons, $\eta=0.65$ (Data Run r0210612)

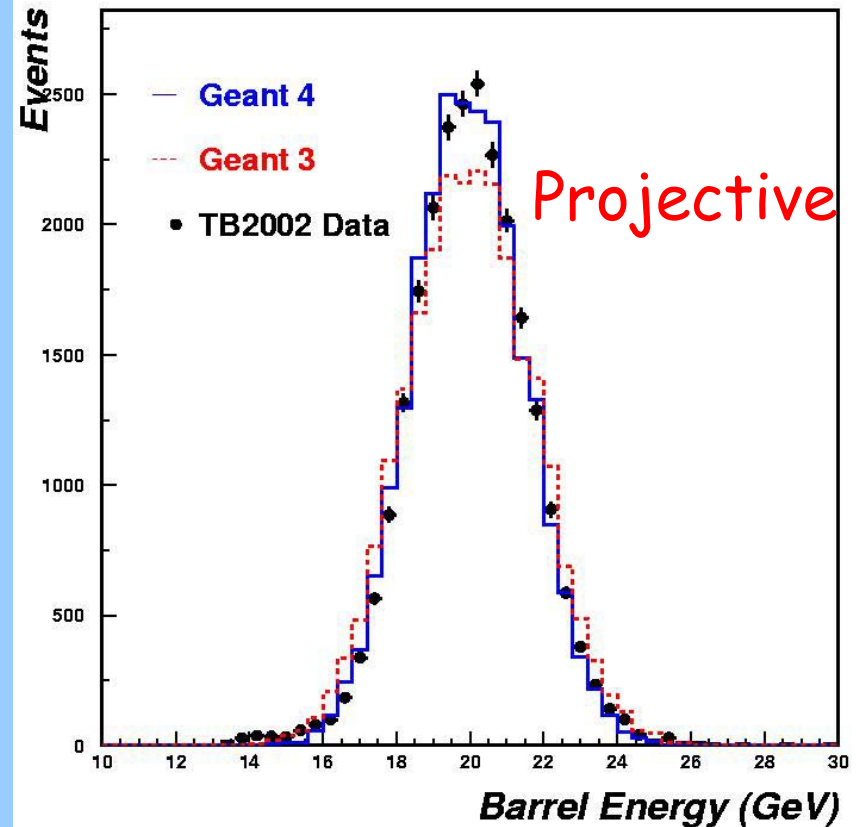


Noise and photostatistic contribution (2/3): electrons $\theta=90^\circ$ and $\eta=-0.65$

20 GeV electrons, $\theta=90$ (Data Run r0210595)



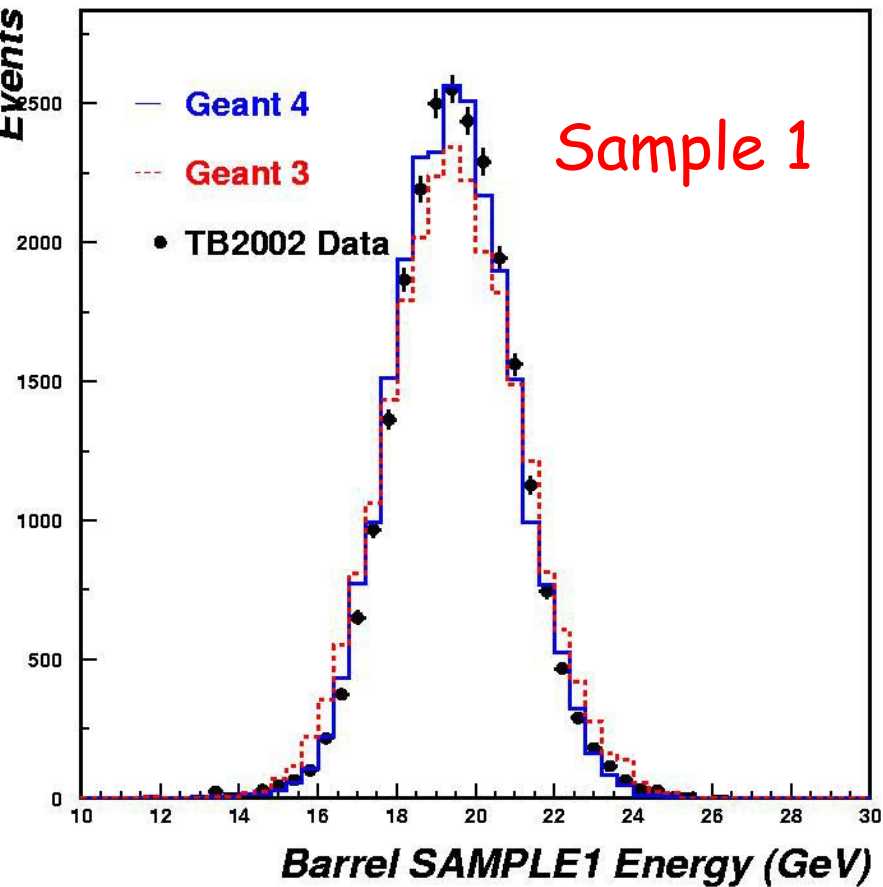
20 GeV electrons, $\eta=0.65$ (Data Run r0210612)



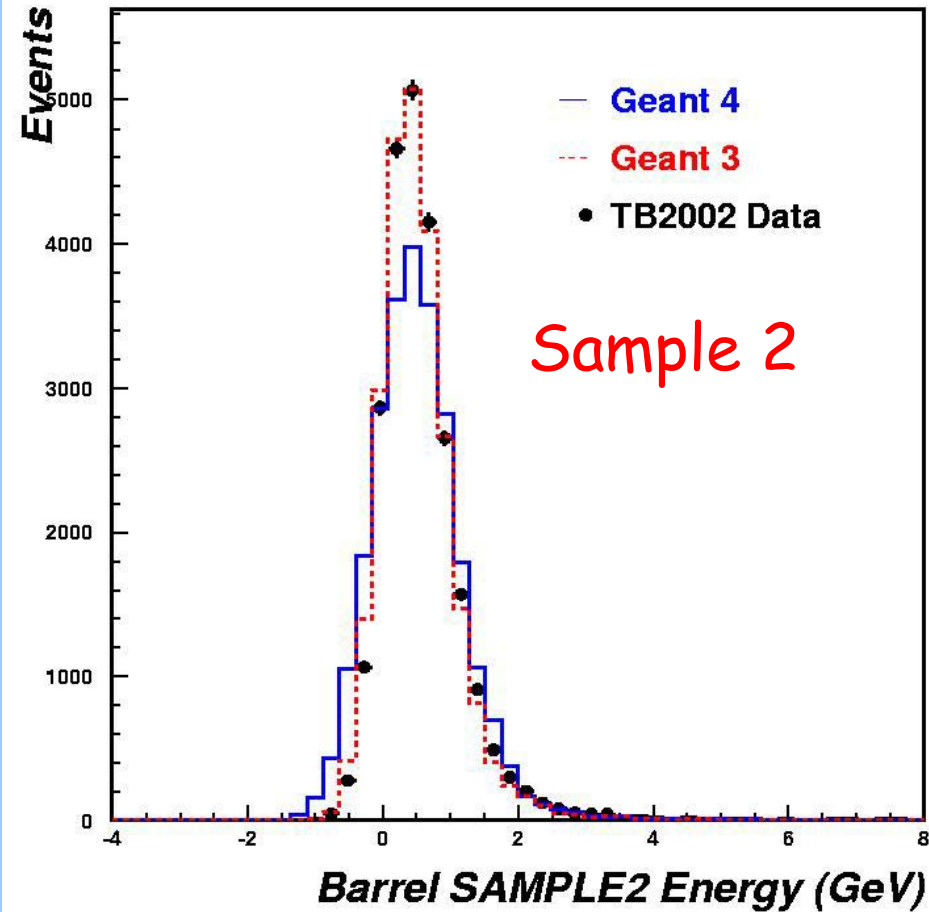
Including noise and photostatistic fluctuations the agreement is even better for both the geometries

Noise and photostatistic contribution (3/3): longitudinal shower profile

20 GeV electrons, $\eta=0.65$ (Data Run r0210612)



20 GeV electrons, $\eta=0.65$ (Data Run r0210612)

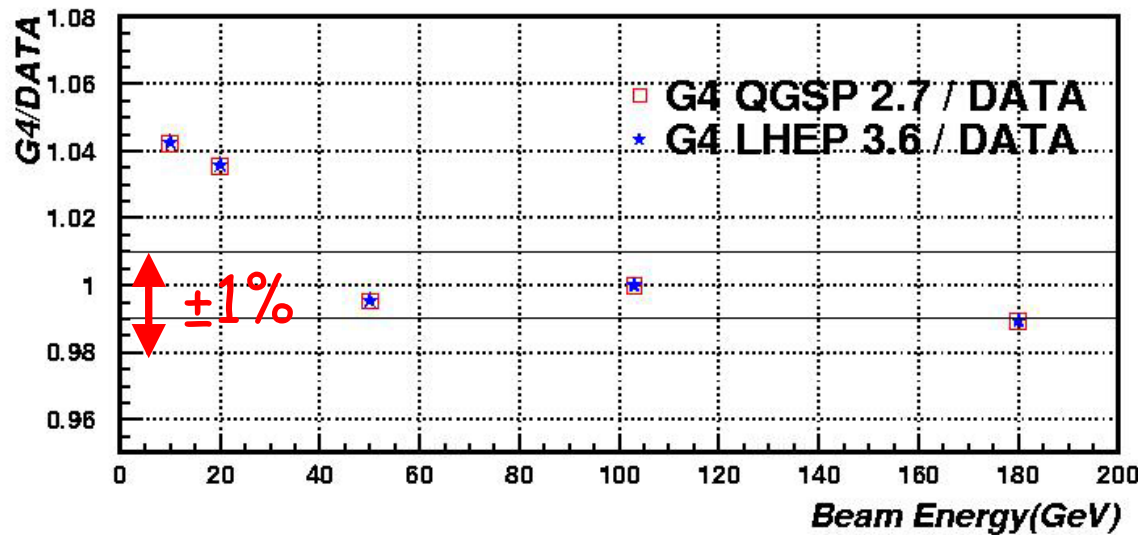
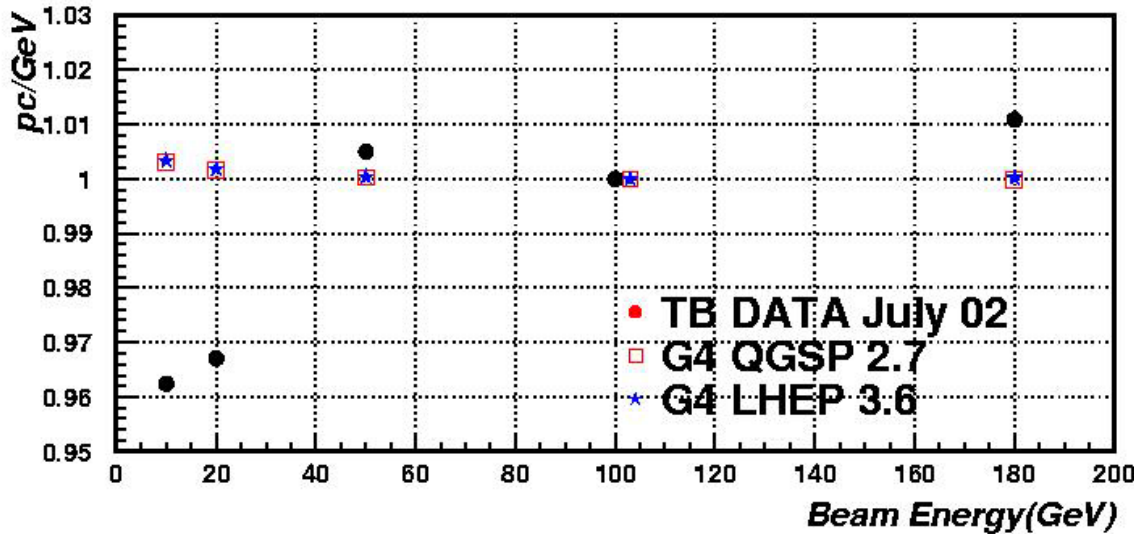


In G4 the shower tail (sample2) is less precisely simulated.

$\theta=90^\circ$ (tile row 5)

Calibration Constant (pC/GeV) vs E

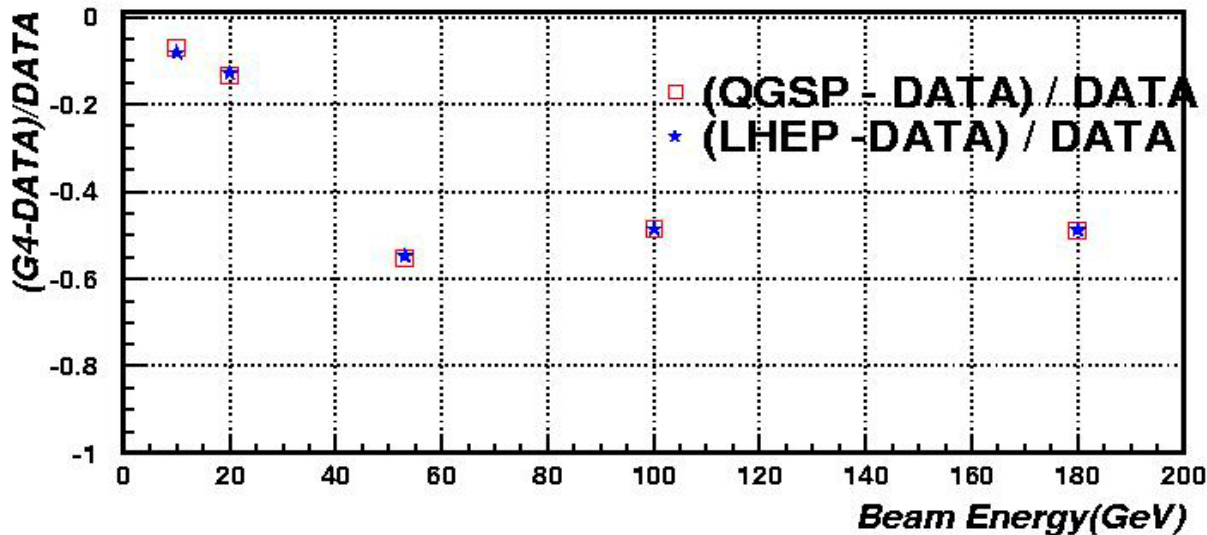
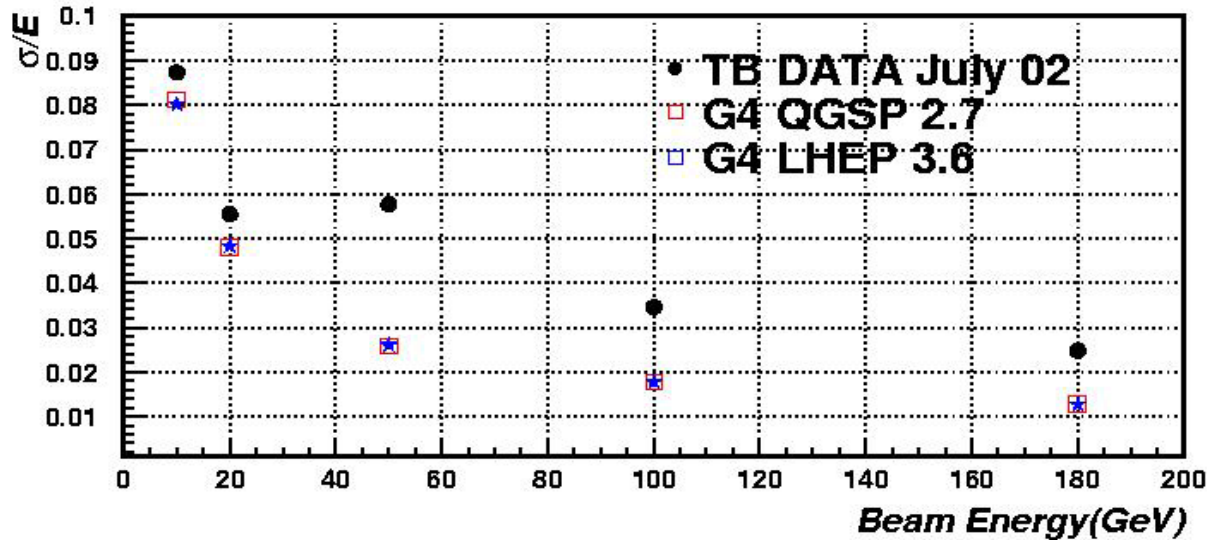
electrons with $\vartheta = -90$, tr5



- As expected no changes in em physics in the two lists.
- Data are normalized to the 100 GeV point (for real data $pC/GeV=1.2$)
- At low energy (10, 20 GeV) electrons have lower signals (not in MC), probably a problem related to: beam energy, electronic amplification... (under investigation)
- Excluding these points the agreement is $\pm 1\%$

σ/E vs E

electrons with $\vartheta = -90$, tr5

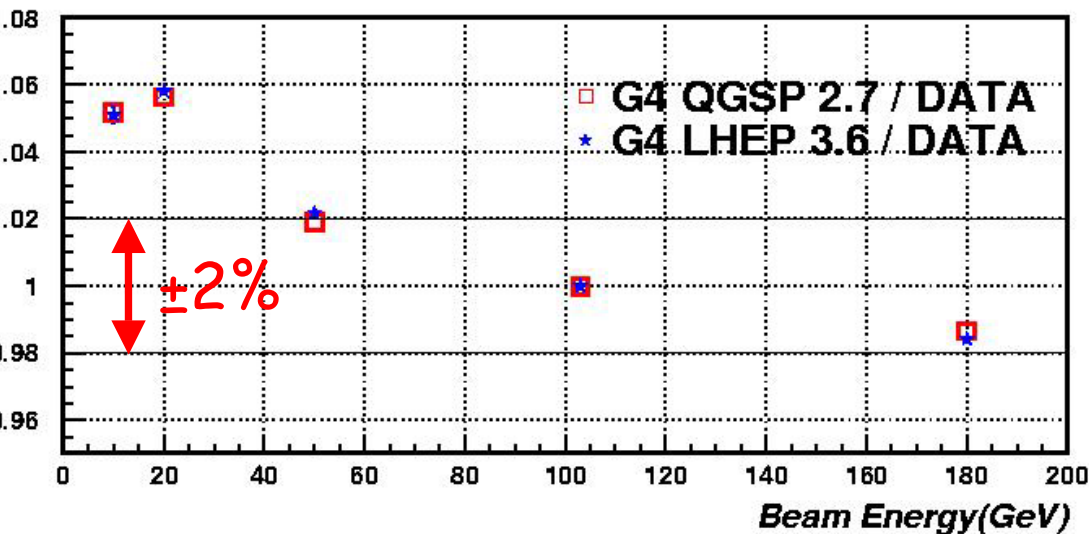
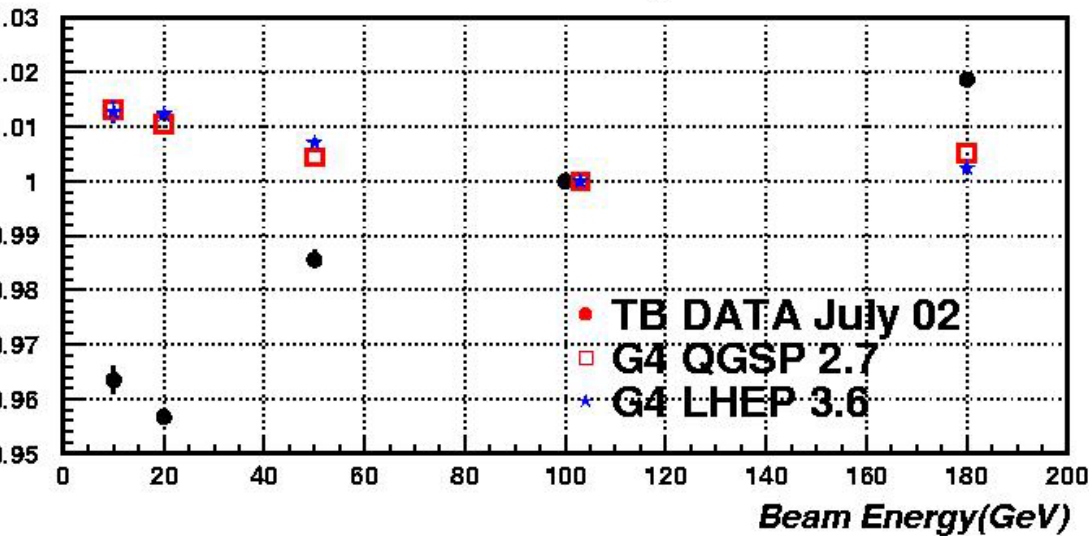


• The agreement in this geometry is not satisfactory

$$\eta = -0.65$$

pC/GeV vs E

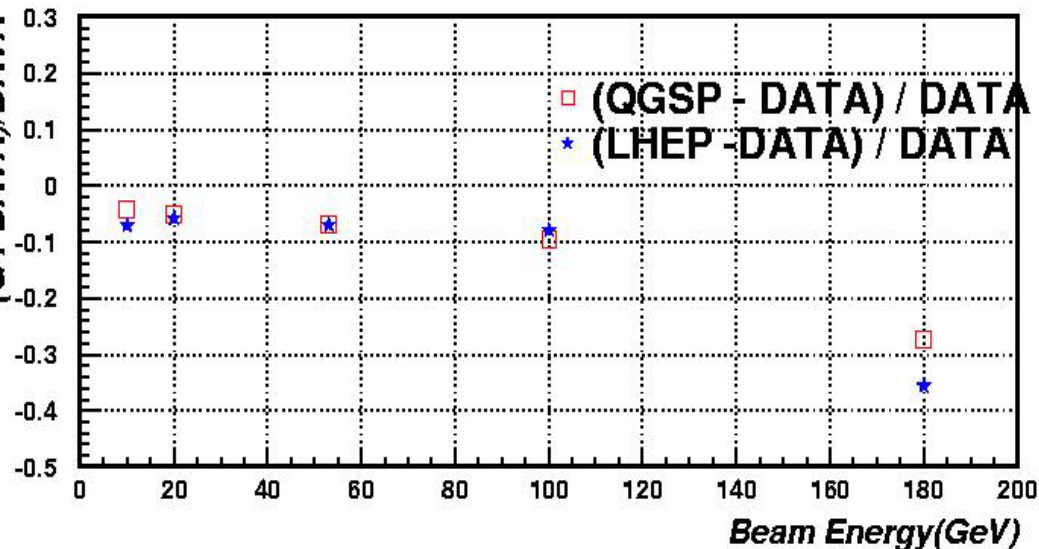
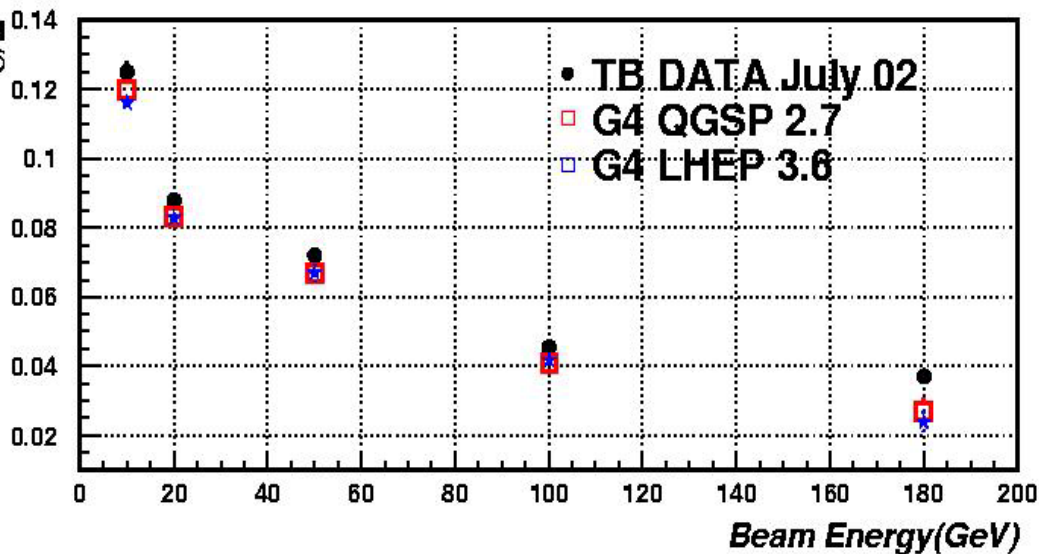
electrons with $\eta = -0.65$



- Here a dependence on E in data is visible ($\sim 6\%$) not visible in $G4$. The problem is under study
- Again at low energies with TB data response is lower than expected
- For $E > 20 \text{ GeV}$ the agreement is $\pm 2\%$

σ/E vs E

electrons with $\eta = -0.65$



• Very Good agreement between data and simulations

Fit results

a/sqrt(E)

b

QGSP	(34±1)%	(0.9±0.2)
LHEP	(34.5±0.8)%	(0.7±0.1)
DATA	(35±1)%	(1.1±0.1)

Conclusions

- The electronic noise and photostatistic is important to be included to correctly simulate the shapes
- Total energy deposit is correctly simulated in *G4*, energy deposit in single samples seem to show larger distributions in shower tails
- pC/GeV constant is well described for high energies ($>20GeV$)
- σ/E for electron is well simulated at all energies for projective data while at 90° data show a worst resolution (by a factor $\sim 50\%$)