



# Calibration of the MINOS Detectors Using Stopping Muons

**Jeff Hartnell**

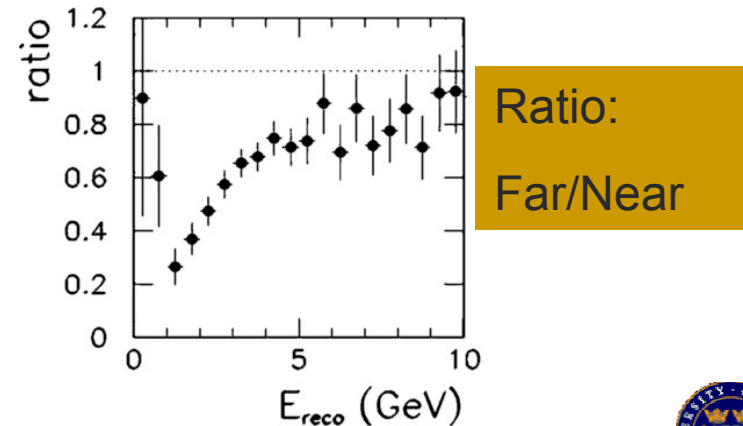
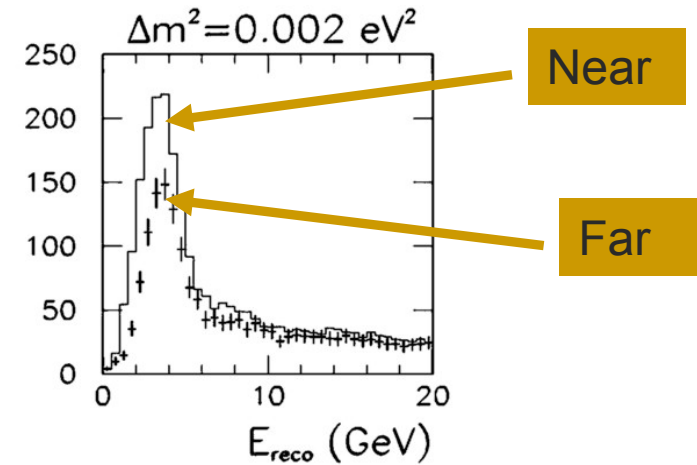
University of Oxford &  
Rutherford Appleton Laboratory  
IoP Particle Physics 2004  
Tuesday 6<sup>th</sup> April

# [ Talk Outline ]

- The MINOS Experiment
- Near, Far and Calibration Detectors
- Calibrating the Calorimeter
- Stopping Muons: Our Standard Candle

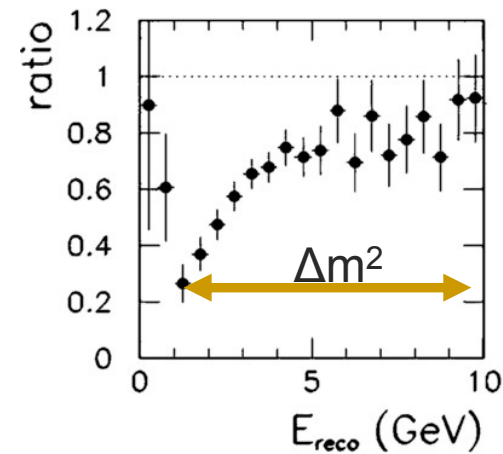
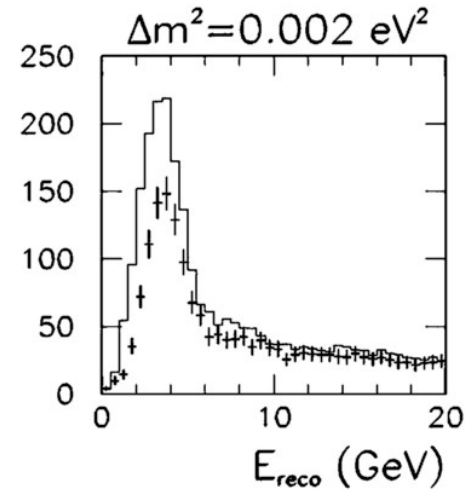
# MINOS Neutrino Physics

- Long-baseline neutrino oscillation experiment.
- Compare Near and Far neutrino energy spectra.
- Looking for  $\nu_\mu$  disappearance as a function of energy – “the dip”.
- Very important that we can accurately reconstruct the neutrino energy.



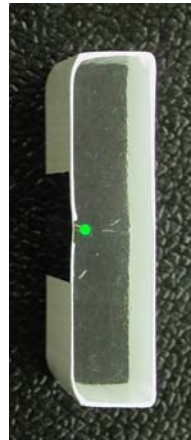
# Calibration Targets

- Will hopefully have statistics for 10% measurement of  $\Delta m^2$ .
- Have a target of 5% for the absolute energy calibration.
- Require that the relative calibration between detectors is 2%.



# [ MINOS Detectors (x3) ]

- Common features:
  - Alternating steel-scintillator, magnetised, tracking calorimeters
  - WLS fibre-optic cables to extract the scintillator light

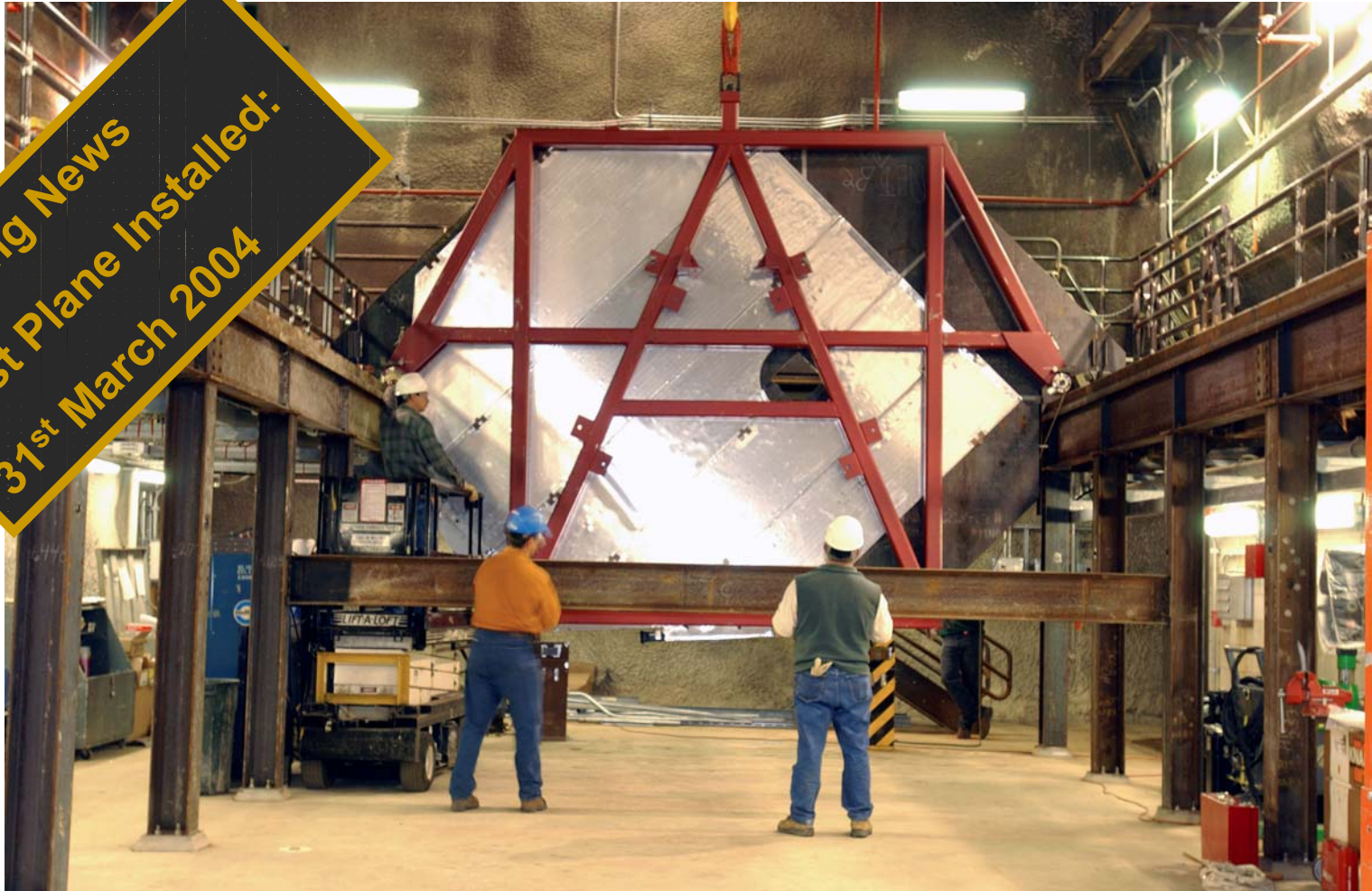


Far Detector plane under construction



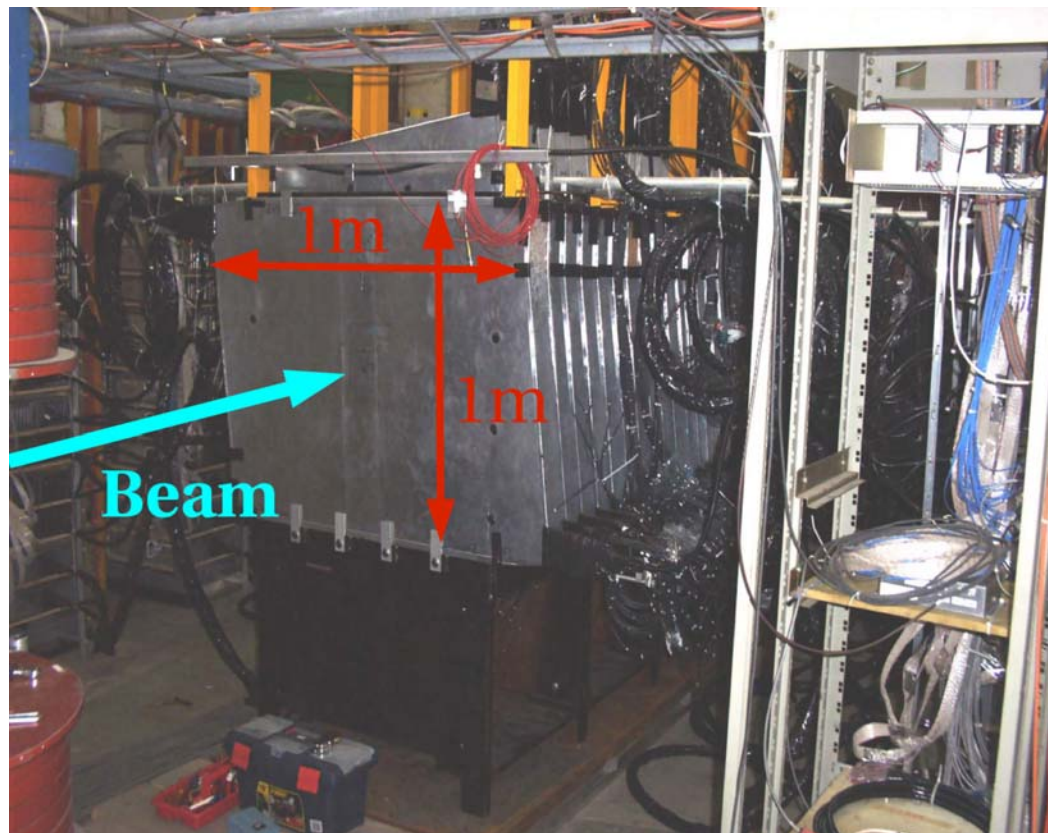
# Near Detector

Breaking News  
First Plane Installed:  
31st March 2004



# Calibration Detector at CERN (CalDet)

- Small version of Near and Far detectors
- Exposed to  $\mu$ ,  $p$ ,  $e$ ,  $\pi$  beams 0.5-10 GeV
- Hadronic & EM energy response
- Tune the MC



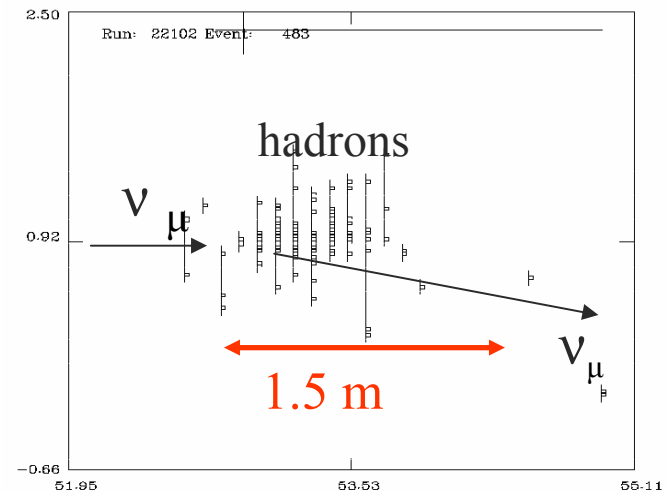
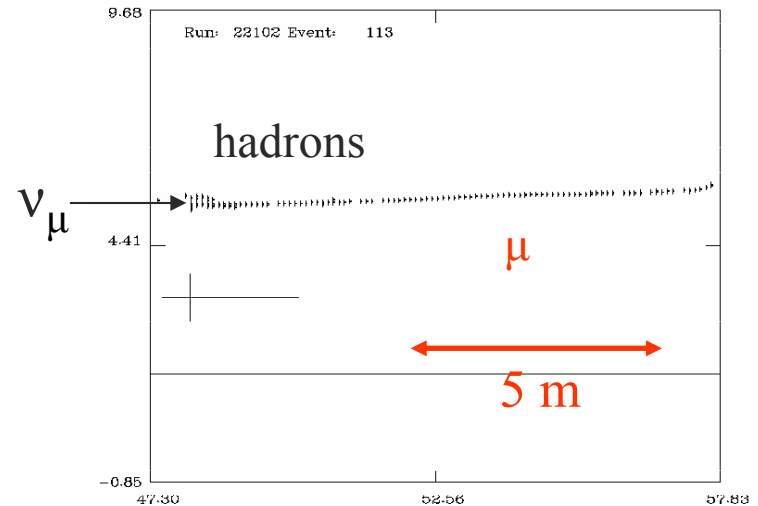
# Neutrino Energy Reconstruction

- Need to accurately determine original neutrino energy to see “the dip”.

$$E_{\nu} = E_{\text{Had}} + E_{\mu} \quad (\text{Visible})$$

- $E_{\mu} = f(\text{Range}) \ \& \ f(\text{B-Curvature})$

- $E_{\text{Had}} = f(\text{Energy Dep.})$





# Calibrating the Calorimeter

## Problems:

- Scintillator output varies by up to 20% from strip to strip.
- Light output depends on particle type and particle energy.



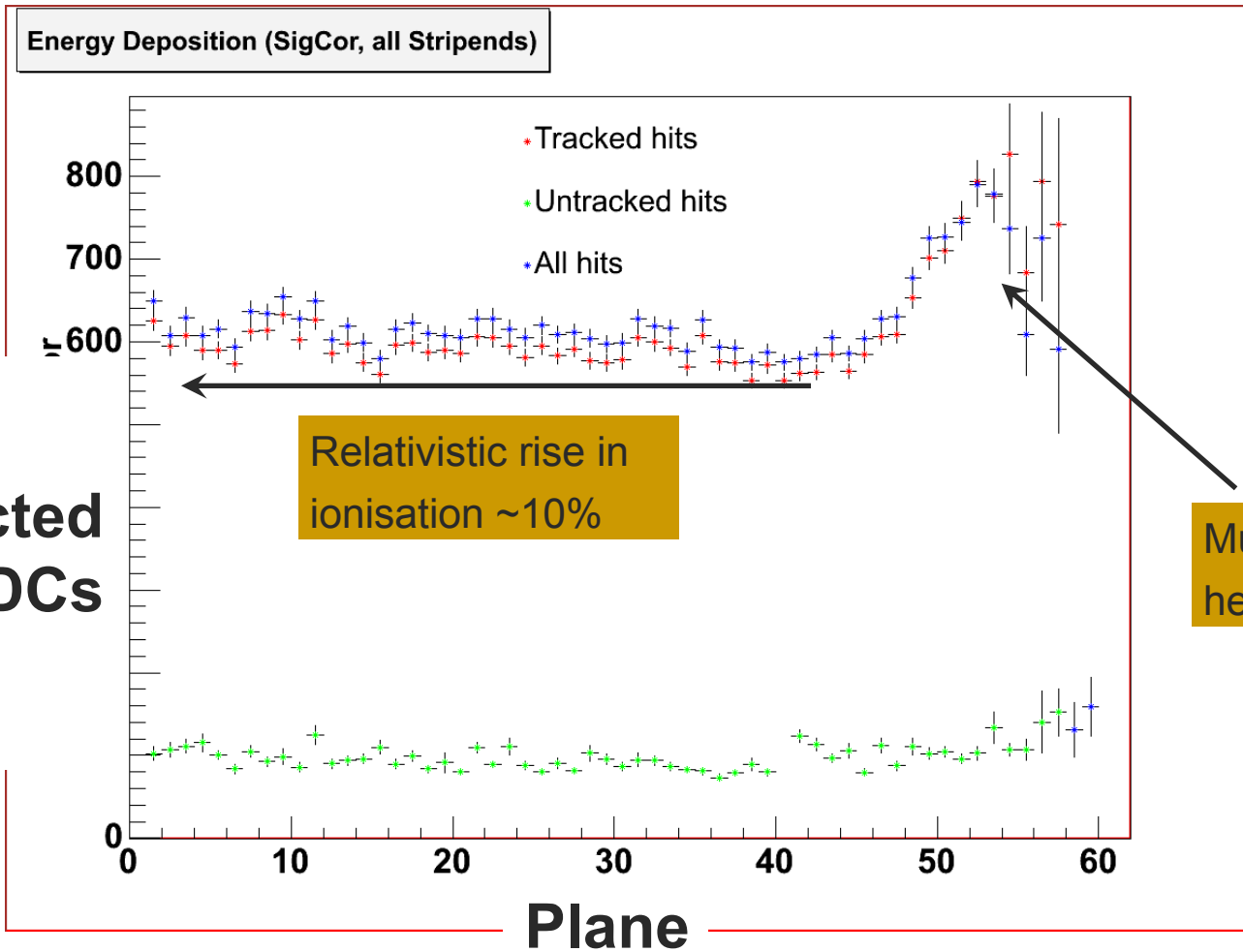
## Solutions:

- Normalise using cosmic muons.
- Measure response with Calibration Detector in MEU/GeV. (MEU=Muon Energy Unit)

# Why Stopping Muons?

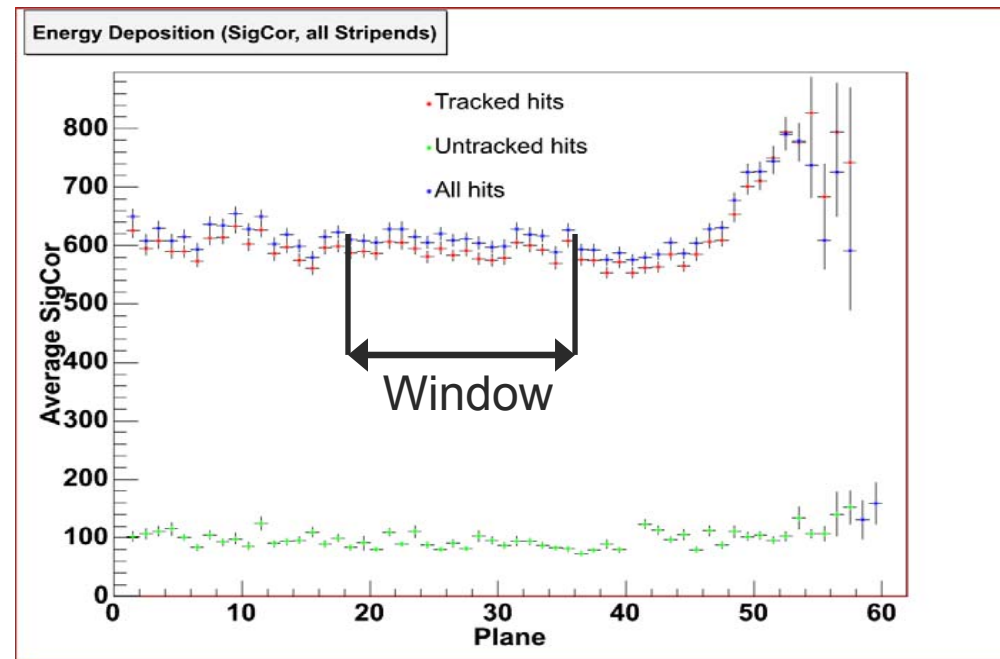
- Need a calibration source of known energy - a “standard candle”.
- Don’t have a “PS” accelerator ½ mile underground to provide beams of particles
- The energy spectrum of cosmic muons (and hence energy deposition) is different for Near and Far detectors and not known accurately enough.
- Working backwards from the end of a stopping muon track you know its energy – Use to define a Muon Energy Unit (MEU).
- Stopping muons are the same at all 3 detectors!

# Stopping Beam Muons in CalDet (at 1.8 GeV/c)



# [ Track Window Method ]

- Want a method that is inherently robust and not sensitive to reconstruction problems.
- Sum up the energy deposition in a window.
- Define a Muon Energy Unit (MEU) to be the average energy deposited in the window by a perpendicular muon in one detector plane.



# Conclusion

- The main physics measurement requires accurate reconstruction of the neutrino energy to 5%.
- Explained how MINOS has 3 detectors in very different environments. Need to set the absolute energy scale to be the same for each detector.
- No beam source of GeV particles for calibration underground so use stopping cosmic muons.
- Take a window on the stopping muon track to measure the detector response.
- Should enable us to meet our 2% relative calibration target and hence accurately measure  $\Delta m^2$ .



**The End!**