



# Commissioning with Physics Data at ATLAS

### Dan Tovey for the ATLAS Collaboration

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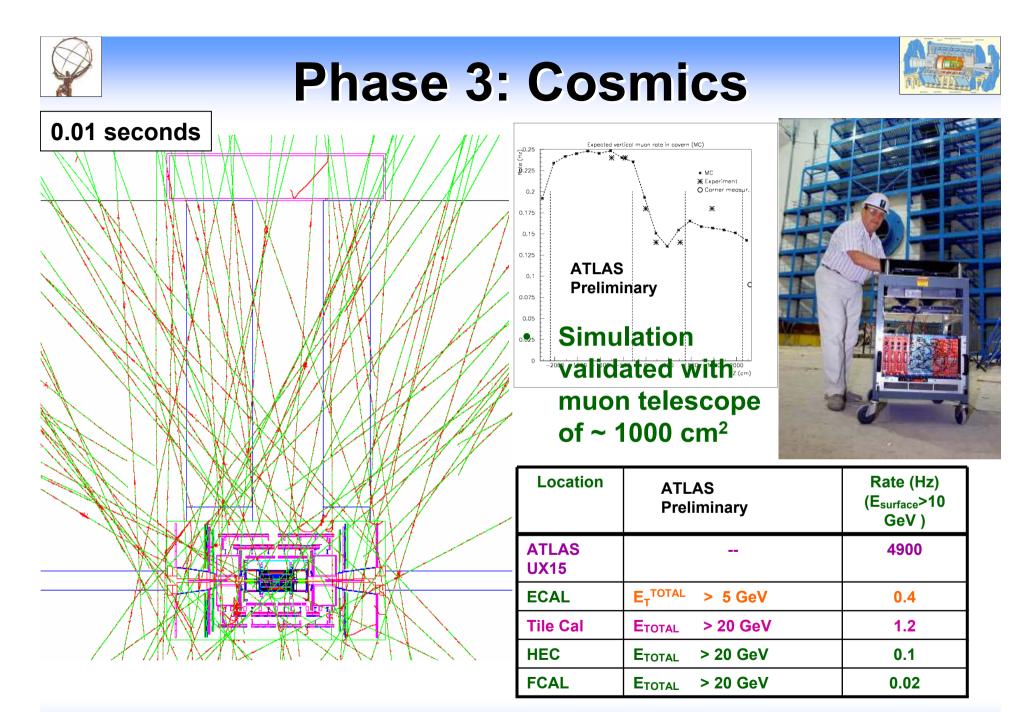


# Strategy



- Commissioning with physics data proceeds in four phases:
  - <u>Phase 3</u> : Cosmics running
    - ➔ initial physics alignment / calibration of the detector
    - ➔ debugging of sub-systems, mapping dead channels etc.
  - Phase 4 : One beam in the machine
    - ➔ beam-halo muons and beam-gas events
    - ➔ more detailed alignment / calibration etc.
  - <u>Phase 5</u> : First pp collisions : prepare the trigger and the detector
    - → tune trigger menus / measure efficiencies
    - ➔ begin to measure reconstruction efficiencies, fake rates, energy scales, resolutions etc.
  - <u>Phase 6</u> : Commissioning of physics channels
    - → Improve measurements
    - → begin to understand backgrounds to discovery channels ...
- Thinking now about what we can learn in each phase / how to use
  the data in practice
- Will give a few examples of recent work / work in progress ...

Pre-collision Phases



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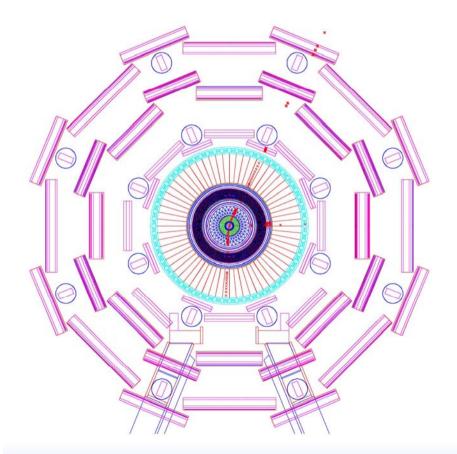
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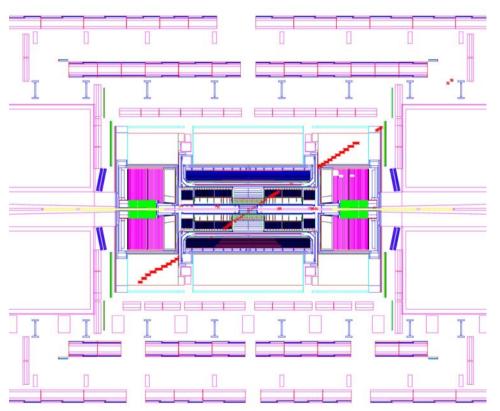


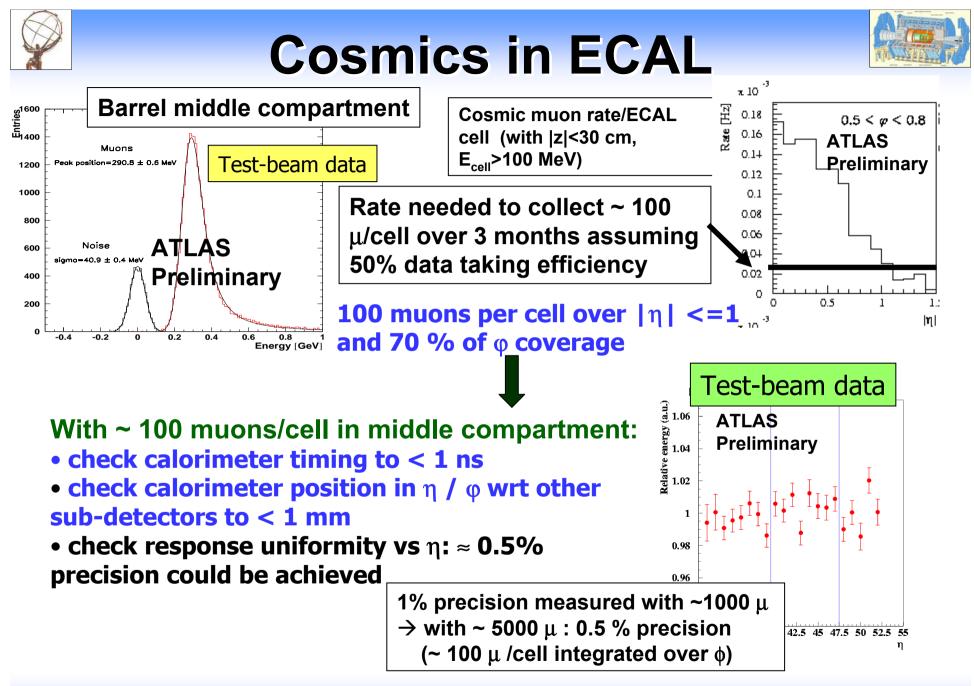
### 'Typical' Event



- One track reconstructed in Muon chambers
- Two tracks reconstructed in Inner Detector
- Will happen every ~ 10 s



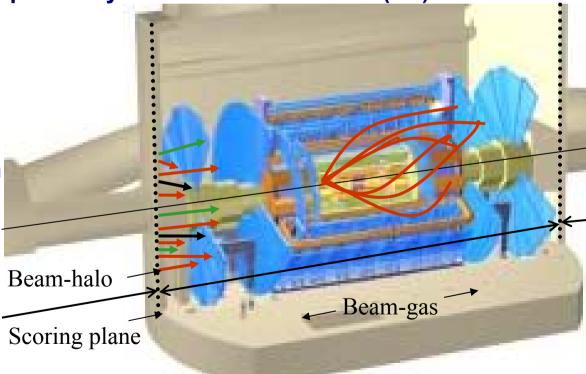




# Phase 4: Single-beam period



- Beam-halo
  - Low  $p_T$  particles from machine.
  - Simulation of machine background by machine experts (V. Talanov):
    - based on MARS; machine optics V 6.4
    - scoring plane at the cavern entrance before ATLAS shielding (z =  $\pm$  23 m from IP)
  - Then particles are transported by ATLAS full simulation (G3)
- Beam-gas
  - Vacuum not perfect
  - p(7 TeV) on p(rest)
  - vertices uniformly distributed over ± 23 m
  - σ(pH, pC, pO, ...) ∝
     σ(pp)xA<sup>0.7</sup> (inelastic only)
  - vacuum estimate:
     ~3x10<sup>-8</sup> Torr (~10<sup>15</sup> mol/m<sup>3</sup>)



# **Cosmics & Beam Gas in ID**

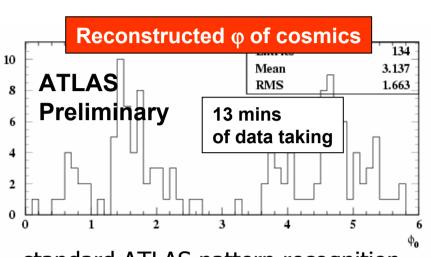
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### Cosmics : O (1Hz) tracks in Pixels+SCT+TRT

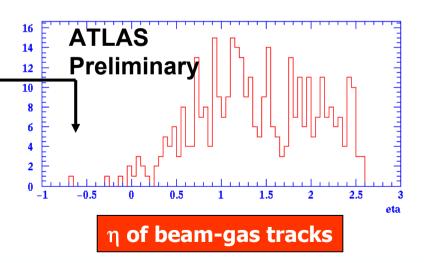
- useful statistics for debugging readout, maps of dead modules, etc.
- check relative position Pixels/SCT/TRT and of ID wrt ECAL and Muon Spectrometer
- first alignment studies: may achieve statistical precision of ~ 10  $\mu m$  in parts of Pixels/SCT
- first calibration of R-t relation in straws

#### Beam-gas :

- ~ 25 Hz of reconstructed tracks with p<sub>T</sub> > 1 GeV and |z|<20 cm</li>
- $\rightarrow$  >10<sup>7</sup> tracks (similar to LHC events) in 2 months
- enough statistics for alignment in "relaxed" environment → exceed initial survey precision of 10-100 µm



standard ATLAS pattern recognition (no optimisation for cosmics ...)



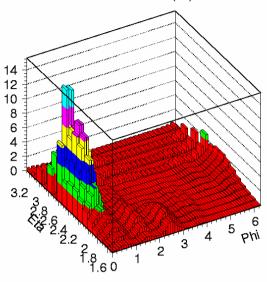




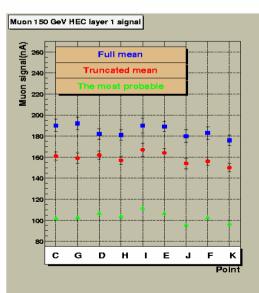
### **Beam Halo in HEC**

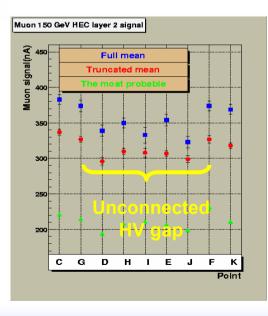


Hit rate distribution (Hz)



- Halo muons:
  - essentially parallel to z-axis
  - look much like test-beam  $\mu$  (esp for endcap)
- HEC-standalone efficiency for muon identification: ~ 25%, S/N ~ 4
  - Max(Min) Rate ~ 3(0.02) Hz / cell
    - 5 x 10  $^{6}$  (3 x 10  $^{4}$  )  $\mu$  in 2 months @ 30 %



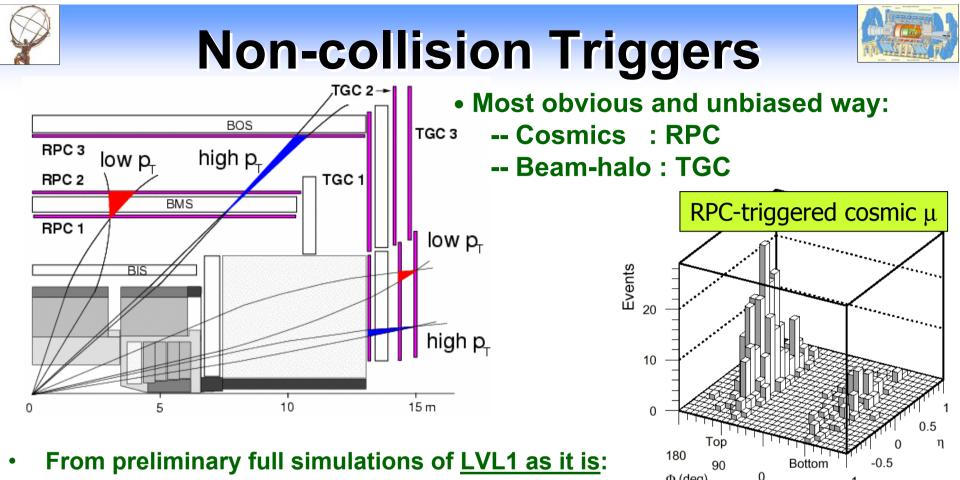


#### From test-beam studies:

- Cell timing: <1 ns</p>
- Cell Energy: <1 %</p>
- Cell, module and wheel alignment: few mm
- Detect unconnected HV gaps

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- -- Cosmic muons : ~ 100 Hz pass low- $p_T$  RPC LVL1 trigger
- -- Beam-halo muons : ~ 1 Hz pass low- $p_T$  TGC LVL1 trigger
  - small enough → not worrying for LHC data taking
  - high enough → useful samples (e.g. > 10<sup>8</sup> cosmics evts in 3 months if ε=50%) for commissioning (triggered muons cross the interaction region)
- Also studying ways of increasing trigger rates during commissioning (dedicated TileCal cosmic trigger, min bias scint. planes in forward regions)

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-90



### **Phase 5: First Collisions**



- With first collisions will begin to understand / calibrate physics objects.
- Assume the detector is already ready for data taking
  - Calorimeters set to EM scale
  - Readout channels reasonably intercalibrated (electronics, cosmics, Cs etc)
  - Hadronic response set with weighting techniques in MC or from testbeam.
  - ID & muon system aligned roughly (initial survey, cosmics etc.)
- Aim to measure
  - energy scales,
  - Resolutions
  - Efficiencies
  - Fake rates etc.
- Requirements from physics e.g.:
  - 0.1% for the electron/muon energy/momentum scale
  - 1% for the jet energy scale
  - Also, uniformity .....
- Initially won't have this precision, e.g. 0.5% for muons from initial field maps and survey, 1-2% for EM from test-beam, 5-10% for JES from test-beam/MC.



### Strategy



- Use isolated charged tracks (e.g. from  $\tau$  decays) to
  - cross-check pre-collision alignment,
  - determine E/p matching precision,
  - determine hadronic energy scale,
  - Intercalibrate calorimeters
- Use  $J/\psi$  (low  $p_T$ ) and  $Z^0$  (high  $p_T$ ) with mass constraint to
  - (Inter)calibrate LAr EM
  - Calibrate e/µ E/p scales
- Use W mass constraint in W→jj from ttbar production to set JES.
- Use  $Z^0/\gamma$  + jet events to calibrate across calorimeters (cracks, dead material) and monitor.
  - $p_T$  balance between jet Z<sup>0</sup>/ $\gamma$
- Later use  $Z^0/\gamma$  + b-jet events to calibrate b-JES. Also measure b-tagging efficiency in situ with top events.

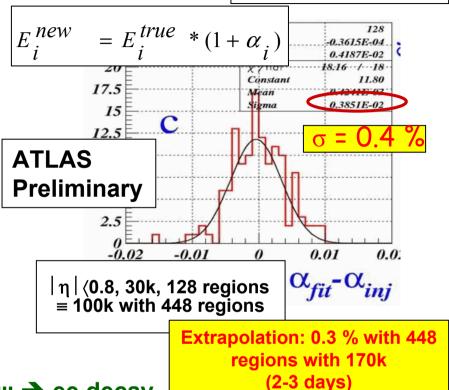


#### LArEM Intercalibration Work in progress



#### From hardware and beam tests: calibration known to 0.5-0.6 % inside 448 windows of $\Delta \eta \times \Delta \Phi = 0.2 \times 0.4$ inside $|\eta| \langle 2.5$

- Need 0.3% intercalib to achieve 0.7% • global constant term
- Use real data to intercalibrate ( $Z^0$ ,  $J/\psi$ , ۲ electron E/p, inclusive  $p_{T}$  distributions, photon conversions,...)
- $Z \rightarrow ee decay$ •
  - High rate (0.5-1 Hz), low background, easy trigger, uniform in  $\eta$  and  $\varphi$ , well known process, 2 correlated electromagnetic objects...
  - Define reference M<sub>ee</sub> distributions and \_ fit to invariant mass of e+e- in given pair of regions by tuning regional 'decalibration' coefficients  $\alpha_i$



#### $J/\psi \rightarrow ee decay$

- ~ 5\*10<sup>5</sup> J/ $\psi$  in 1 year of low lumi (reconstruction eff=20%), trigger  $pt(\mu) > 6 GeV$
- Gives check on linearity at low energy
- **Expected** intercalibration precision of 0.6%

# **EM**/ $\mu$ Scales from Z<sup>0</sup>

0.11

0.108

0.106

0.104

0.102

0.098

0.096

0.094

0.092

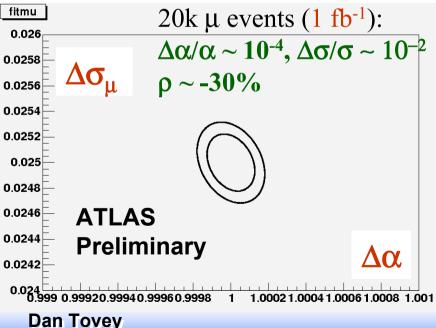
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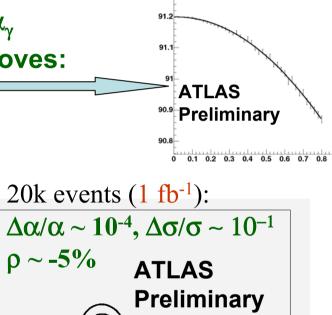
0.1

 $\Delta \sigma_{e}$ 

0.09 0.999 0.99920.99940.99960.9998

- Measure  $e/\mu/\gamma$  energy scales using  $Z \rightarrow ee(\gamma)/\mu\mu(\gamma)$ .
- Create reference distributions for each channel
- Then minimize  $\chi^2$  comparing reference distributions and data varying the e,  $\mu$ ,  $\gamma$  E/p scales  $\alpha_e$ ,  $\alpha_\mu$ ,  $\alpha_\gamma$
- Also consider <u>concurrently</u> as accuracy improves:
  - resolution effects (can shift peak)
  - PDFs (""")
  - FSR (" " ")





Work in progress

lass Shift vs. Resolution





1.0002 1.0004 1.0006 1.0008 1.001

1

Λα



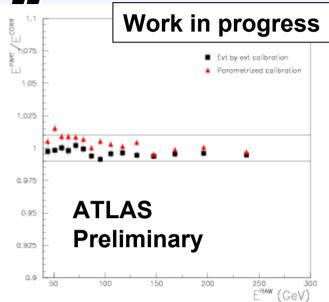
### JES from W→jj

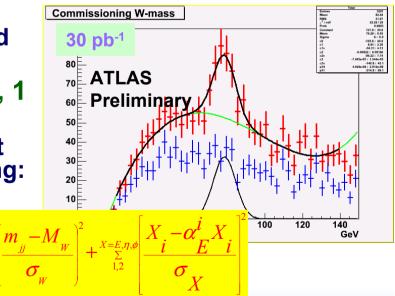


 Use the mass constraint of the W in ttbar events, to set the JES / rescale jet to parton energy α = E<sub>parton</sub> / E<sub>jet</sub>

$$M_{jj} = \sqrt{2E_{j1}E_{j2}(1-\cos\theta_{j1j2})} = M_W$$

- Take into account E,  $\eta$  and  $\phi$  in the minimization procedure and corrected energies and angles.
- E of parton and jet agree within ~ 1% over the range 50-250 GeV
- Pros: Good statistics, easily triggerable, small physics backgrounds.
- Cons: Only light q jets, limitations in E and  $\eta$  reach.
- More recently: investigating cases with 0, 1 or 2 b-tags.
  - Consider more sophisticated approach: fit to W mass dist rather than simple rescaling:
  - Takes into account variation of rescaling parameter with energy and correlation between energies and opening angle.





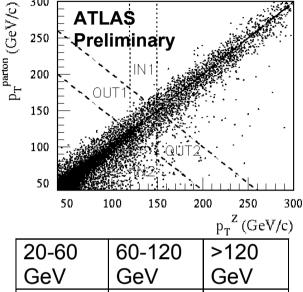


# JES with γ/Z<sup>0</sup>+jet



- Use the p<sub>T</sub> balance between Z or photon (precisely measured) and highest p<sub>T</sub> jet
  - Reconstructed jet  $p_T$  rescaled to balance the Z  $p_T$ .
- Distribution systematically skewed, esp by ISR (and FSR)
- <u>Pros</u>:
  - Enlarged E and (especially) η reach wrt W**→**jj,
  - includes 6% of b-jets,
  - potentially large statistics available: γ+jet with p<sub>T</sub>>20
     GeV: ~10K events/min. (not incl. eff. & trigger)
- <u>Cons</u>:
  - Easy to introduce biases in the selection procedure,
  - sensitivity to ISR modeling, esp at low p<sub>T</sub>,
  - background to the  $\gamma$  or Z<sup>0</sup> may bring additional bias
  - p<sub>T</sub> range covered with good statistics limited.
  - The effect of the trigger has also to be considered (standard menu or downscaled)
- Also : dijet calibration, E<sub>T</sub><sup>miss</sup> projection method
- Also use Z<sup>0</sup> + b-jet to calibrate b-JES Work in progress

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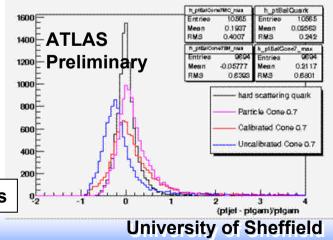


#### 1% difficult below 60 GeV

0.004

0.015

0.049

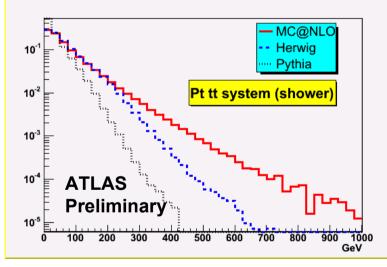




# Phase 6: First Physics



- Vast topic in principle as many background estimation techniques as analyses
- In practice large degree of commonality, although different emphases.
- Need to
  - Minimise most poorly estimated backgrounds (at expense of statistics?);
  - Estimate remaining backgrounds from combination of data and MC;



Large differences between NLO/LO MC codes → Use even NLO codes with caution!

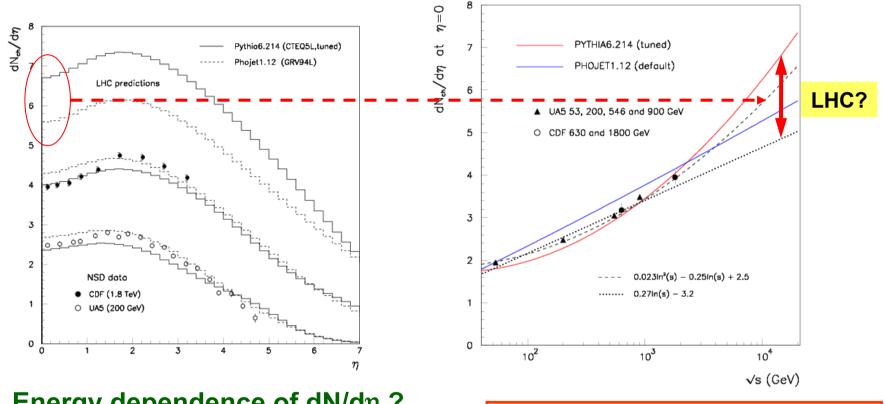
• We can learn a lot from Runll <u>but</u> one big difference:

There will have been no previous measurements at similar  $\sqrt{s!}$ 

• Will concentrate on three case studies: Min bias, Top and SUSY







- Energy dependence of dN/dη ?
- Vital for tuning UE model (see later)
- Only requires a few thousand events.

• PYTHIA models favour ln<sup>2</sup>(s);
• PHOJET suggests a ln(s) dependence.

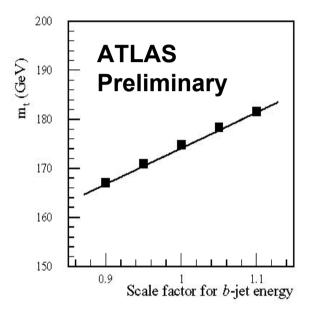






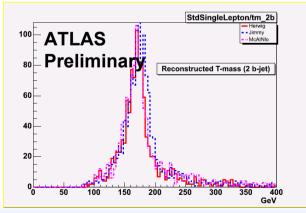
- Assume low luminosity and/or detector pessimistic scenarios
  - Partly or non-working b-tagging at startup
  - Dead regions in the LArEM
  - Jet energy scale
- Initially uncertainty on b-jet energy scale expected to be dominant:

b-jet scale uncertainty	δ
1%	0.7 GeV
5%	3.5 GeV
10%	7 GeV



Cf: 10% on q-jet scale  $\rightarrow$  3 GeV on M<sub>top</sub>

Important to understand UE (see earlier)
 → can have a large effect (as large as 5 GeV on m<sub>t</sub>)



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- **QCD** Multijet Background
- Not possible to realistically generate this background
  - Crucially depends on Atlas' capabilities to minimize mis-identification and increase  $e/\pi$  separation
- This background has to be obtained from data itself
  - E.g. method developed by CDF during run-1:

Use missing ET vs lepton isolation to define 4 regions:

A. Low lepton quality and small missing  $\mathbf{E}_{\mathrm{T}}$ 

Mostly non-W events (i.e. QCD background)

B. High lepton quality and small missing  $E_{T}$ 

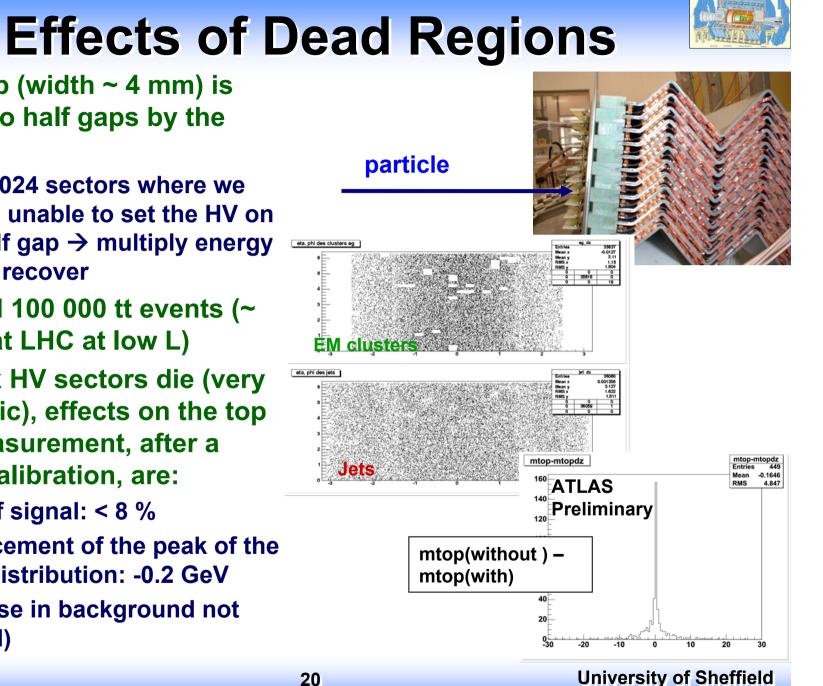
Observation of reduction in QCD background by isolation cut

C. Low lepton quality and high missing  ${\rm E}_{\rm T}$ 

W enriched sample with a fraction of QCD background

- **D. High lepton quality and high missing E**<sub>T</sub> W enriched sample
- The QCD reduction factor B/A can be applied to the "W enriched sample " (region C and D).
- The non-W candidate in D will therefore be (B/A)xC. Therefore, the fraction of non-W events in the region D will be:

(B.C)/(A.D)



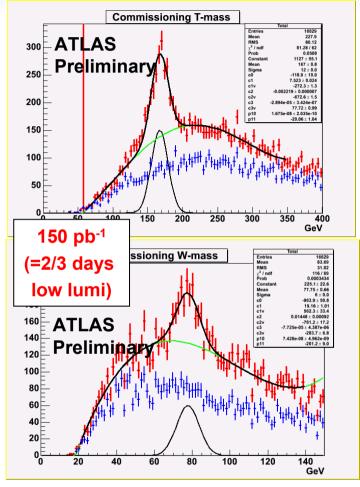
- Argon gap (width  $\sim$  4 mm) is split in two half gaps by the electrode
  - ~ 33 / 1024 sectors where we may be unable to set the HV on one half gap  $\rightarrow$  multiply energy by 2 to recover
- Simulated 100 000 tt events (~ 1.5 days at LHC at low L)
- If 33 weak HV sectors die (very pessimistic), effects on the top mass measurement, after a crude recalibration, are:
  - Loss of signal: < 8 %</p>
  - Displacement of the peak of the mass distribution: -0.2 GeV
  - (Increase in background not) studied)



### **Top Mass without B-tag**



- Most important background for top: W+4 jets
  - Leptonic decay of W, with 4 extra 'light' jets
- Selection:
  - Isolated lepton with P<sub>T</sub>>20 GeV
  - Exactly 4 jets ( $\Delta R=0.4$ ) with P<sub>T</sub>>40 GeV
- Reconstruction:
  - Select 3 jets with maximal resulting P<sub>T</sub>
- Try to identify W peak (also useful for JES calibration)
- Select highest  $p_T 2$  jet combination
  - W peak visible in signal
  - No peak in background
  - Better ideas possible?



		Prel	A5 iminary				
1	mean	σ(stat)		0 20 40 60 80 100 120 140 GeV			
	167.0	0.8	Health warning: Systematics not included / fast simulation used.				
	77.8	0.7	Cu	Currently under detailed study			

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150 pb<sup>-\*</sup>

Mtop

Mw

ATLAC



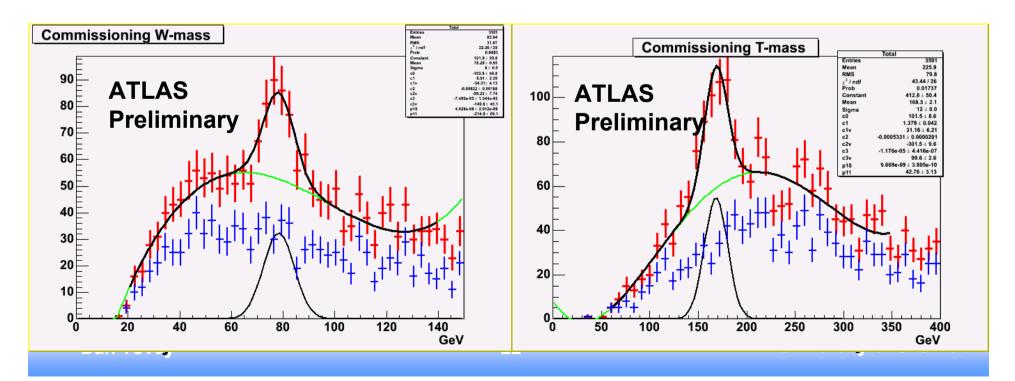
### Lower luminosity?



- Go down to 30 pb<sup>-1</sup>
  - Both W and t peaks already observable
  - See something!

Health warning: Systematics not included / fast simulation used. Currently under detailed study

		ATLAS Preliminary		
30 pb <sup>-1</sup>	mean	σ(stat)		
Мtop	170.0	3.2		
Mw	78.3	1.0		

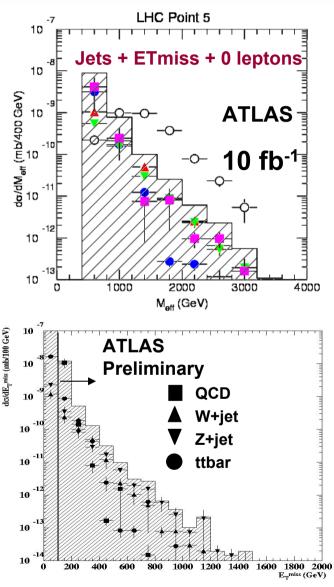




### SUSY



- Inclusive signature: jets + n leptons + E<sub>T</sub><sup>miss</sup>
- Main backgrounds:
  - Z + n jets
  - W + n jets
  - ttbar
  - QCD
- Greatest discrimination power from E<sub>T</sub><sup>miss</sup> (R-Parity conserving models)
- Generic approach to background estimation:
  - Select low E<sub>T</sub><sup>miss</sup> background calibration samples;
  - Extrapolate into high E<sub>T</sub><sup>miss</sup> signal region.
- Extrapolation is non-trivial.
  - Must find variables uncorrelated with E<sub>T</sub><sup>miss</sup>



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- Aim to use techniques developed at CDF/D0 + some new ones
- W/Z + n jets
  - Z → vv + n jets, W → Iv + n jets, W → τv + (n-1) jets (τ fakes jet)
  - Estimate from  $Z \rightarrow I^+I^- + n$  jets (e or  $\mu$ )
  - Tag leptonic Z and use to validate MC / estimate  $E_T^{miss}$  from  $p_T(Z) \& p_T(I)$
- QCD / fake E<sub>T</sub><sup>miss</sup> (from gaps in acceptance, dead/hot cells, nongaussian tails etc.)
  - Much harder : simulations require detailed understanding of detector performance (not easy with little data).
  - Strategy (learn from Tevatron):
    - 1) Initially choose channels which minimise contribution until well understood
    - 2) Reject events where fake  $E_T^{miss}$  likely: beam-gas and machine background, bad primary vertex, hot cells, CR muons,  $E_T^{miss}$  vector pointing in (opposite) direction of (to) jets (jet fluctuations), jets pointing at regions of poor response, large Missing  $E_T$  Significance
    - 3) Choose hard cuts which minimise contribution to background.
    - 4) Estimate background using data and/or calibrated fast MC: need to estimate jet resolution functions using e.g. E<sub>T</sub><sup>miss</sup> projection



# **Top Background**



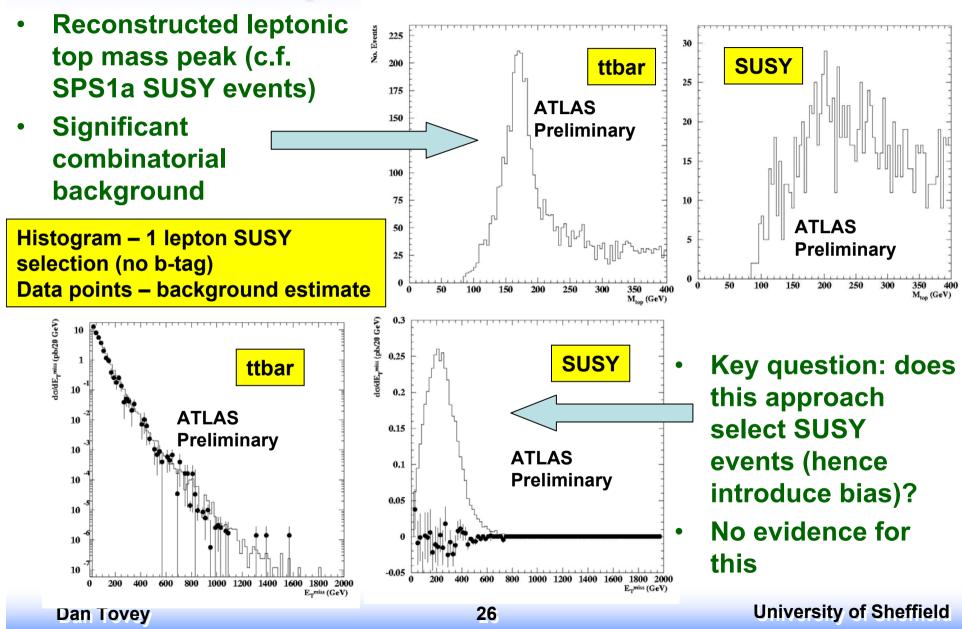
- Estimation using simulation possible (normalised to data ttbar selection) cross-check with data ?
- Standard (TDR) semileptonic top cuts look rather like SUSY cuts with looser E<sub>T</sub><sup>miss</sup> requirement!

Process	_p <sub>1</sub> <sup>1</sup> > 20GeV ε <sub>1</sub> <sup>misa</sup> > 20GeV	As before, plus N <sub>jet</sub> ≥ 4	As before, plus N <sub>b-jet</sub> ≥2	Events per 10 fb <sup>-1</sup>	
ft signal	64.7	21.2	5.0	126 000	Physics TD
W+ jets	47.9	0.1	0.002	1658	
Z+ jets	15.0	0.05	0.002	232	
ww	53.6	0.5	0.006	10	
WZ	53.8	0.5	0.02	8	
Z	2.8	0.04	0.008	14	
Total background				1922	
S/B				65	

- If harden E<sub>T</sub><sup>miss</sup> cuts top sample contaminated with SUSY signal (bias) ...
- Possible approach?
  - Select semi-leptonic candidates (standard cuts what btag available?);
  - Fully reconstruct top from E<sub>T</sub><sup>miss</sup> & W mass constraint;
  - Reduce combinatorics with highest p<sub>T</sub> W candidate
  - Reject (SUSY) background with mass cut & mtop sideband subtraction;
  - Use to validate top production in MC / estimate remaining E<sub>T</sub><sup>miss</sup> background.

# **Top Reconstruction**







### Conclusions



- Lots of work currently being carried out preparing for first data.
- Detailed studies of calibration & alignment with cosmics and beam halo / beam-gas
- Preliminary studies of commissioning using collision data completed – more on-going.
- Physics Working Groups studying techniques needed to estimate/reduce backgrounds to specific channels 
   also requires development of new tools.
- ATLAS will be ready to make optimum use of first physics data when it arrives.





# **Backup Slides**

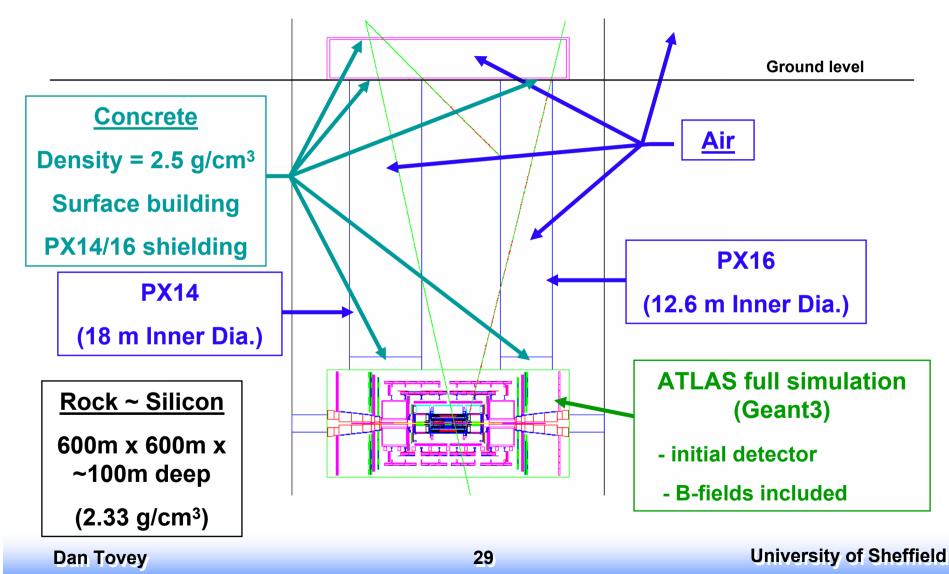
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### **Phase 3: Cosmic** μ



• Full simulation of cosmic ray muons in ATLAS developed (G3)





### **Expected Cosmics Rates**



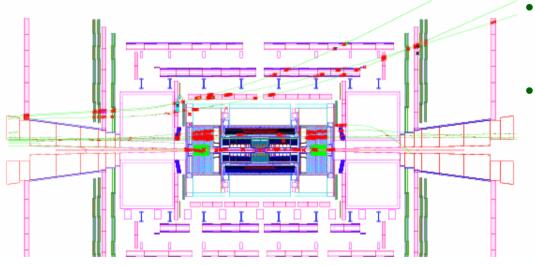
	Condition	Rate	(Hz)
ATLAS Preliminary		E <sub>surface</sub> > 10 GeV ("PDG" approximation)	E <sub>surface</sub> >10 GeV ("ALE" generator)
ATLAS UX15		5900	4900
Any G3 digit		2800	2300
Through	RPC <sub>Y&gt;0</sub> x RPC <sub>Y&lt;0</sub> x ID <sub>DIGI</sub>	28	24
going	RPC <sub>Y&gt;0</sub> x RPC <sub>Y&lt;0</sub> x PIX <sub>DIGI</sub>	0.6	0.4
Pass by	Z <sub>DIGI</sub>   < 300,  R <sub>DIGI</sub>   < 60 cm	12.2	10.2
≈ origin	Z <sub>DIGI</sub>   < 100,  R <sub>DIGI</sub>   < 30 cm	2.3	1.9
	Z <sub>DIGI</sub>   < 60,  R <sub>DIGI</sub>   < 20cm	0.6	0.5
	E <sub>T</sub> <sup>CELL</sup> > 5 GeV	0.1	0.1
EM Cal	E <sub>T</sub> <sup>CLUSTER</sup> > 5 GeV	0.2	0.2
	E <sub>T</sub> <sup>TOTAL</sup> > 5 GeV	0.4	0.4
Tile Cal	E <sub>TOTAL</sub> > 20 GeV	1.4	1.2
HEC	E <sub>TOTAL</sub> > 20 GeV	0.1	0.1
FCAL	E <sub>TOTAL</sub> > 20 GeV	0.02	0.02
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### **Beam halo**

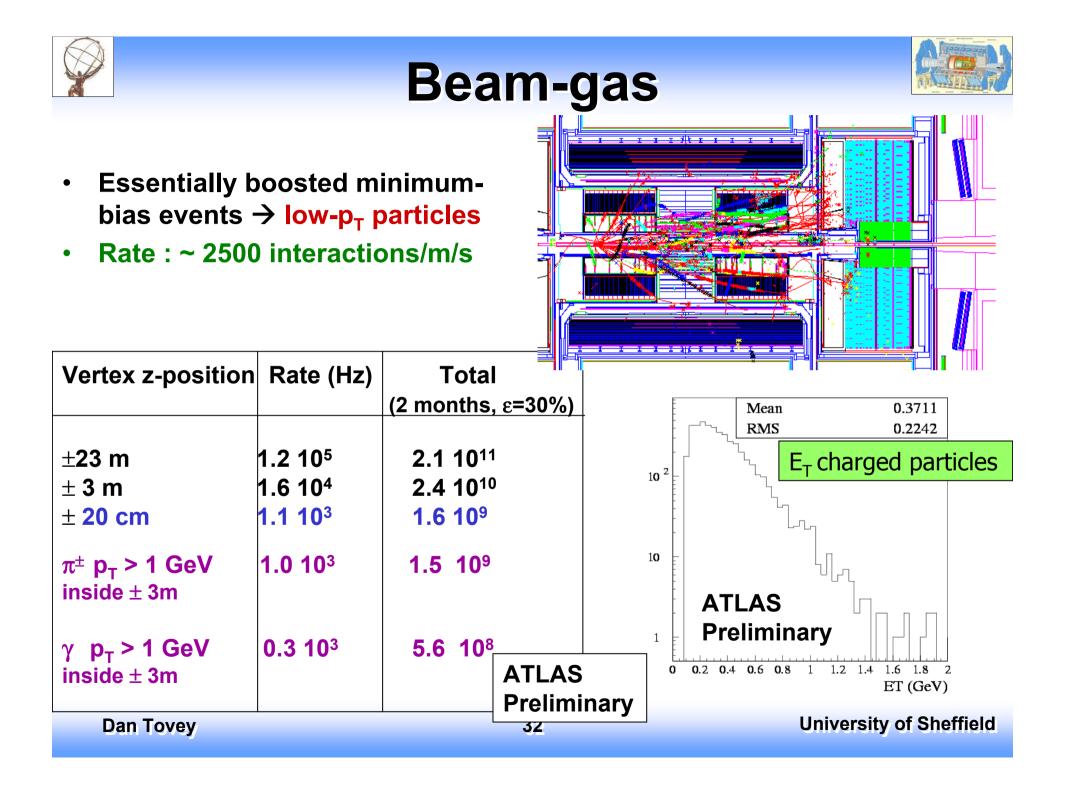


- Rates for initial period scaled from highluminosity rates by assuming
- 3 x 10<sup>10</sup> p per bunch and 43 bunches → ~ 200 times lower current (but assuming same vacuum, etc.)
- Total rates assume two months single-beam w/ 30% data taking eff.



Detector	Rate (B-field off )	Total (B-field off)	Rate (B-field on)	Total (B-field on)
MDT barrel	15 Hz	2.5 10 <sup>7</sup>	72 Hz	1.5 10 <sup>8</sup>
MDT end-cap	145 Hz	2.5 10 <sup>8</sup>	135 Hz	2.5 10 <sup>8</sup>
Pixel/SCT	1.8/17 Hz	3 10 <sup>6 /</sup> 3 10 <sup>7</sup>	2/19 Hz	3 10 <sup>6 /</sup> 3 10 <sup>7</sup>
EM E > 5 GeV	2 Hz	3.5 10 <sup>6</sup>	1 Hz	1.7 10 <sup>6</sup>
Tile/HEC E > 20 GeV	1.7/1.2 Hz	2.9/2.1 10 <sup>6</sup>	1.6/0.9 Hz LAS Prelin	2.8/1.6 10 <sup>6</sup>

- Simple definition of "useful tracks": 2-3 segments in MDT+ 3-4 disks in ID end-cap
- More recently: results from simulation of machine conditions in the commissioning period (including more realistic vacuum estimates, etc.) give rates ~ 7 lower

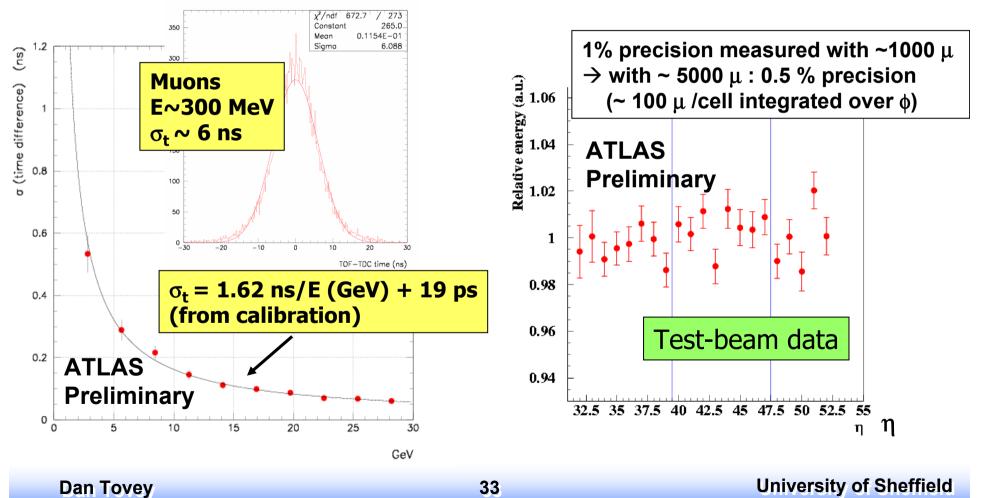




### **Cosmics in ECAL**



- With ~ 100 muons/cell in middle compartment:
- check calorimeter timing to < 1 ns</li>
- check calorimeter position in  $\eta$  /  $\phi$  wrt other sub-detectors to < 1 mm
- check response uniformity vs  $\eta$ :  $\approx$  0.5% precision could be achieved







### Cosmics in µ System

- Cosmic rate high enough for polar angles up to  $\theta$ =75°: ~1Hz/strad for muons going through the ID (almost projective) and  $p_{\mu}$ >10 GeV
  - Study of all barrel sectors (probably except sectors 1-9 with vertical chambers) and part of the forward chambers
- First test of the full reconstruction (field off/reduced/full field)
- Map dead channels, chase/replace faulty FE cards
- Tube efficiency, R-t relation (autocalibration):
  - 1000 (no field)-10000 (with field)  $\mu/tube \ \Rightarrow$  ~10-100 days
- Check/calibration of the (barrel only?) alignment system with straight tracks (<30µm level): 2000µ/chamber ~10 hours</li>
- Alignment  $\mu$  barrel/  $\mu$  End cap,  $\mu$  spectro/ID

# **B-Tagging Efficiency**



- $\epsilon_{tag}$  = probability to tag at least one jet in a top event
  - $-\varepsilon_{tag} = \varepsilon_{b-tag} + \varepsilon_{non-b} (\varepsilon_{b-tag} \cdot \varepsilon_{non-b})$
  - $\quad \varepsilon_{\text{non-b}} = \varepsilon_{\text{c-tag}} + \varepsilon_{\text{nonhf}}$
- $\varepsilon_{b-taq}$  is the sum of these possibilities:
  - Probability to tag 1 b-jet in the event, when 1 is found in the detector
  - Probability to tag 1 b-jet when 1 is found in the detector
  - Probability to tag 2 b-jets when 2 are found in the detector
- First simple evaluation (counting method):
  - Select a very pure ttbar sample with tight kinematical cuts
  - Count the number of events with <u>at least</u> 1 tagged b-jet
  - Divide this number by the number of pre-tag candidate events
- $\epsilon$ 's are measured in MC. Account for difference in tagging between MC and data with Scale Factor:

$$\varepsilon_{b-tag}^{event} = F_{1b} \cdot \varepsilon_{btag} \cdot SF + F_{2b} \cdot \varepsilon_{btag}^2 \cdot SF^2 + 2 \cdot F_{2b} \cdot \varepsilon_{btag} \cdot SF \cdot (1 - \varepsilon_{btag})$$

Probability to tag oneProbability to tag twoPB-jet when one is foundB-jets when two are foundB

Probability to tag one B-jet when two are found

- F<sub>1b</sub> = fraction of events with 1 taggable jets
- $F_{2b}$  = fraction of events with 2 taggable jets