

# *Infrared Finite Scattering Amplitudes*

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## Overview

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- A *very* brief overview of calculating physical observables and scattering amplitudes.
- Highlight the source of infrared (IR) singularities.
- Discuss ways of avoiding such IR singularities, **IR finite scattering amplitudes**.
- An example demonstrating **IR finite scattering amplitudes** in the calculation of the total cross section for  $e^+e^- \rightarrow 2 \text{ Jets}$  at NLO.

## Calculation of physical observables

- The aim of theoretical calculations is to produce predictions of the **physical observables** measured in experiments.
- This is achieved by integrating over the *phase space* with the amplitude “squared” times some definition function for the *physical observable* being calculated,

$$\int dLips |A|^2 \times J(k_i, \dots)$$

- The important part here is the matrix element  $A$ , this is calculated from the  **$S$ -Matrix**,

$$A = \langle out | S | in \rangle$$

- The  $S$ -Matrix maps the basis of the initial states onto the final states and contains all the details of the **interactions**.

# The $S$ -Matrix

- Scattering calculations are performed by calculating the overlap of an *initial* and *final* eigenstate of the **full** Hamiltonian,  $H$ ,

$$\langle \Psi_{out}(\infty) | \Psi_{in}(-\infty) \rangle = \langle \Psi_{out} | U(\infty, t') U^\dagger(-\infty, t') | \Psi_{in} \rangle = \langle \Psi_{out} | S | \Psi_{in} \rangle$$

- To calculate this we usually place it in the **interaction picture**,

$$\begin{aligned} \langle \Psi_{out} | U_{H_0}(\infty, t') U_{H_0}^\dagger(\infty, t') S U_{H_0}(-\infty, t') U_{H_0}^\dagger(-\infty, t') | \Psi_{in} \rangle \\ = \langle \psi_{out}^I | S^I | \psi_{in}^I \rangle \end{aligned}$$

- The *in* and *out* fields are now eigenstates of the **free** Hamiltonian,  $H_0$ .
- The  $S$ -Matrix now evolves in time with the **interaction** Hamiltonian,  $H_I = H - H_0$ .
- For this transformation to be valid the **interaction** must “*turn off*” at infinity.

# Infrared divergences

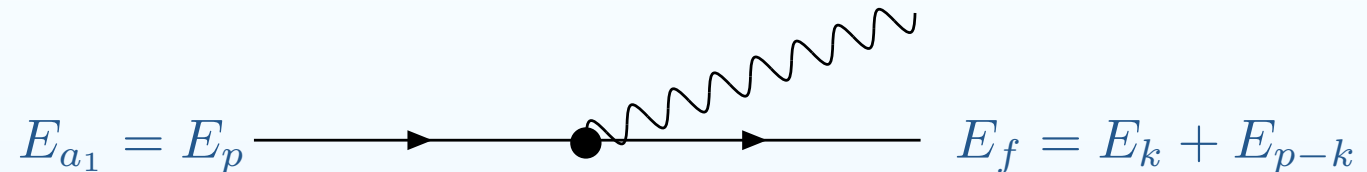
- The  $S$ -Matrix is then expanded perturbatively in terms of some parameter  $\alpha$ .
- At each order in the **time ordered** perturbative expansion we have terms of the form,

$$\alpha^n \int_{-\infty}^{\infty} dt_1 H_I e^{-it_1(E_f - E_{a_1})} \dots \int_{-\infty}^{t_{n-1}} dt_n H_I e^{-it_n(E_{a_n} - E_i)}$$

- For the perturbative expansion to make sense these *time* integrals must **converge**. [P.P.Kulish & L.D.Faddeev Theor.Math.Phys. **4**, 745 (1970)]
- For the interaction to “*turn off*” we require  $\exp(-it_1(E_f - E_{a_1}))$  to vanish at  $t_1 \rightarrow \pm\infty$ , this is usually achieved by using an *adiabatic* factor.
- We can quickly see though that if  $E_f - E_{a_1} = 0$ , then  $\exp(-it_1(E_f - E_{a_1})) \rightarrow 1$  and this integral will in fact **diverge**.

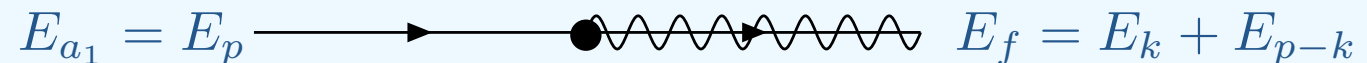
# The appearance of infrared divergences

- Situations where  $E_f - E_{a_1} = 0$  can occur are,
  - **Soft** emission,  $E_{p-k} + E_k - E_p \rightarrow 0$  as  $E_k \rightarrow 0$ ,



- **Collinear** emission,

$$E_{p-k} + E_k - E_p \rightarrow (1 - \lambda)E_p + \lambda E_p - E_p = 0 \text{ as } E_k \rightarrow \lambda E_p,$$



- These divergences will **only** appear when the particles involved are “*on-shell*”, so **IR divergences** will only appear in **initial** and **final** states (as internal particles are “*off-shell*”).
- States at infinity are **not** free and well separated. The interaction does **not** “*turn off*”.

## So whats going on?

- The problem is that the energy of the states  $|p, k\rangle_{E_k \rightarrow 0, E_k \rightarrow \lambda E_p}$  is **degenerate** with those of  $|p\rangle$ .
- By assuming that the *in* and *out* states are eigenstates of  $H_0$  we mistakingly split up states that “look” the same.
- A possible **solution** is to use **asymptotic states** that are eigenstates of the **asymptotic** Hamiltonian,  $H_A = H_0 + H_{IR}$ , these *combine*  $H_0$  states with the same energy into *one* state.

$$|\{p\}\rangle \equiv \text{---} \rightarrow \text{---} + \text{---} \rightarrow \bullet \text{---} \rightarrow \text{---}$$

- We can do this by switching to the **asymptotic interaction picture**. The *in* and *out* states then become **asymptotic** states.
- **Energy** is now an eigenvalue of the **asymptotic** Hamiltonian  $H_A$  and  $E_f - E_{a_1} \neq 0$  as the state  $\langle\{f\}| \neq \langle\{a_1}\|$ , leading to a finite result as **degenerate** states are **no longer separate**.

## Infrared finite scattering amplitudes

- **Problem!** this is *very hard* to do. We need to solve for the eigenstates of  $H_A$ , which we do **not** know how to do.
- Instead, perform a different **unitary** transformation on the  $S$ -Matrix. Changing into the **asymptotic interaction** picture involves,

$$A = \langle \psi_{out}^I | U_{H_A-H_0} U_{H_A-H_0}^\dagger S^I U_{H_A-H_0} U_{H_A-H_0}^\dagger | \psi_{in}^I \rangle = \langle \psi_{out}^A | S_A^A | \psi_{in}^A \rangle$$

Instead **remain** in the **interaction picture** but alter the  $S$ -Matrix,

$$\mathcal{A} = \langle \psi_{out}^I | U_{H_A-H_0}^\dagger S^I U_{H_A-H_0} | \psi_{in}^I \rangle = \langle \psi_{out}^I | S_A^I | \psi_{in}^I \rangle$$

- *Physical observables* remain **unaltered** by this transformation as  $\int |A|^2 \equiv \int |\mathcal{A}|^2$ , only the amplitudes **change**.
- The amplitudes,  $\mathcal{A}$ , are called **infrared finite scattering amplitudes** and it is known that these are *also free* of **IR singularities**. [J.Frenkel et.al. Nucl. Phys. B194, 172 (1982) and others]



## How does this work?

- **IR finite scattering amplitudes** differ from the usual approach by the basis of states used,
  - The  $S$ -**Matrix** usually acts on the **complete** basis of  $H_0$  states including the **degenerate** states.
  - The  $S_A$ -**Matrix** acts on a **reduced** basis of  $H_0$  states that does **not** include any of the **soft** or **collinear** states. So the amplitude is finite.
- So we can use our knowledge of **energy** eigenstates of  $H_0$  to produce amplitudes that are **free** of IR divergences.
- There is one problem, unlike in the usual case where  $[S, H_0] = 0$ , **IR finite scattering amplitudes** use  $S_A$  and  $[S_A, H_0] \neq 0$  and so we do not expect the amplitude to contain a single energy conserving delta function.

## An example: $e^+e^- \rightarrow 2 \text{ Jets}$ at NLO

- Check that **IR finite scattering amplitudes** give the correct answer for  $e^+e^- \rightarrow 2 \text{ Jets}$  at NLO. [Forde & Signer Nucl. Phys. B**684** 125, (2004), arXiv:hep-ph/0311059]
- We choose a definition of  $S_A$  that **removes** all the *singularities* in the three point vertex. This once chosen is **general** for **all** processes involving three point vertices.
- There are **two** amplitudes which will contribute to this process.
  - $\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)$ , which has **one** incoming **photon** and **two** outgoing **“quarks”**.
  - $\mathcal{A}(\{q(p_1), \bar{q}(p_2), g(p_3)\}; \gamma)$ , has **one** incoming **photon**, **two** outgoing **“quarks”** and **one** outgoing **“gluon”**.
- They will both be IR finite unlike the usual technique.

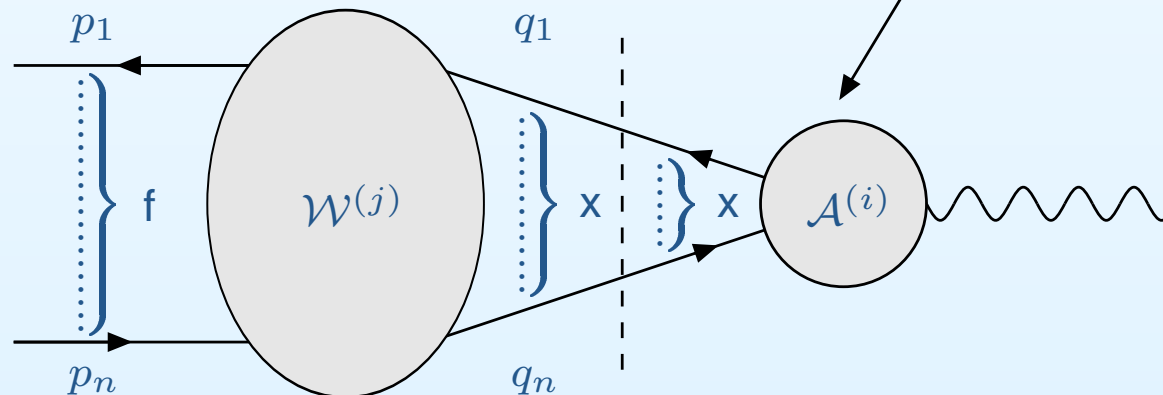
## An example: $e^+e^- \rightarrow 2$ Jets at NLO

- Unlike the normal situation we do not want to calculate this **directly**, instead we rewrite the amplitude as,

$$\langle f|S_A|i\rangle \equiv \langle f|U_{H_A-H_0}^\dagger S U_{H_A-H_0}|i\rangle = \langle \{f\}|S|\{i\}\rangle$$

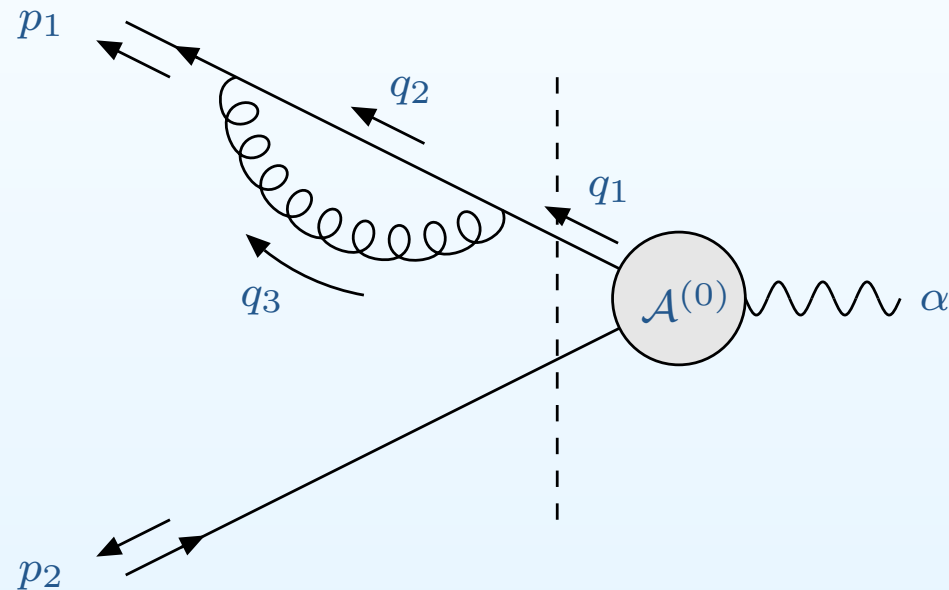
- So the **amplitudes** can be split up into **sub-amplitudes** as,

$$\mathcal{A}(\{f\}; \{i\}) = \langle \{f\}|S|\{i\}\rangle = \sum_x \underbrace{\langle f|U_{H_A-H_0}^\dagger|x\rangle}_{\mathcal{W}^{(j)}} \underbrace{\langle x|S|\gamma\rangle}_{\mathcal{A}^{(i)}}$$



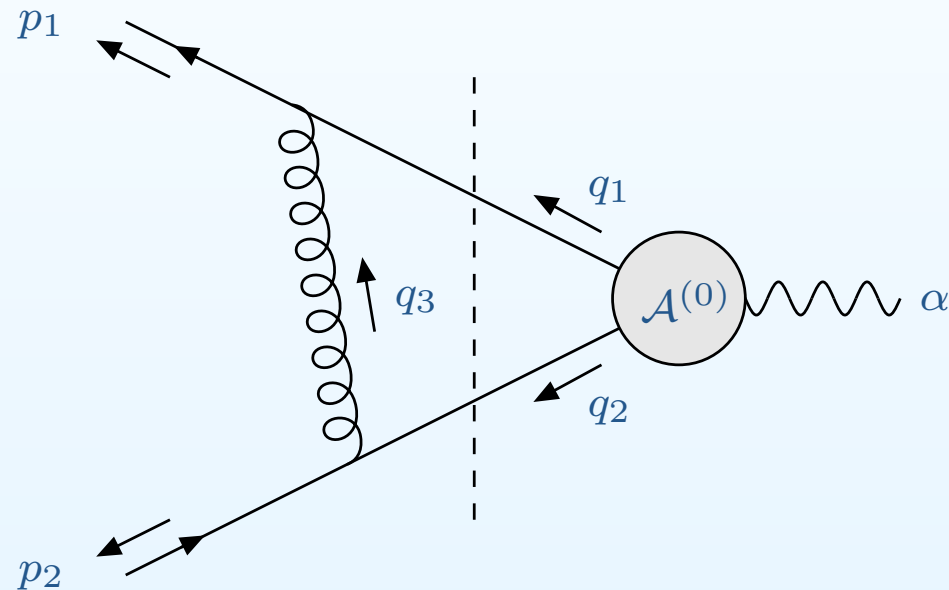
## IR free amplitudes at $\mathcal{O}(g^2)$ for $e^+e^- \rightarrow 2$ Jets at NLO

- $\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)$  is made up of **seven sub-amplitudes**.
- The **first** pair of *sub-amplitudes* is given by the **self-interaction** terms, one for each leg,



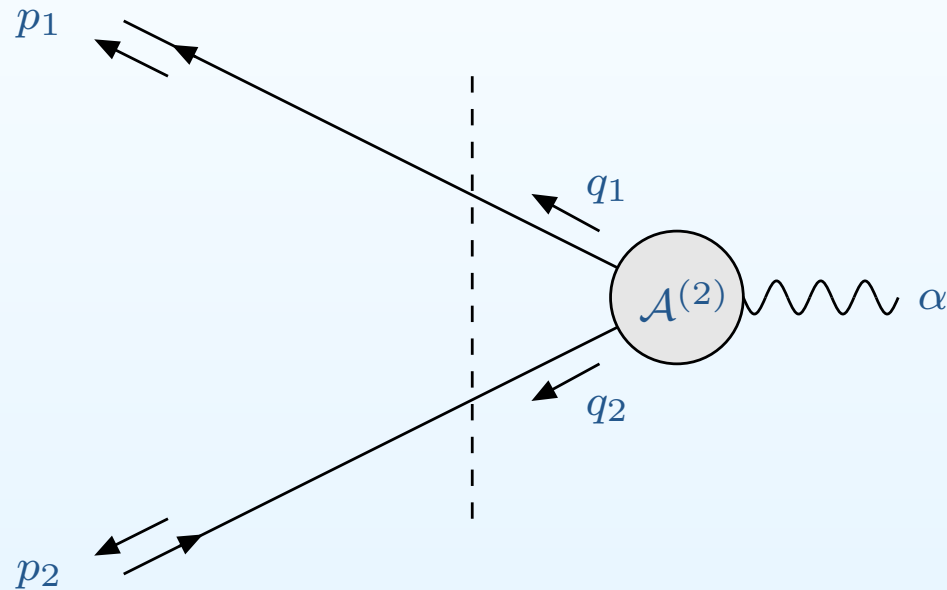
## IR free amplitudes at $\mathcal{O}(g^2)$ for $e^+e^- \rightarrow 2$ Jets at NLO

- $\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)$  is made up of **seven sub-amplitudes**.
- The **second** pair of *sub-amplitudes* is given by the 1-gluon exchange diagrams, one for each time ordering,



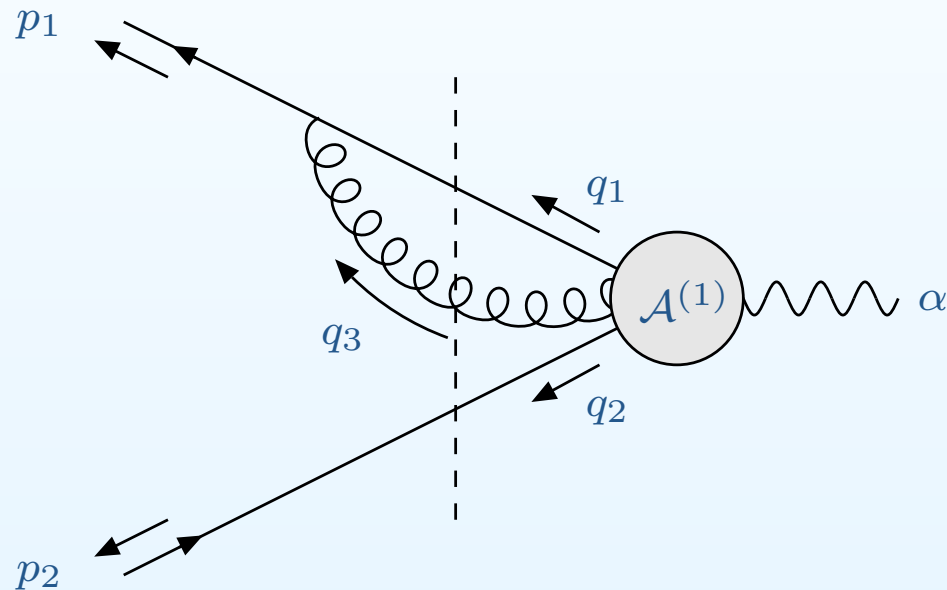
## IR free amplitudes at $\mathcal{O}(g^2)$ for $e^+e^- \rightarrow 2$ Jets at NLO

- $\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)$  is made up of **seven sub-amplitudes**.
- The **fifth sub-amplitude** is given by the usual virtual correction term,



## IR free amplitudes at $\mathcal{O}(g^2)$ for $e^+e^- \rightarrow 2$ Jets at NLO

- $\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)$  is made up of **seven sub-amplitudes**.
- The **final** pair of *sub-amplitudes* is given by the two 3-particle cut diagrams,



## An IR Finite Amplitude

- We can now **combine** these *sub-amplitudes* together.
- All the infrared singular pieces **cancel** between the *sub-amplitudes*, leaving an **infrared finite amplitude**,

$$\begin{aligned}
 \mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma) = & \\
 & \left( 1 + C_F \left( \frac{\alpha_s}{2\pi} \right) \left( g_1(\Delta) + g_2(\Delta) + g_3(\Delta) - 4 + \frac{\pi^2}{12} \right) \right) \\
 & \mathcal{A}^{(0)}(q(p_1), \bar{q}(p_2); \gamma(P)) \\
 & + (-ie) \delta_{ij} \langle p_1 | \gamma^\alpha | p_2 \rangle (2\pi)^{(D-1)} \delta^{(D-1)}(\vec{P} - \vec{p}_1 - \vec{p}_2) \\
 & \int d\tilde{q}_3 \left( f_1(p_1, p_2, q_3) \delta(\sqrt{S} - \omega(\vec{p}_1) - \omega(\vec{p}_2)) + f_2(p_1, p_2, q_3) \right. \\
 & \left. + f_3(p_1, p_2, q_3) \right)
 \end{aligned}$$



# A Cross Section Calculation

- Using the two **IR finite amplitudes**,  $\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)$  and  $\mathcal{A}(\{q(p_1), \bar{q}(p_2), g(p_3)\}; \gamma)$ , we can calculate the **total cross section** for  $e^+e^- \rightarrow 2$  Jets at NLO.
- The cross sections are,

$$\begin{aligned}\sigma_{\{q\bar{q}\}} &= \int |\mathcal{A}(\{q(p_1), \bar{q}(p_2)\}; \gamma)|^2 \\ &= \left(1 + C_F \frac{\alpha_s}{\pi} \left(-\frac{1}{2} + \log 4 - \frac{3}{2} \log\left(\frac{\Delta}{2}\right) - \log^2\left(\frac{\Delta}{2}\right)\right)\right),\end{aligned}$$

$$\begin{aligned}\sigma_{\{q\bar{q}g\}} &= \int |\mathcal{A}(\{q(p_1), \bar{q}(p_2), g(p_3)\}; \gamma)|^2 \\ &= C_F \left(\frac{\alpha_s}{\pi}\right) \left(\frac{5}{4} - \log 4 + \frac{3}{2} \log\left(\frac{\Delta}{2}\right) + \log^2\left(\frac{\Delta}{2}\right)\right).\end{aligned}$$

- Adding these two pieces together gives,

$$\frac{1}{\sigma_0} \sigma = \frac{1}{\sigma_0} (\sigma_{\{q\bar{q}\}} + \sigma_{\{q\bar{q}g\}}) = \left(1 + \frac{\alpha_s}{\pi} + \mathcal{O}(\alpha_s^2)\right)$$

the same as the usual result given by  $\sigma = \sigma_{q\bar{q}} + \sigma_{q\bar{q}g}$ .

## Conclusion

- Given a *very brief* overview of infrared singularities
- Shown how **IR finite amplitudes** can be produced.
- Shown how this technique can be **applied** to an example process,  $e^+e^- \rightarrow 2 \text{ Jets}$  at NLO, producing the **known** result but using amplitudes that are **free** of IR singularities.
- Future work,
  - Develop the technique for **direct** calculations of **IR finite amplitudes** without the intermediate sub-amplitudes.
  - Develop techniques to calculate *physical observables* from these amplitudes.
  - ...