

Lecture Plan

- I. History of Antimatter
- II. Antimatter and the Universe
- III. Production and trapping of antiparticles
- IV. Precision tests of particle-antiparticle symmetry
- V. AD Physics and Antihydrogen**
- VI. Antimatter technologies**



today

ANTIHYDROGEN

ATHENA (AD-1)

ATRAP (AD-2)

Phase 1 (2000-2004)

Production of slow antihydrogen (completed)

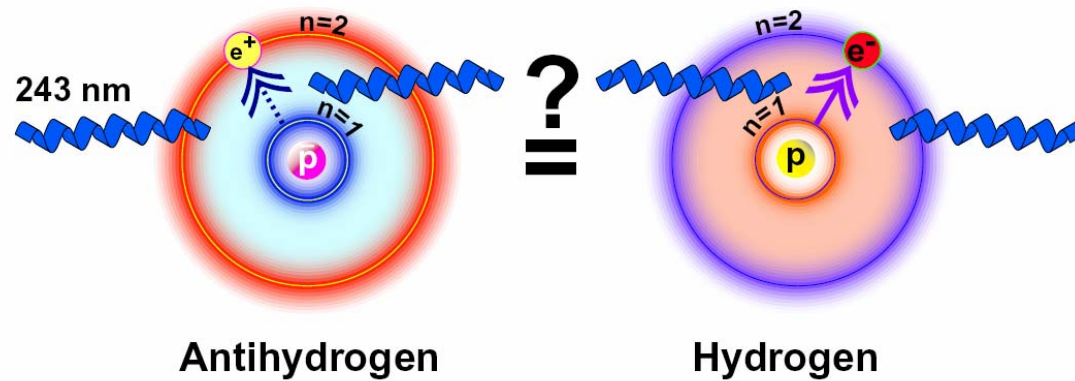
Phase 2 (2006-)

Trapping and cooling of antihydrogen

Phase 3 (?)

Precision experiments

Phase 3 - High precision comparison of hydrogen and antihydrogen



"Shelving" Scheme with single (trapped) atom:

Strong Lyman- α is excited and fluoresces

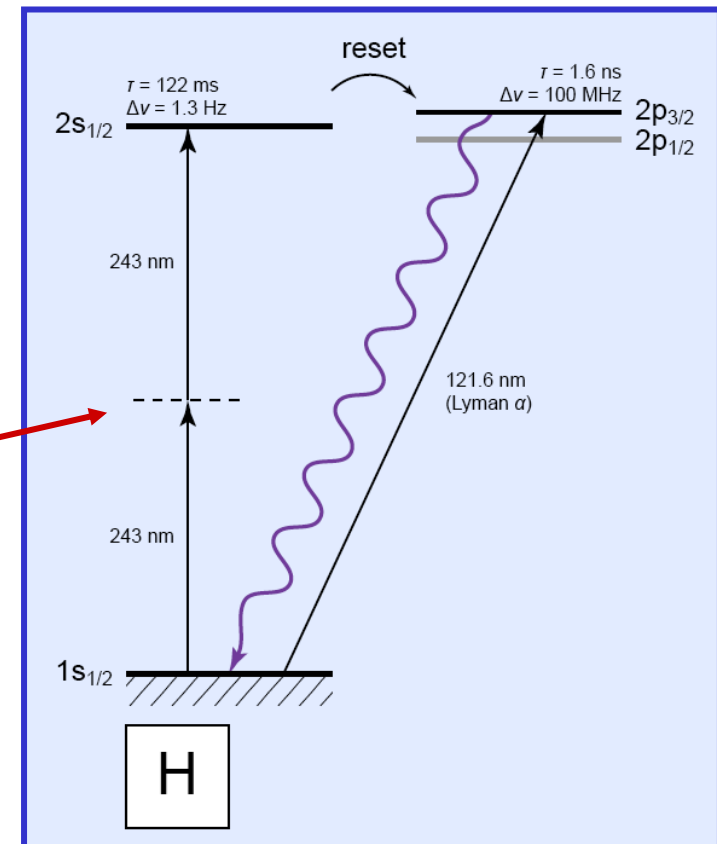
2s state is populated by 2-photon excitation

'Shelving' in (metastable) 2s suppresses fluorescence

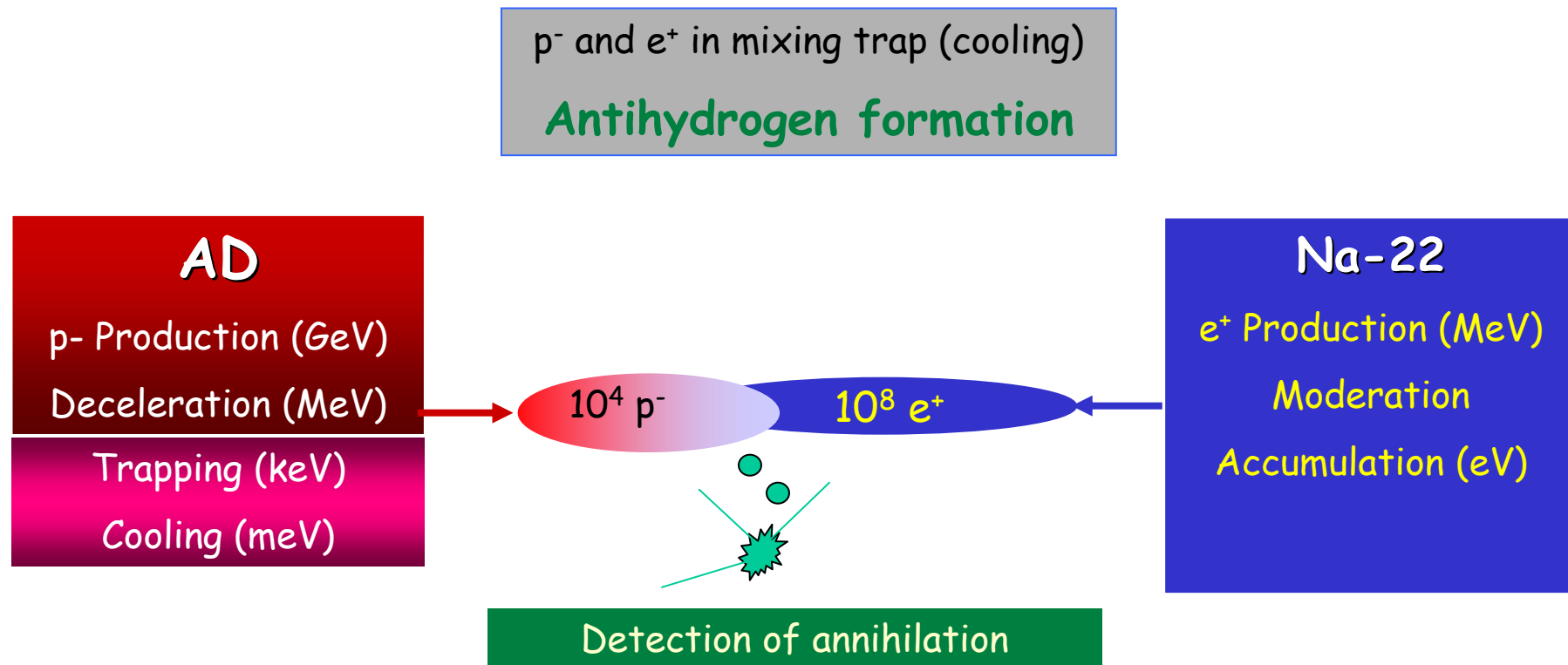
2s state is 'reset' with microwave transition into 2p

Natural line width $4 \cdot 10^{-16}$

[J. Walz et al., Hyp. Int. 127 (2000) 167]



How to make antihydrogen (Phase 1)



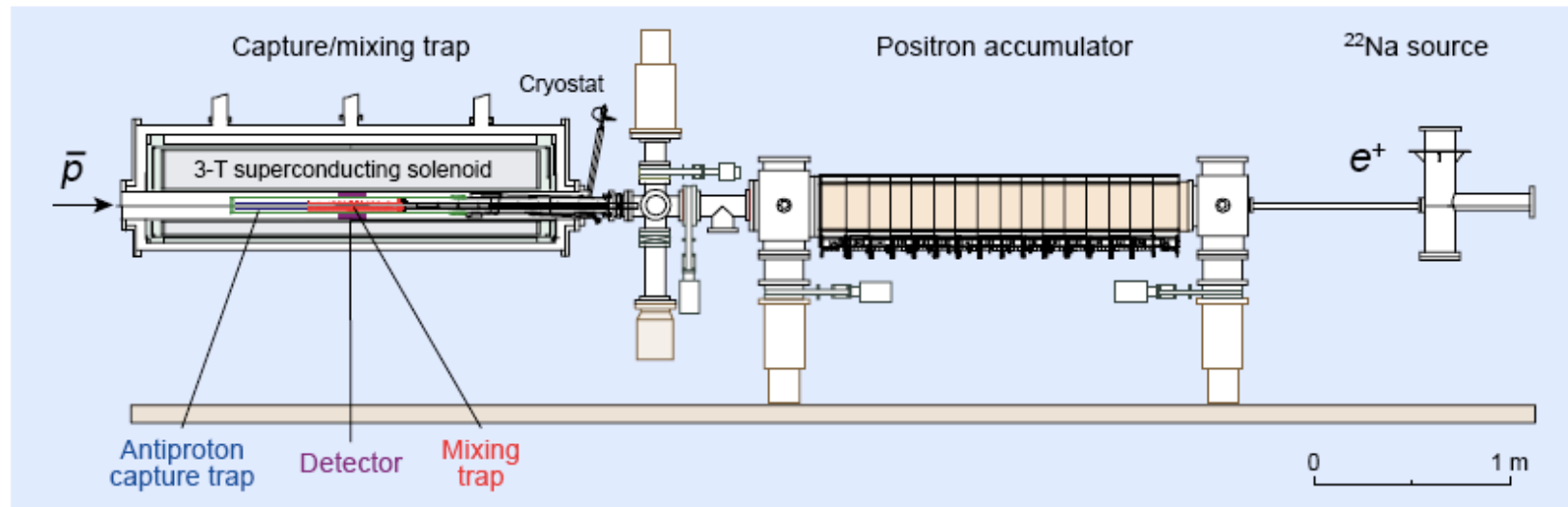
ATHENA - Schematic view

Antiproton capture trap

Deceleration and capture of antiprotons
Penning trap in 3-T field at 15 K
Cooling and accumulation in e^- plasma

^{22}Na source

Positron production via $^{22}\text{Na}(\beta^+)^{22}\text{Ne}$ at 5.5 K
Positron accumulator
Penning trap in 0.14-T field at 300 K



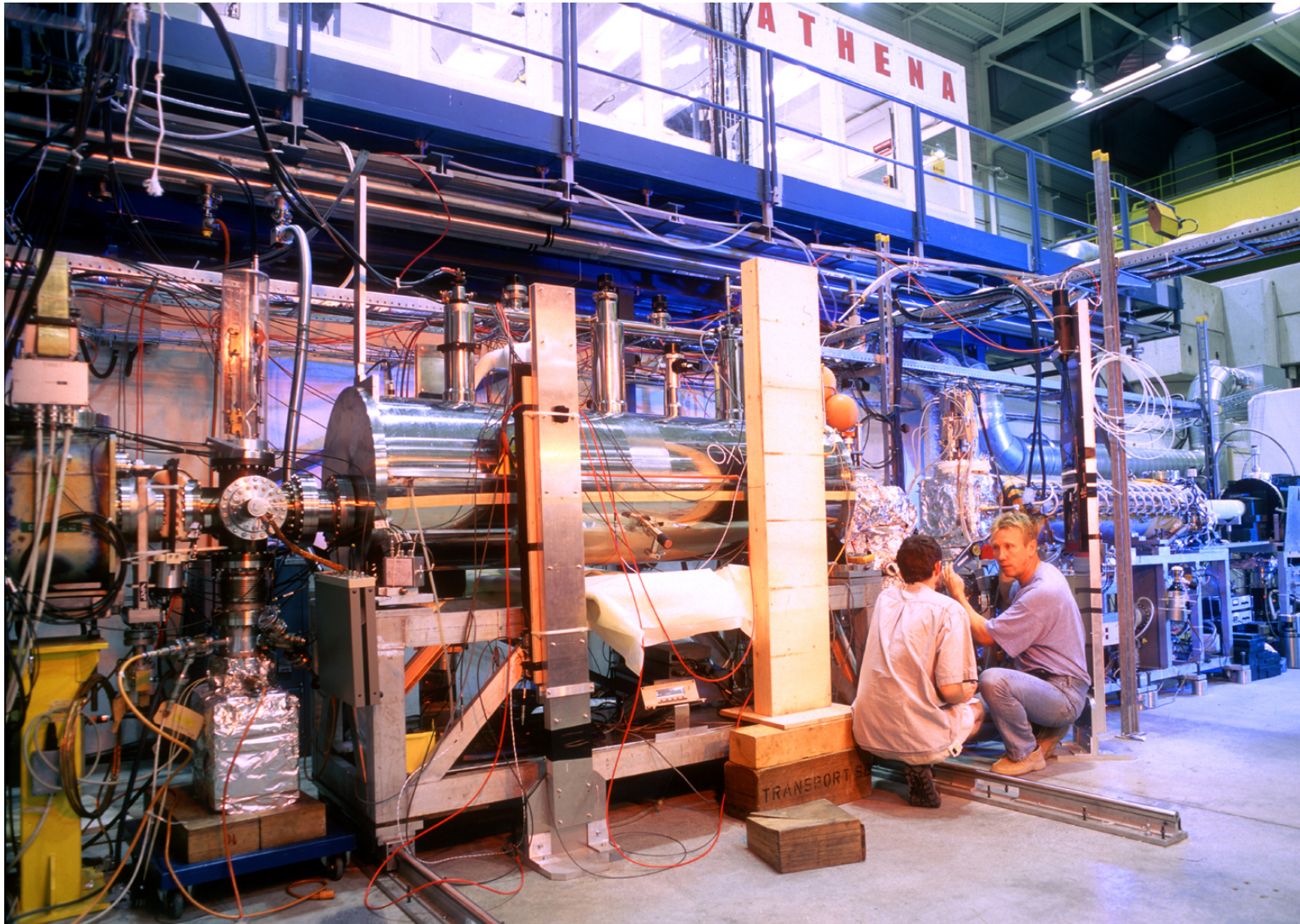
Mixing trap

Antihydrogen production
Nested Penning trap in 3-T field at 15 K

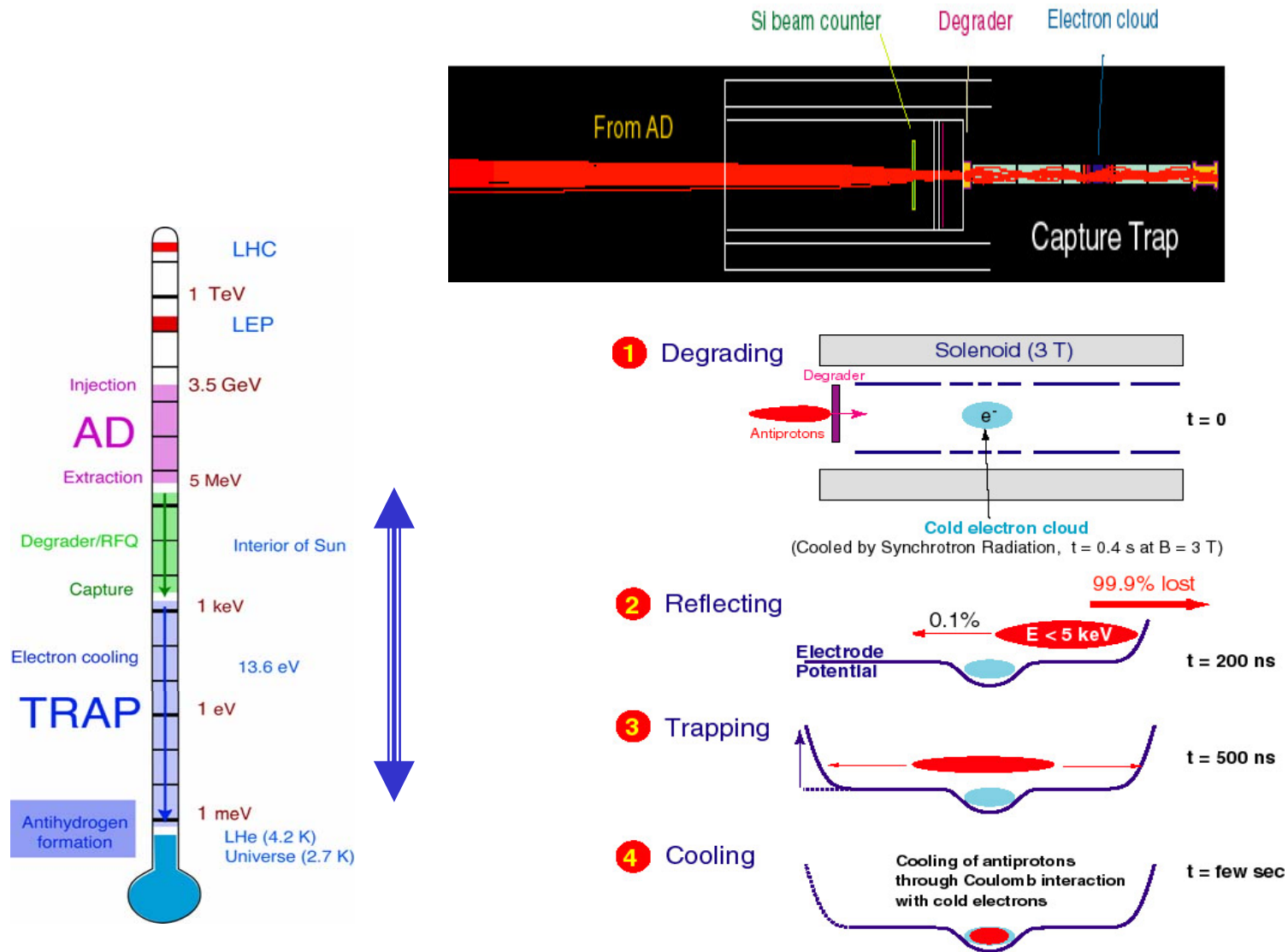
Detector

[M. Amoretti *et al.*,
NIM A 518 (2004) 679]

ATHENA (AD-1)

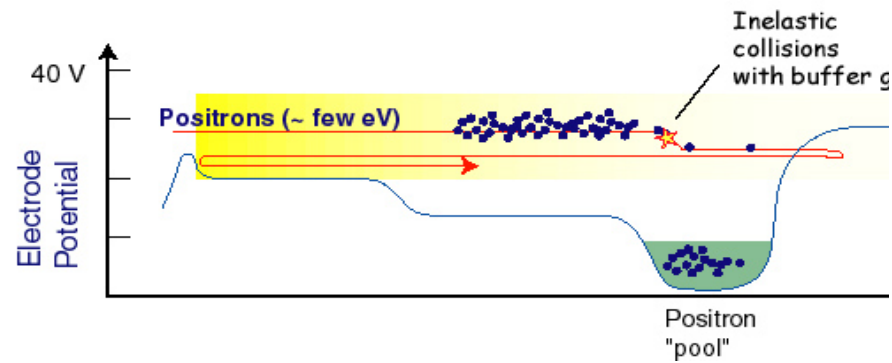
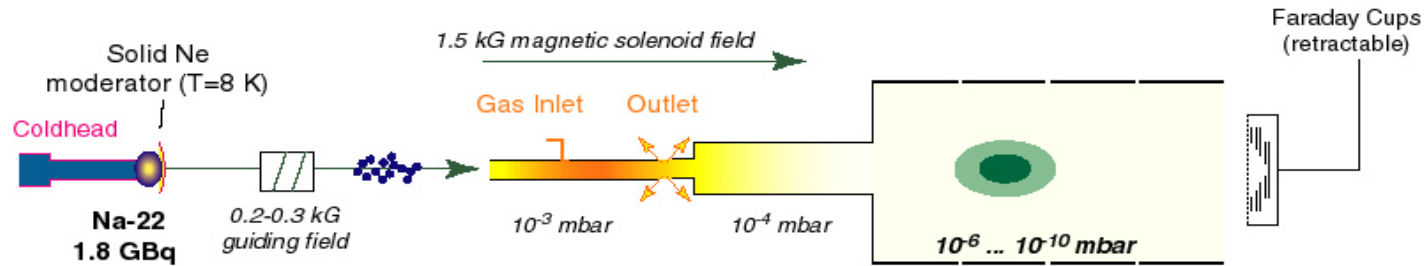


Antiprotons (capture and cooling)

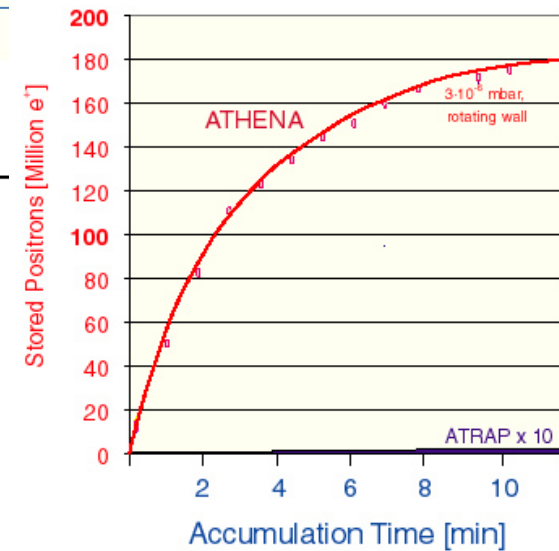


Positron Accumulation

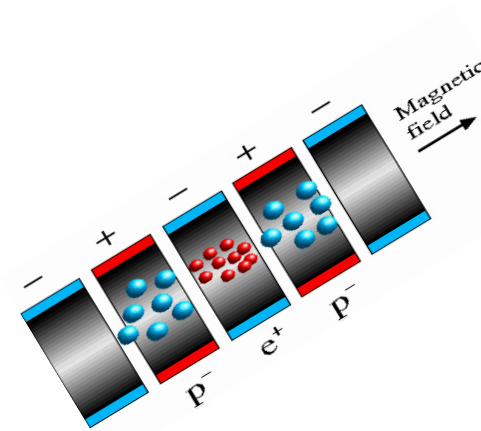
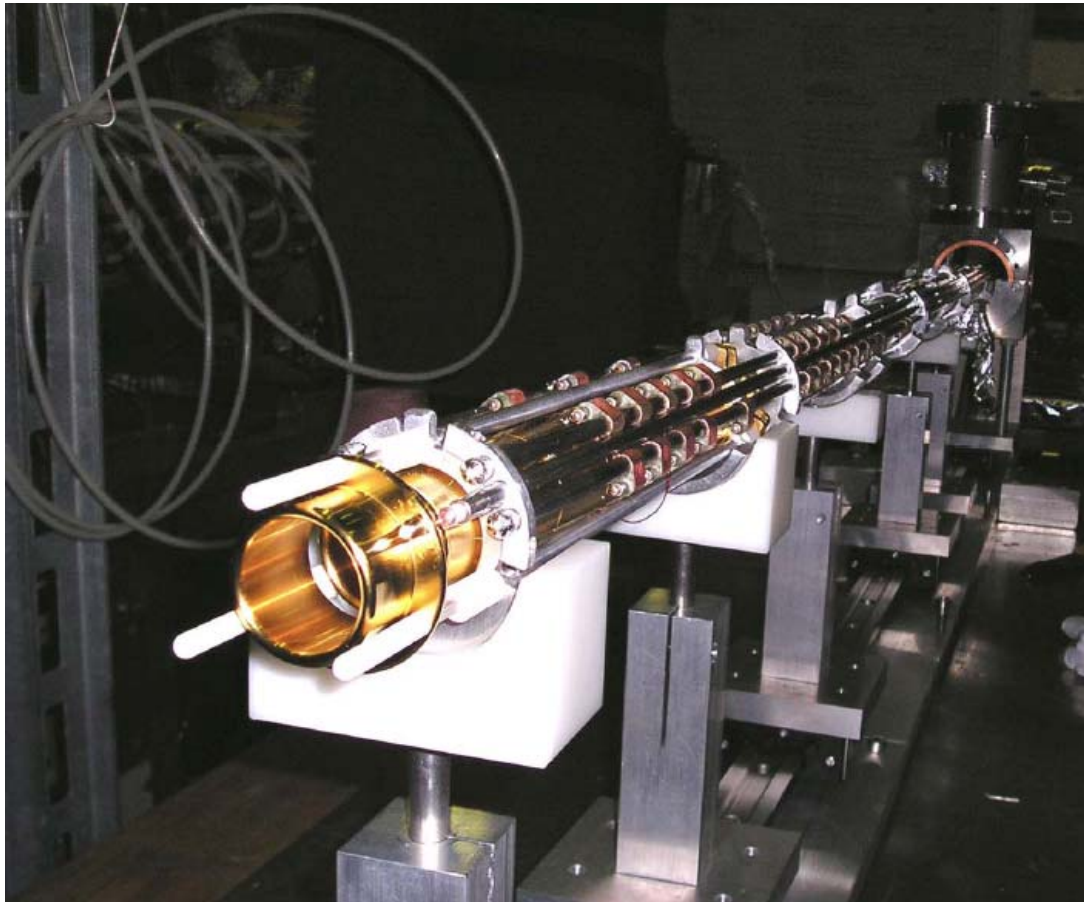
ATHENA - Positron Accumulation Scheme



Accumulated positrons vs time

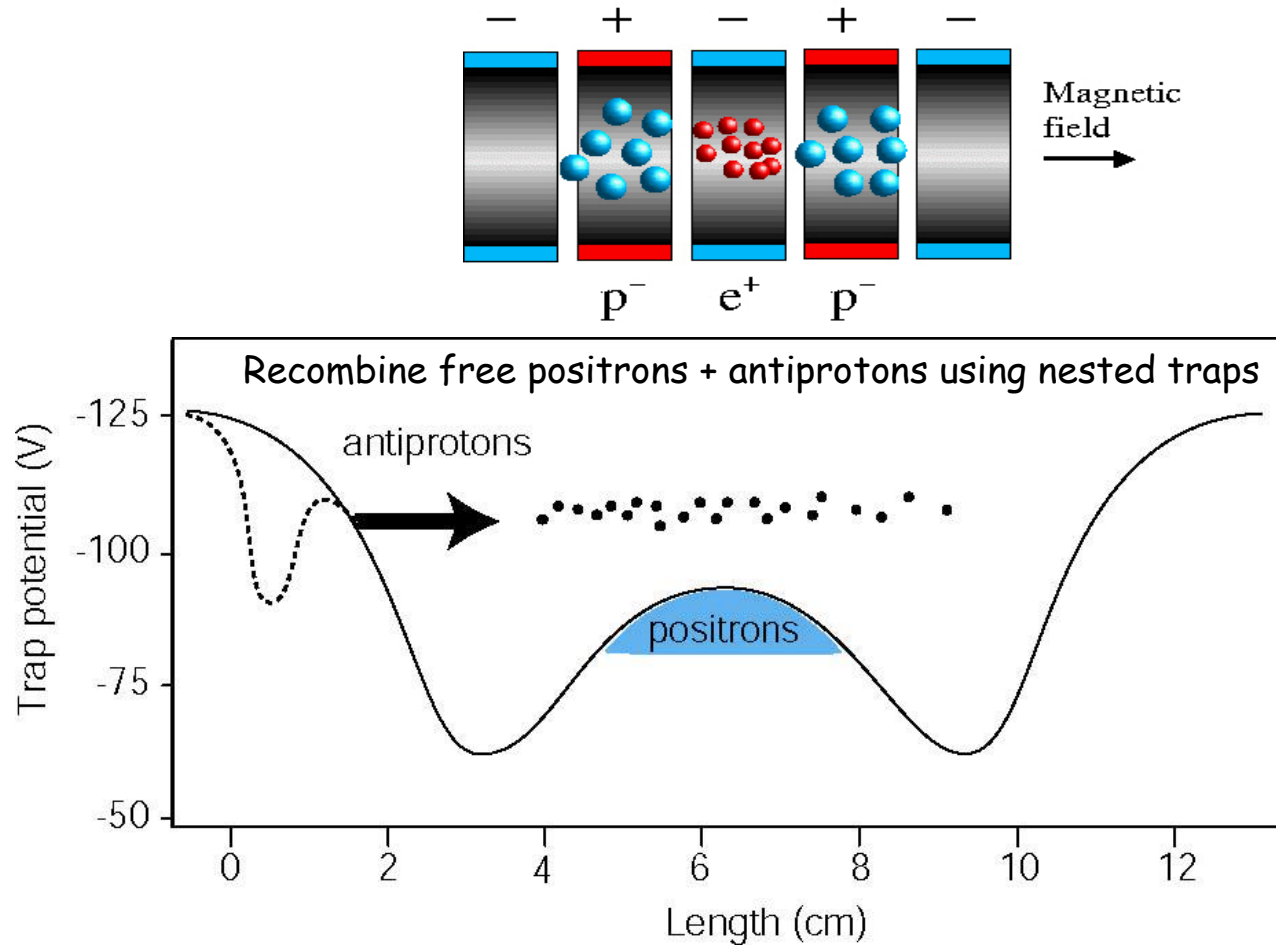


Electrical Landscapes



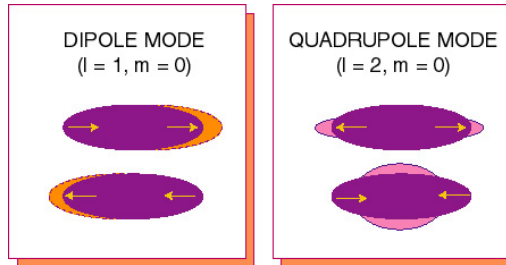
ATHENA multi-electrode trap

Antiproton-Positron Recombination Scheme

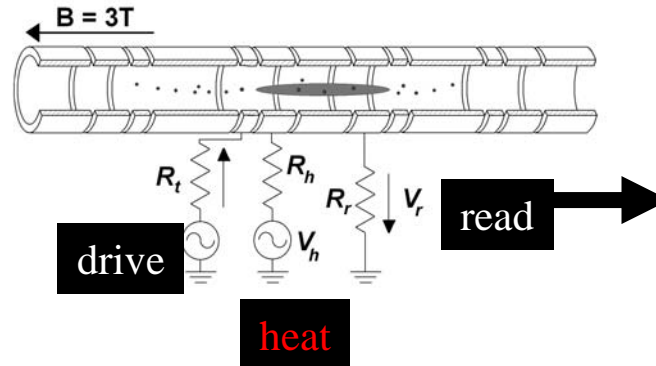


*D.S. Hall, G. Gabrielse, Phys. Rev. Lett. 77, 1962 (1996)

Cold vs hot - Positron plasma 'heating'



'Shake' positron plasma with RF fields
Heating leads to measurable shift of 'quadrupole frequency'



Recombination of antiprotons and positrons
is suppressed for hot positron plasma

Antihydrogen- Detection

Charged particles

2 layers of Si microstrip detectors

511 keV gammas

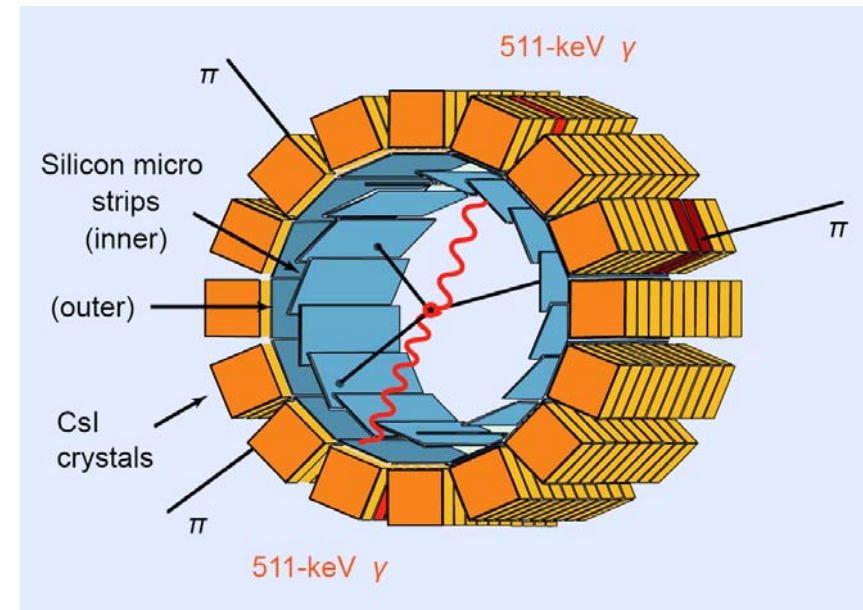
192 CsI crystals

Inner radius 4 cm, thickness \sim 3 cm

70% solid angle coverage

Operates at 3 Tesla, 140 Kelvin

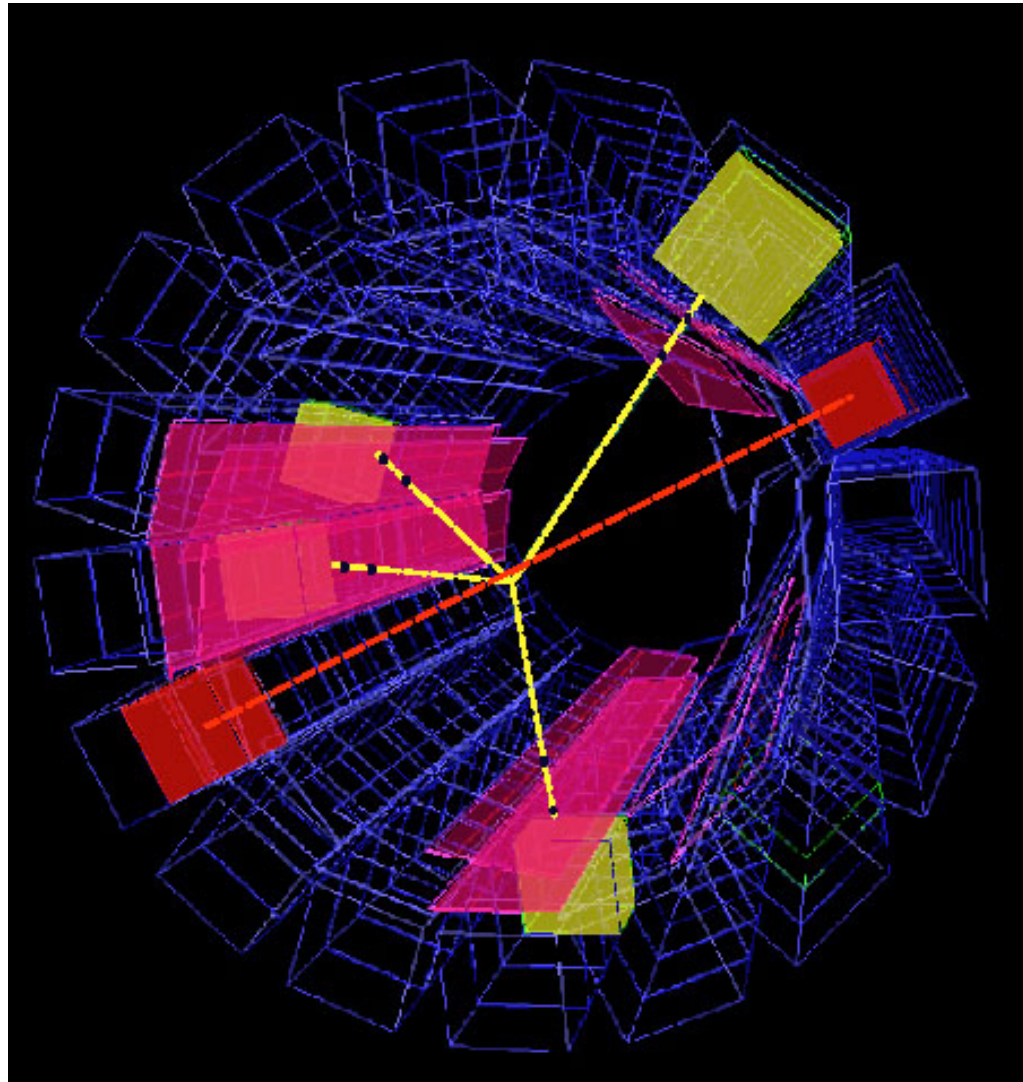
(C. Regenfus et al., NIM **A501**, 65 (2003))



Event analysis:

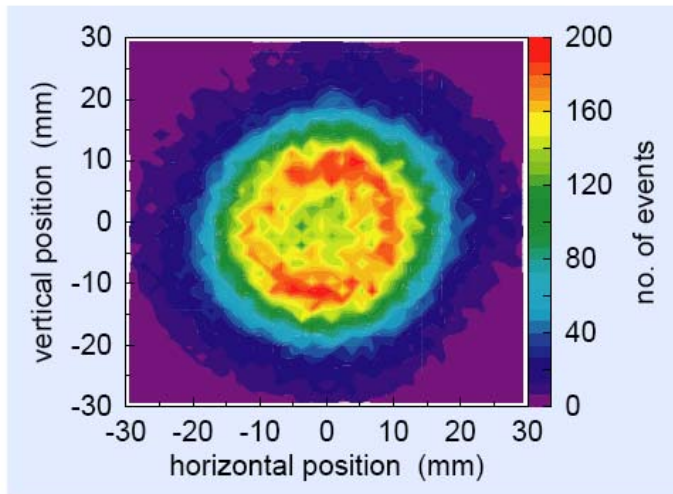
1. Reconstruct vertex from tracks of charged particles
2. Identify pairs of 511 keV γ -rays in time coincidence
3. Measure opening angle between the two γ -rays

ATHENA: Event Display

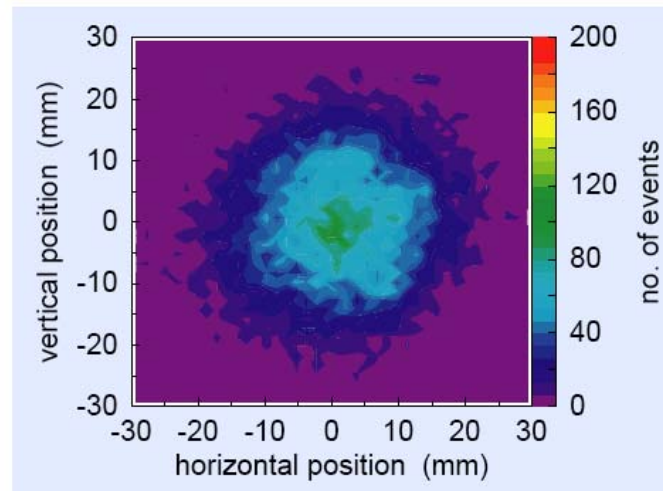


ATHENA: First observation of cold antihydrogen

1) Distribution of annihilation points



Cold positrons



Hot positrons
(RF excitation
of axial e^+
plasma modes)

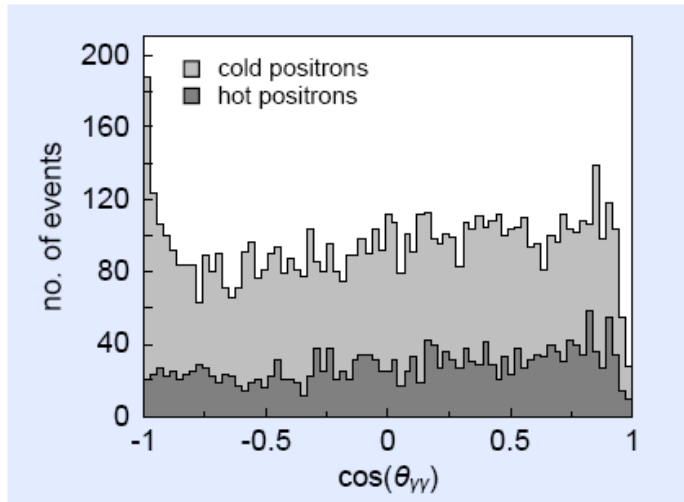
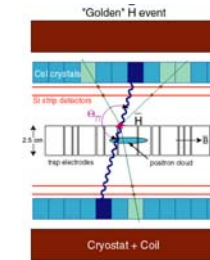
[M. Amoretti *et al.*, Nature **419** (2002) 456]

Neutral antihydrogen annihilates on trap walls (not trapped)

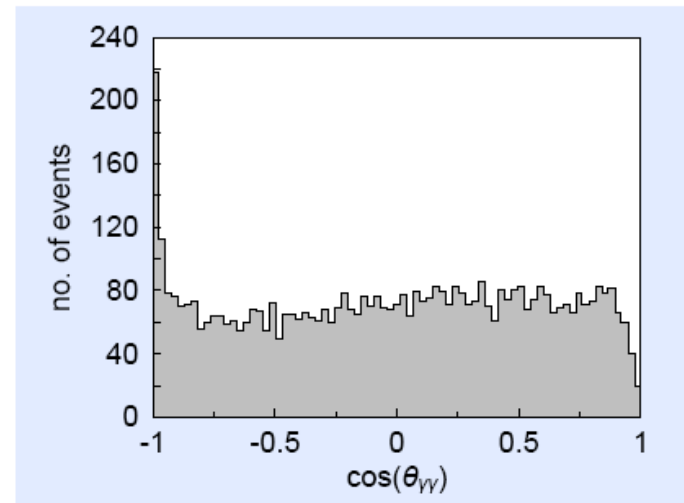
Heating of positrons suppresses antihydrogen formation

ATHENA - Antihydrogen (2)

2) Opening Angle Distribution



Data



Monte Carlo

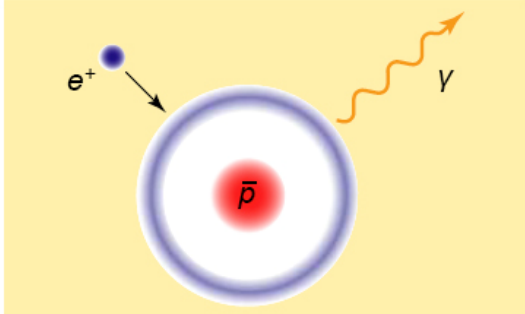
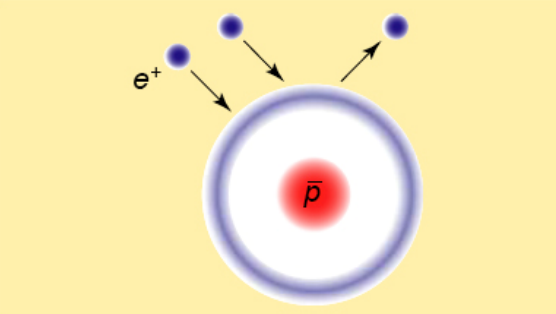
[M. Amoretti *et al.*, Nature 419 (2002) 456]

Peak from back-to-back 511 keV photon pairs

Disappears when positrons are 'hot'

Correcting for detection efficiency: $> 100,000$ anti-atoms

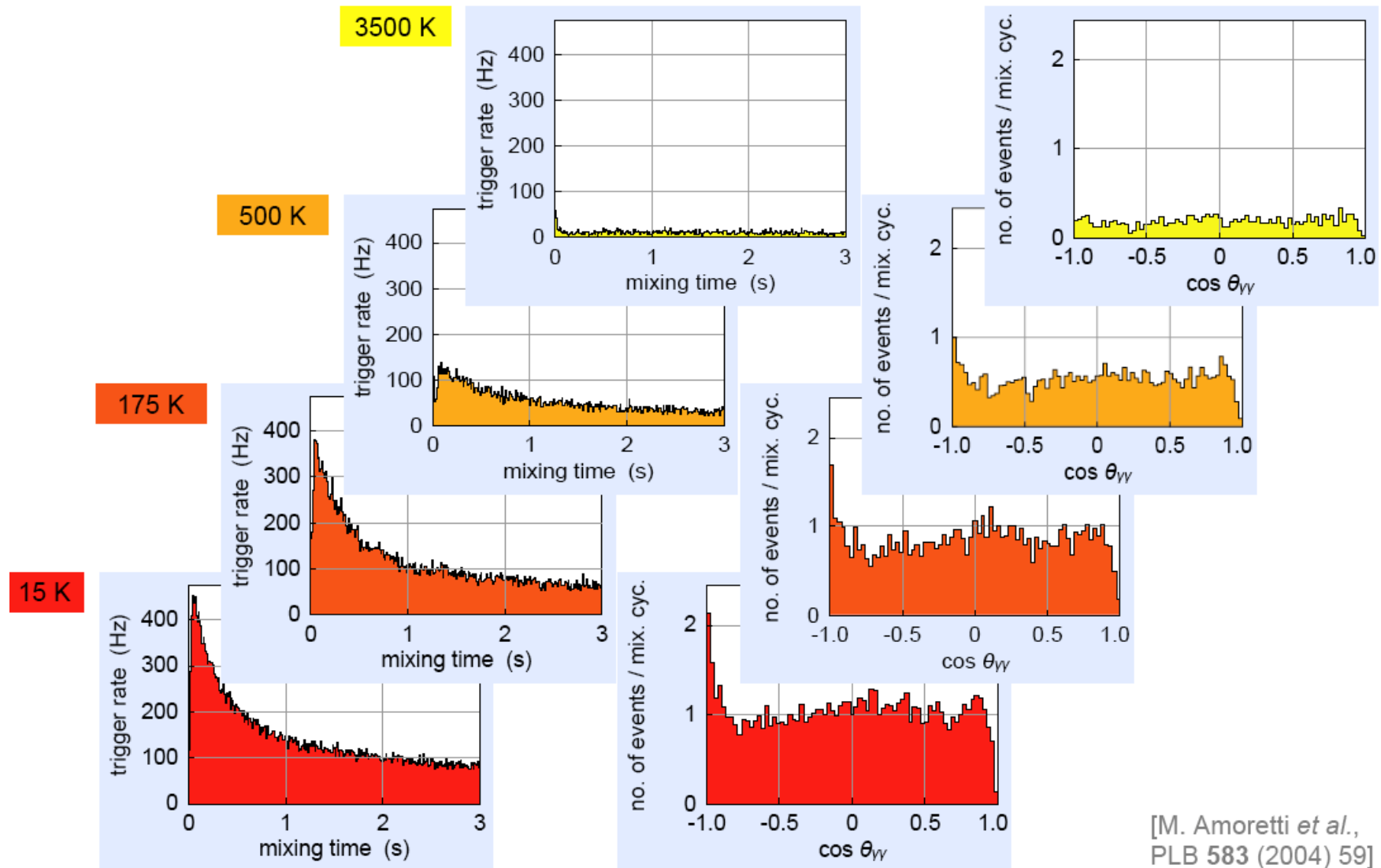
Theoretical models of recombination

	Radiative Recombination	Three-Body Recombination
Principle		
Temperature depend.	$\propto T^{-2/3}$	$\propto T^{-9/2}$
e^+ density dependence	$\propto n_e$	$\propto n_e^2$
Cross-section at 1 K	10^{-16} cm^2	10^{-7} cm^2
Final internal states	$n < 10$	$n \gg 10$

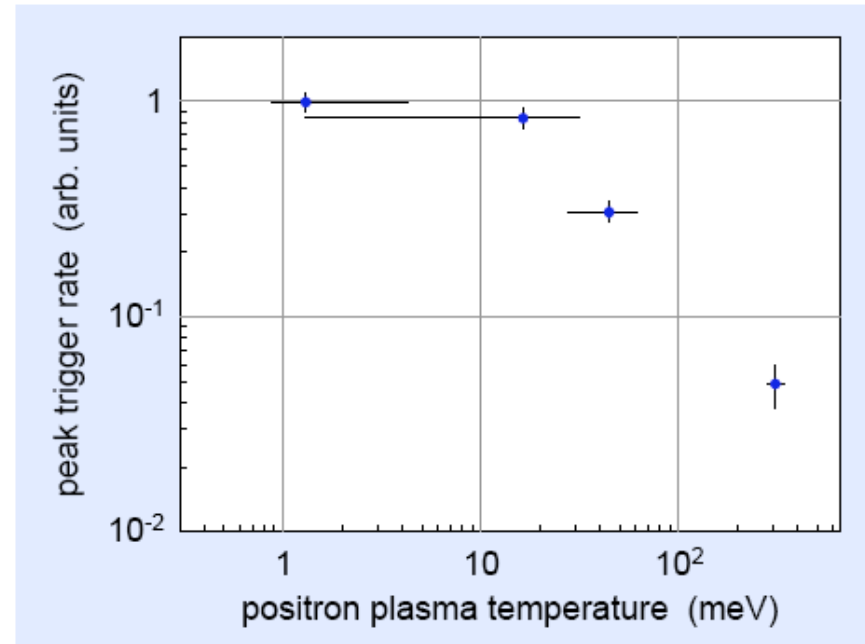
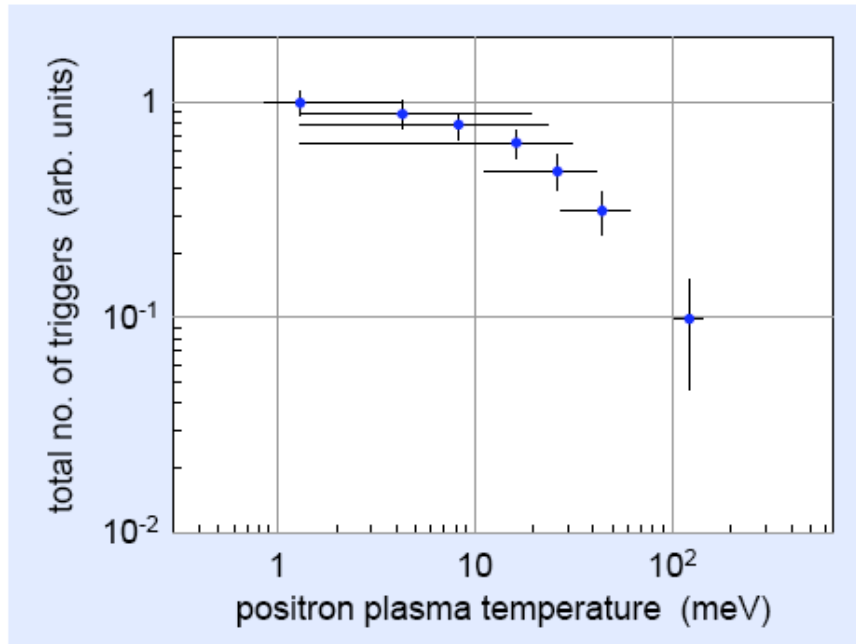
[J. Stevefelt *et al.*, PRA 12 (1975) 1246]

[M. E. Glinsky *et al.*, Phys. Fluids B 3 (1991) 1279]

Dependence of production rate on positron temperature



Recombination rate vs Positron Temperature



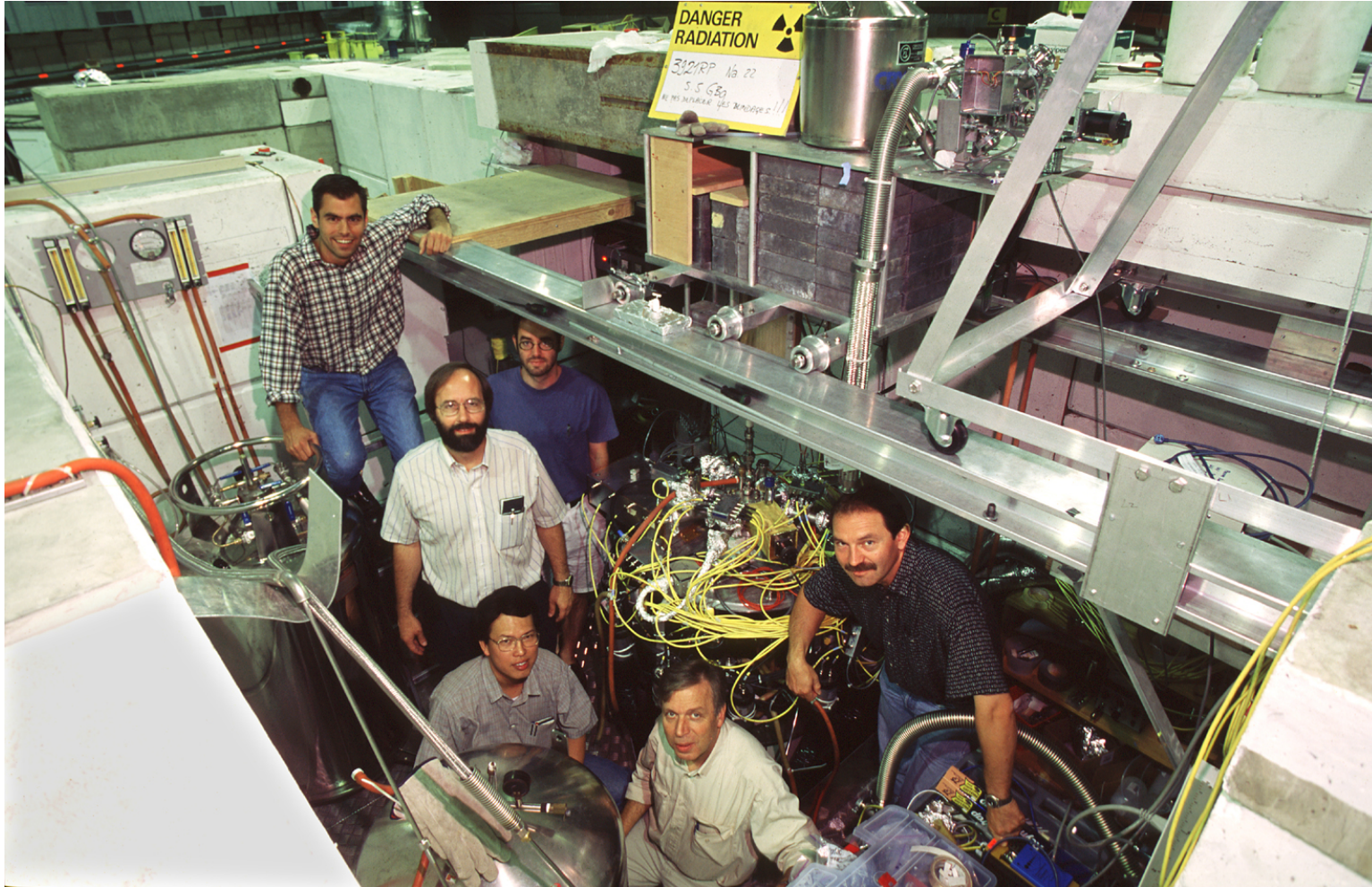
Production decreases with increasing temperature

No strong increase at very low temperature

Simplistic power law fit to data yields $\sim T^{-0.7 \pm 0.2}$

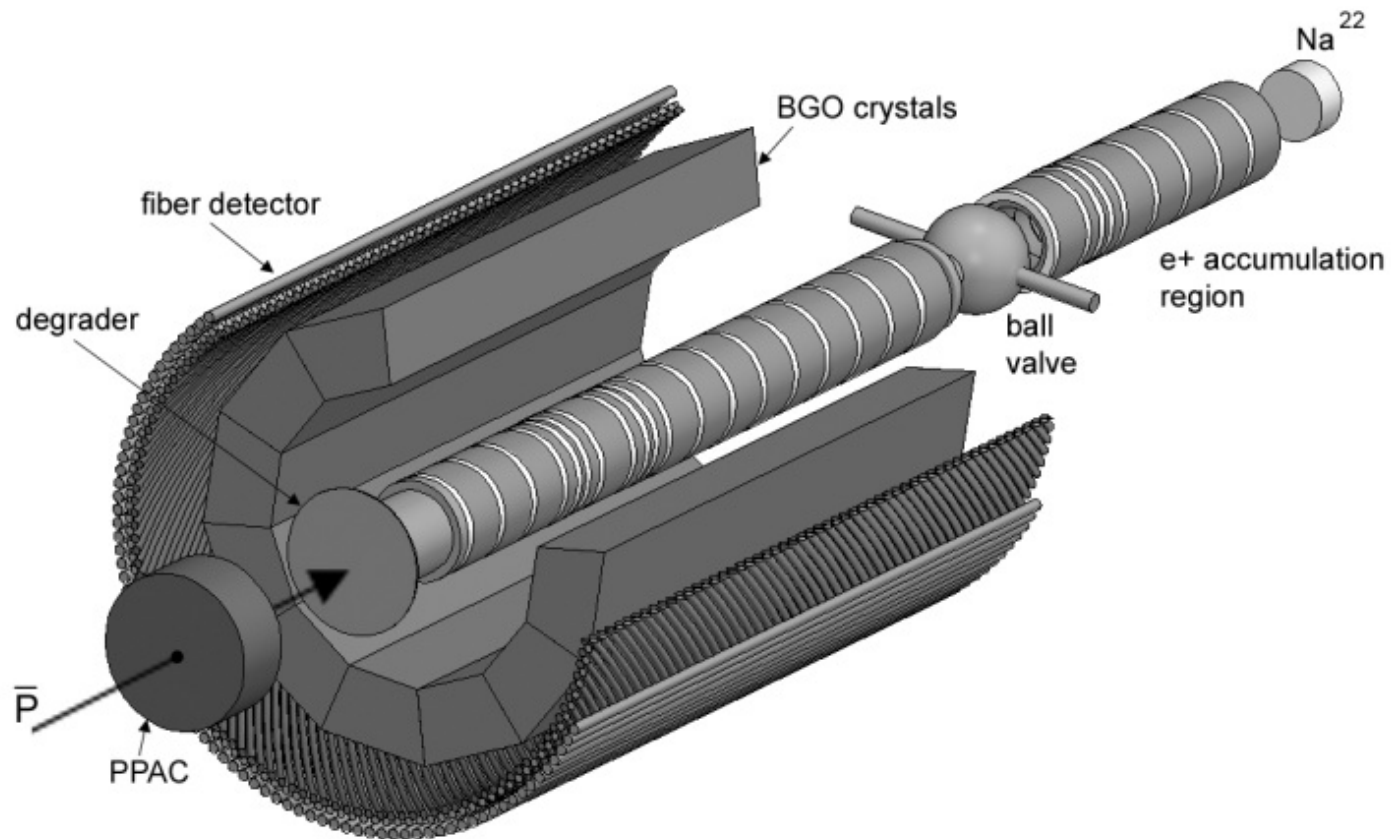
Note: corrections for plasma dynamics, magnetic field

ATRAP (AD-2)

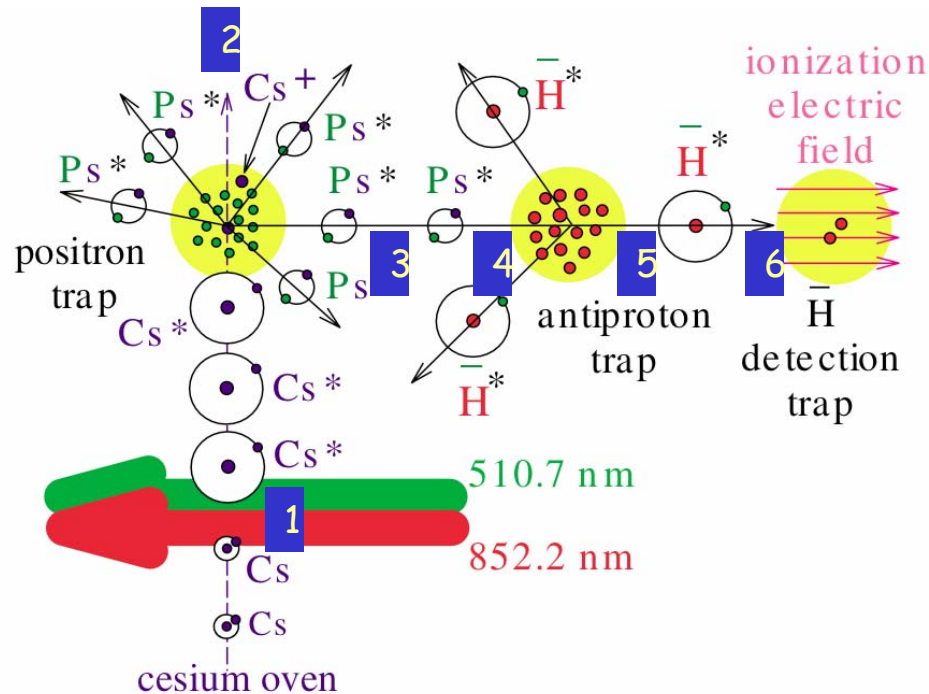


ATRAP (AD-2) - Overview

Nested Penning Trap



ATRAP uses positronium to make antihydrogen



- 1 Cs atoms are excited to $n \sim 50$
- 2 Cs* atoms collide with positron cloud
- 3 Formation of excited positronium
- 4 Ps* collides with cold antiprotons
- 5 Antihydrogen formation ($n \sim 45$)
- 6 Antihydrogen detection by ionisation

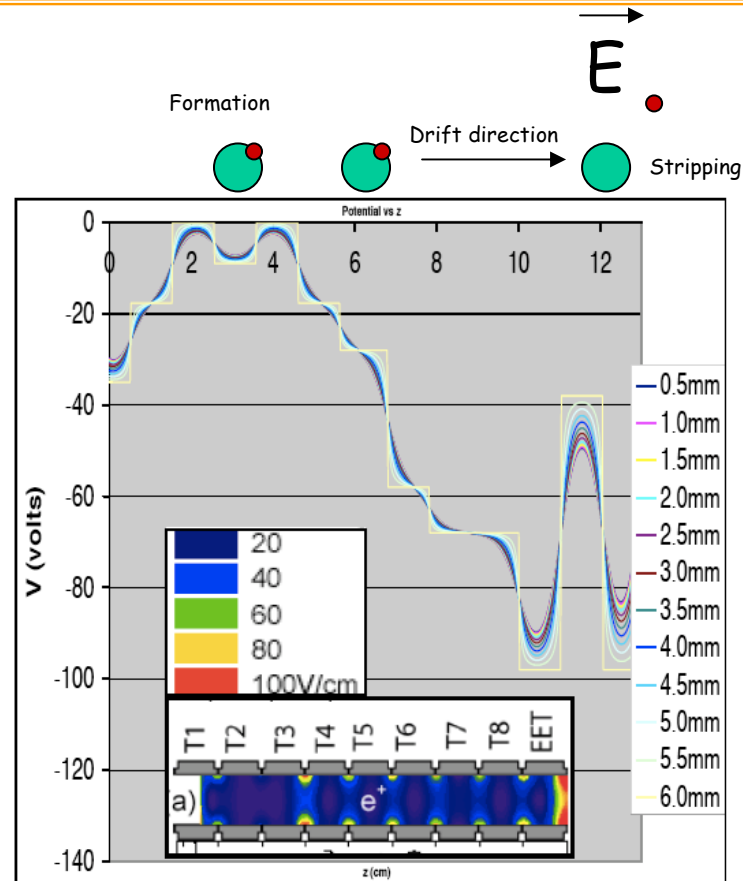
Advantage:

Disadvantage:

Positron attached to 'resting' antiproton

Low rate (14 antihydrogen atoms detected)

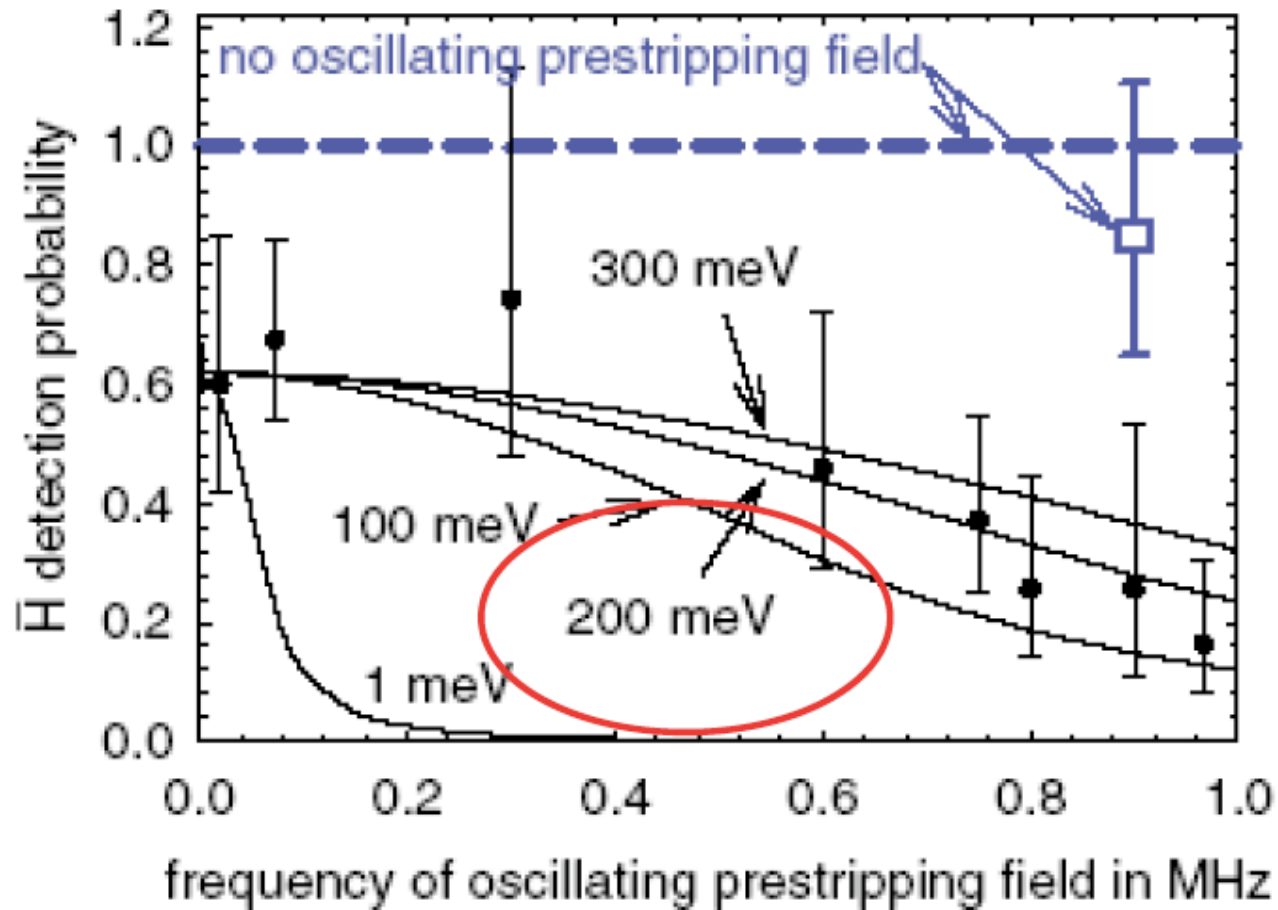
ATRAP - Detection by ionisation



Only works for antihydrogen in high-n states ($n > 35$) moving along magnetic axis
Field ionization - 'rip' positron off, antiproton left and detected by annihilation

ATRAP - Kinetic energy

By using **two** time-variable ionisation+analysis fields, ATRAP measured $E_{\text{kin}} \sim 200 \text{ meV}$ (?)



Summary - Cold antihydrogen production

Two different techniques have been demonstrated:

nested traps (large production rate)

Ps* collisions (small production rate)

Recombination process: a **mixture** of radiative and 3-body transitions

n-state distribution? ATRAP can only observe high-n states (field ionisation)

Velocity of antihydrogen: unknown (1 - 200 meV ?)

MANY IMPORTANT FEATURES NOT YET UNDERSTOOD

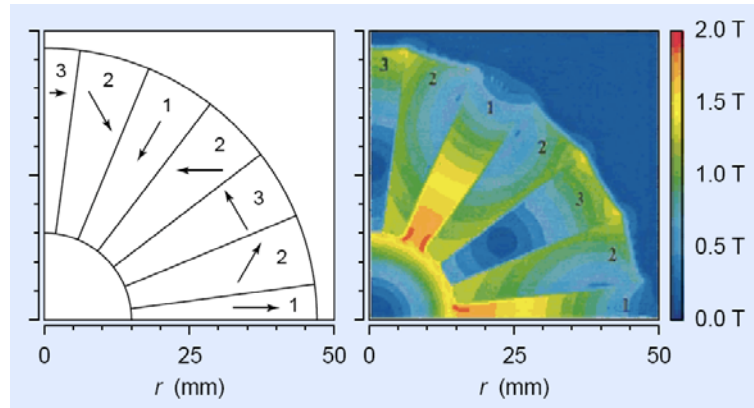
Phase 2 - Trap antihydrogen

How to trap (neutral) antihydrogen?

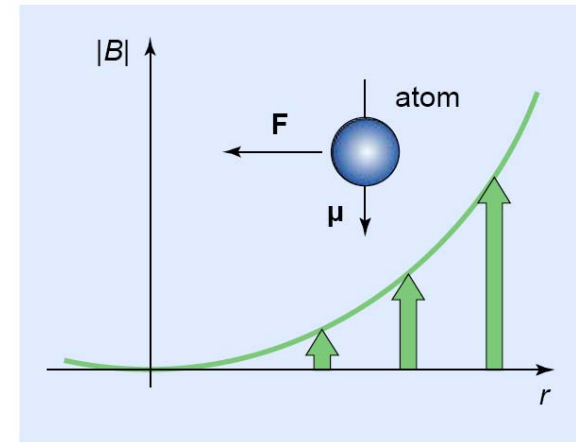
- 1) magnetic moment ($\sim \mu_{e^+}$) - ATRAP/ALPHA approach
- 2) induced electric dipole moment (Stark deceleration) ?
- 3) Formation of positive antihydrogen ions (additional e^+) ??

Magnetic bottle ?

Recombination e.g. inside sextupole magnet



$$U = -\vec{\mu} \cdot \vec{B}$$
$$\vec{F} = -\vec{\nabla} U$$



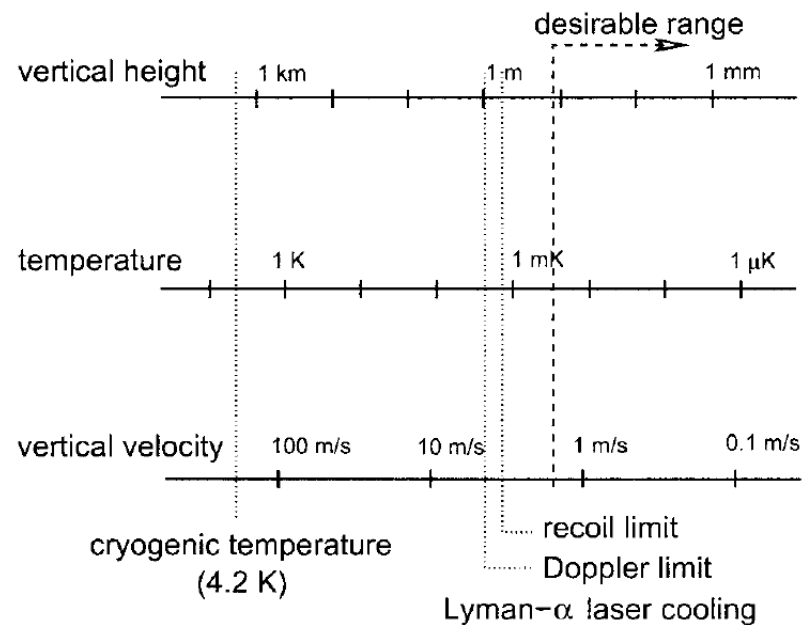
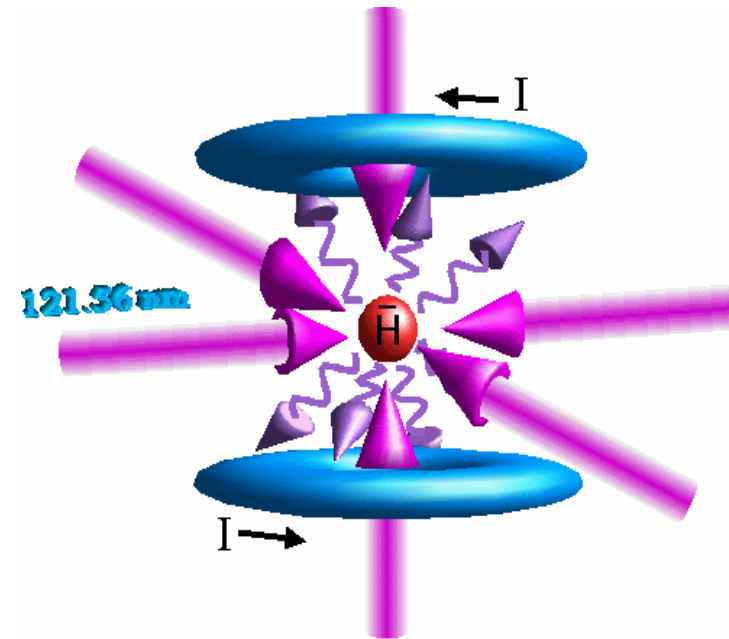
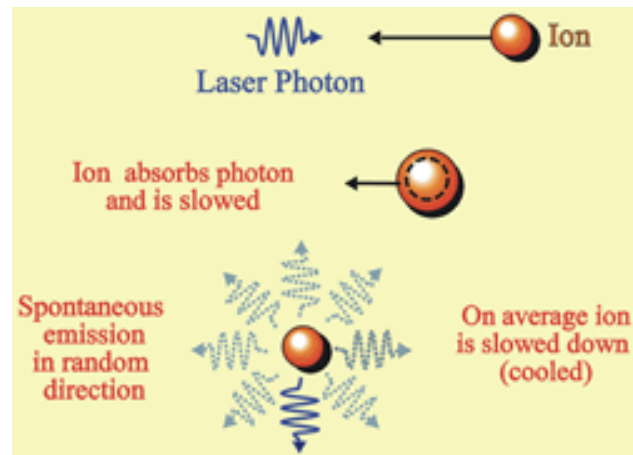
Trap 'low field seeking' atoms (50%) in middle

Very shallow potential (~ 0.07 meV/T)

Realistic $\Delta B \sim 0.2-0.3$ T $\Rightarrow E < 0.02$ meV

(reminder: $E \sim 1-200$ meV)

Antihydrogen cooling



121 nm laser needed
 Prototype at MPI Munich
 ... only 50 nW

Phase 3 - Ideas for precision experiments

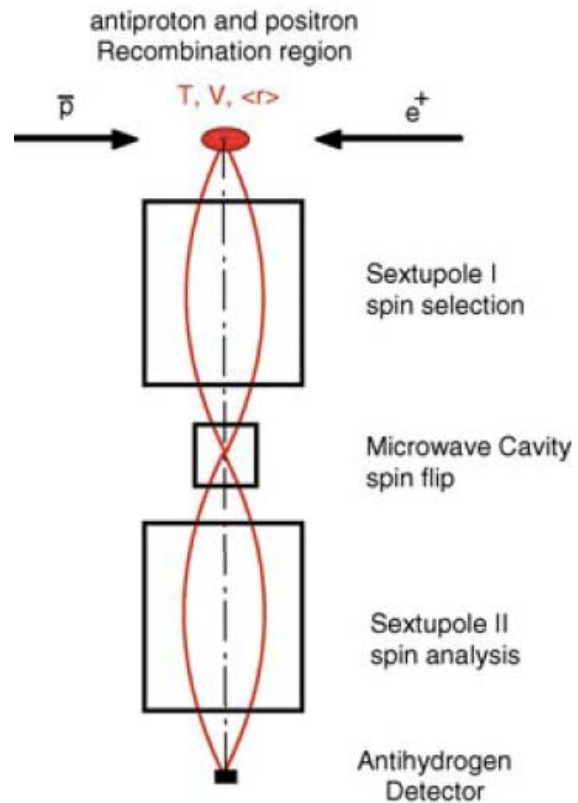
1S-2S Two-photon laser spectroscopy

Hyperfine Structure

Ballistic gravity measurement

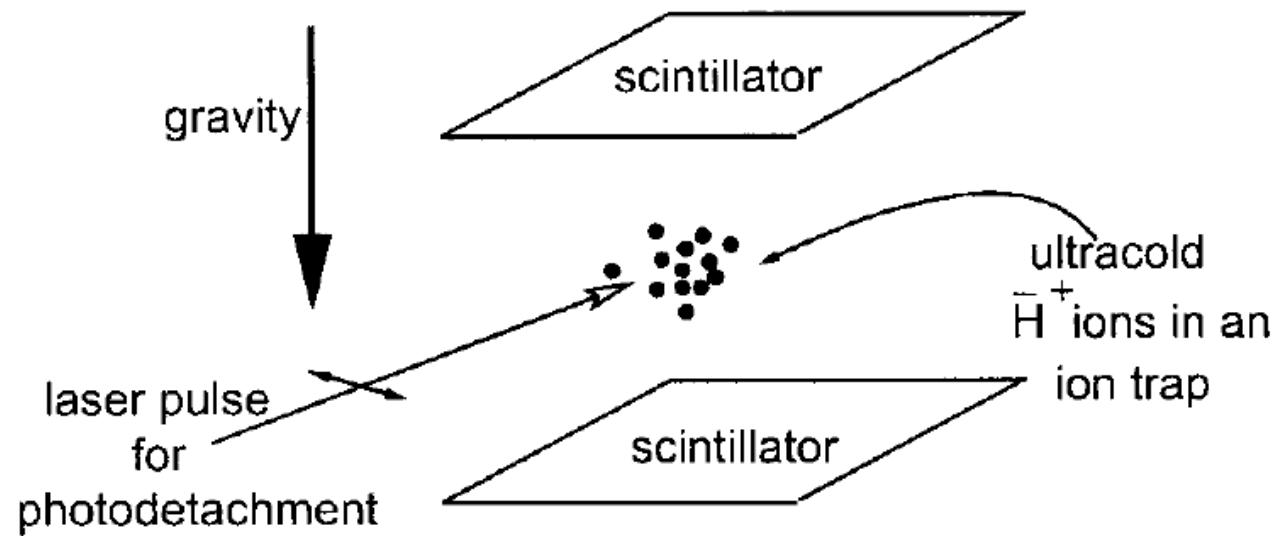
Atomic interferometry

Hyperfine Structure



Conceptual design (E. Widmann et al., CERN-SPSC-2003-009)

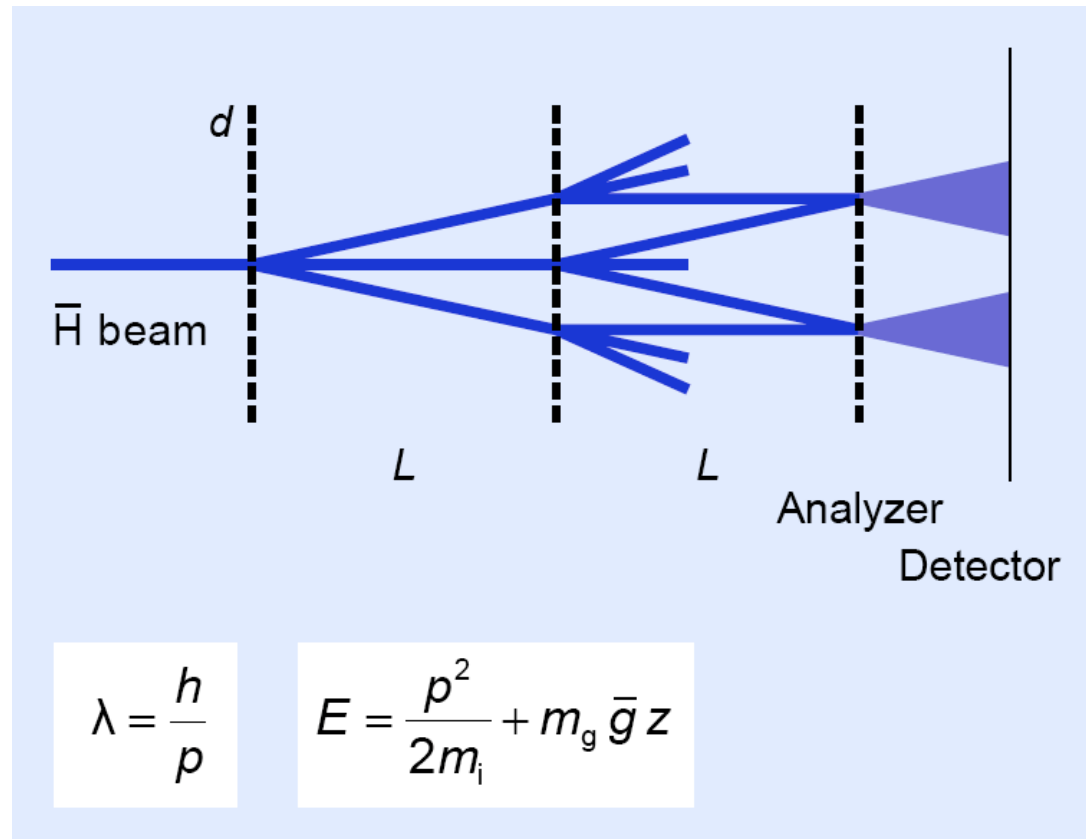
Gravity experiment - ballistic experiment



J. Walz, T. Hänsch, *Gen. Rel. Gravit.* 36 (2004) 561

Gravity experiment - atom interferometer

Mach - Zehnder interferometer



1 mio Hbar:

$$\delta g/g \sim 10^{-4}$$

T.J. Philips, Hyp. Int. 109 (1997) 357

VI. ANTIMATTER - APPLICATIONS AND SPECULATIONS

The wonderful and the weird ...

PET scan

Positron microscope

Tumour therapy

PET isotope production

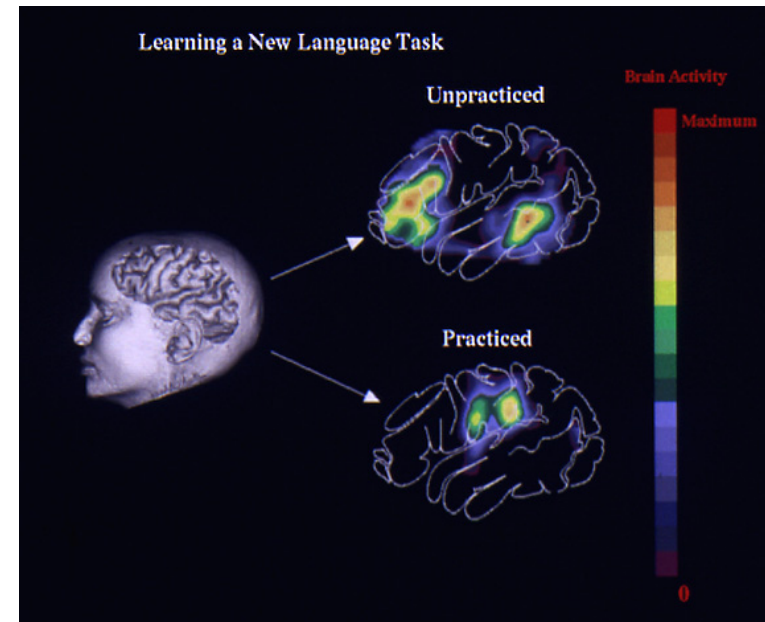
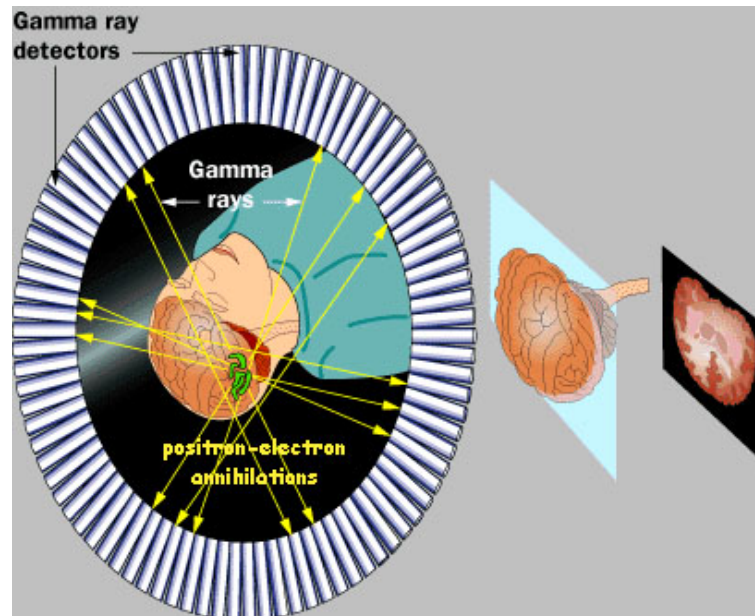
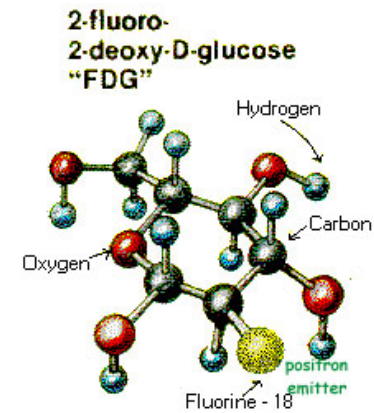
Antiproton Induced Fission/Fusion

Antimatter Propulsion

Positron Emission Tomography (PET)

Insert e^+ emitting isotopes (C-11, N-13, O-15, F-18) into physiologically relevant molecules (O_2 , glucose, enzymes) and inject into patient.

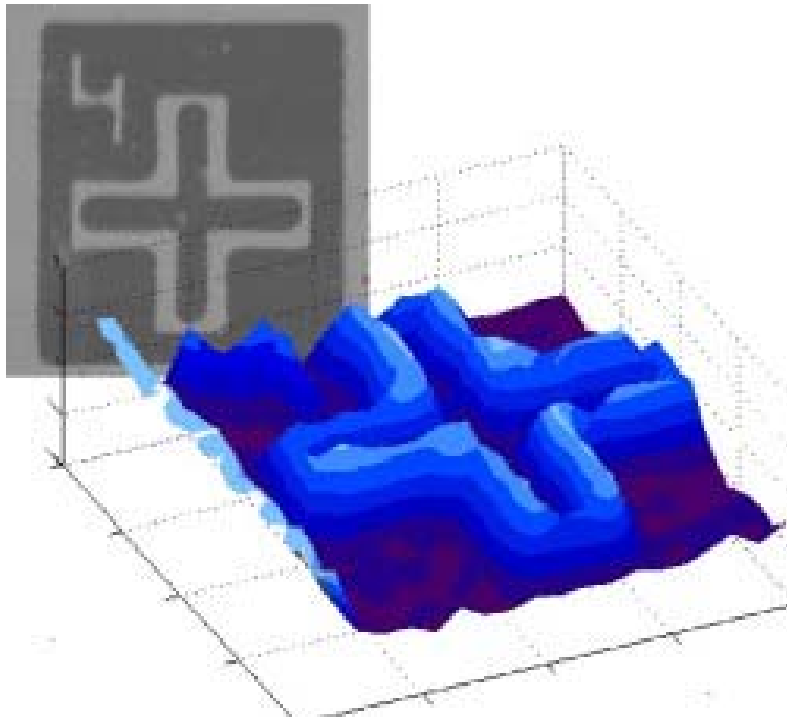
Study positron annihilation with crystal calorimeter (Positron Emission Tomography, PET)



Positron Microscope

Detection of defects in semiconductor surfaces

Positrons at vacancy sites live longer (smaller electron density): **long positron lifetimes indicates vacancies**. Scan surface with pulsed (0.2 ns) positron beam from Na-22 source, $\sim 1 \mu\text{m}$ wide. Time difference from emission to annihilation measures lifetime.



Cross-shaped pattern $\sim 120 \mu$ across
(Pt on SiO surface; positron lifetime in SiO $\sim 2x$
longer than in Pt)

Scanning Positron Microscope: A. David et al., Phys. Rev. Lett. 87 (2001) 067402

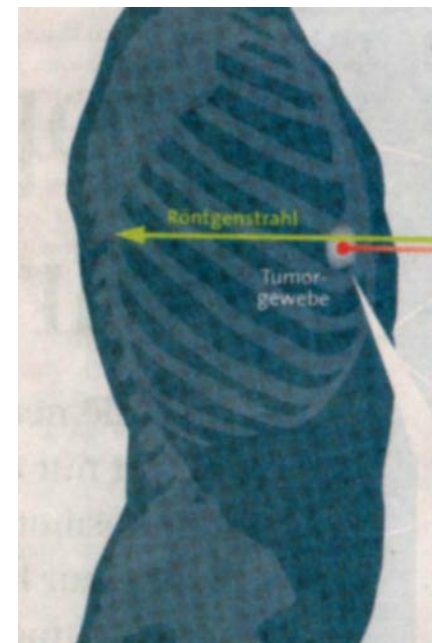
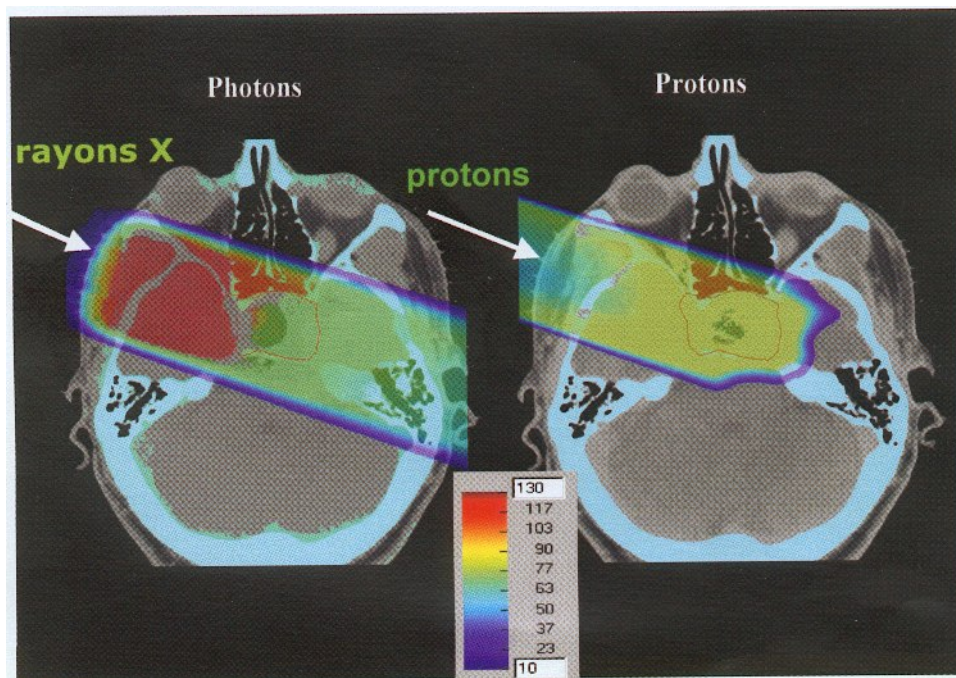
Antiproton Beam Therapy

Goal: destroy tumour without (too much) harm to healthy tissue

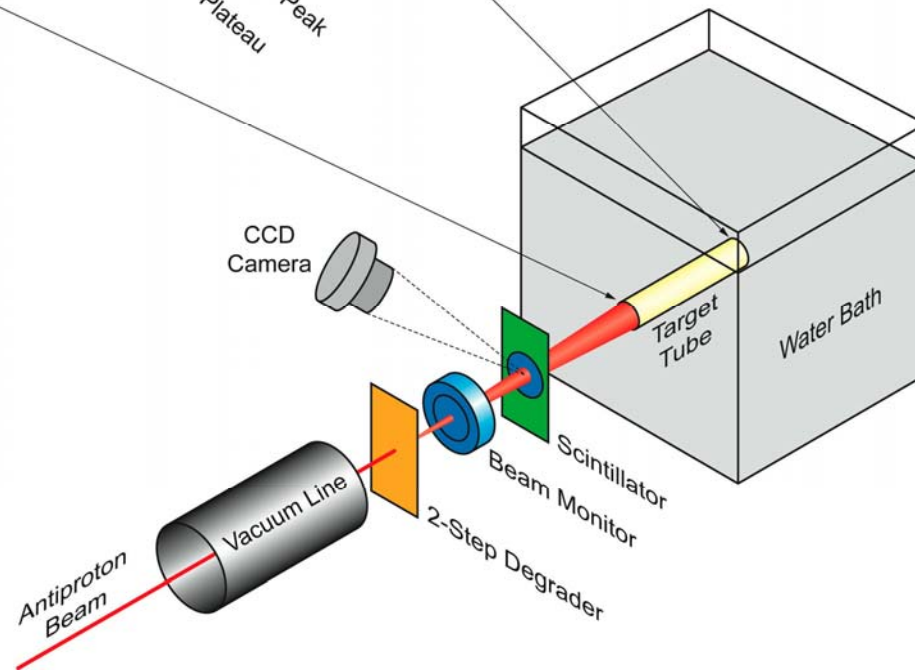
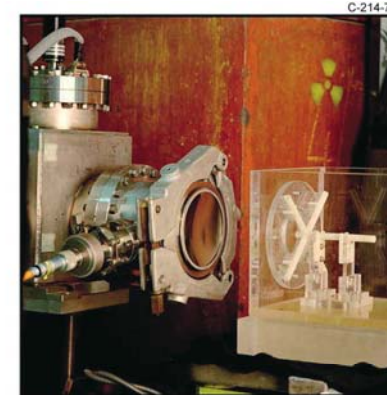
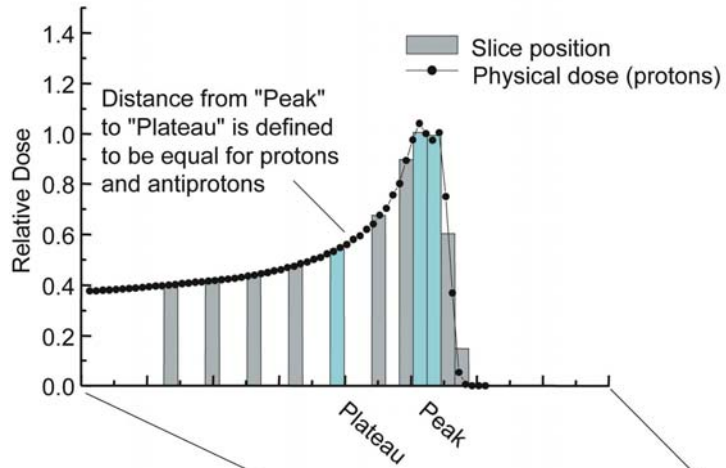
Gammas: exponential decay (peaks at beginning)

Charged particles: Bragg peak (Plateau/Peak better for high Z)

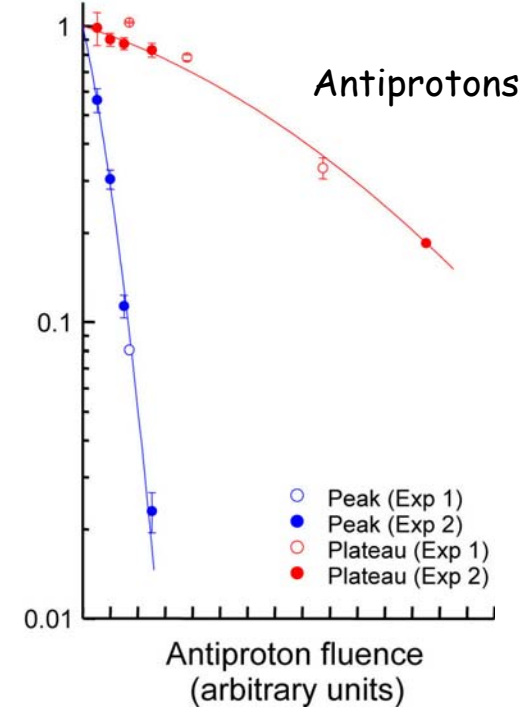
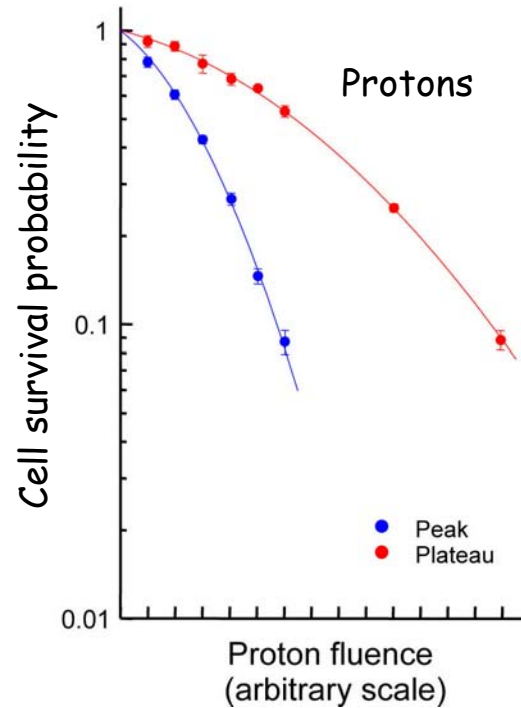
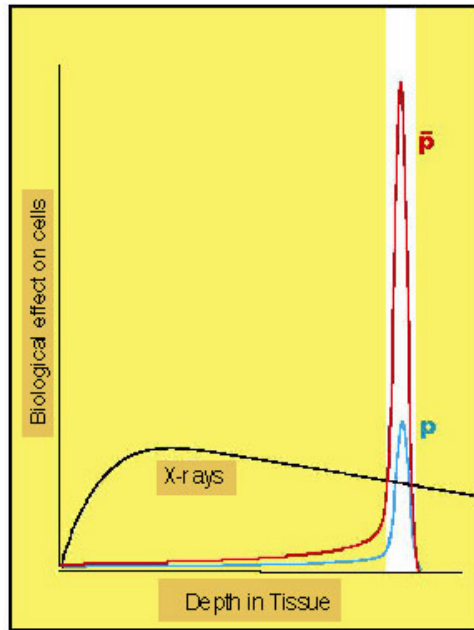
Antiprotons: like protons, but enhanced Bragg peak from annihilation



Antiproton Cell Experiment ACE (AD-4)



Antiprotons beams for tumour treatment ?



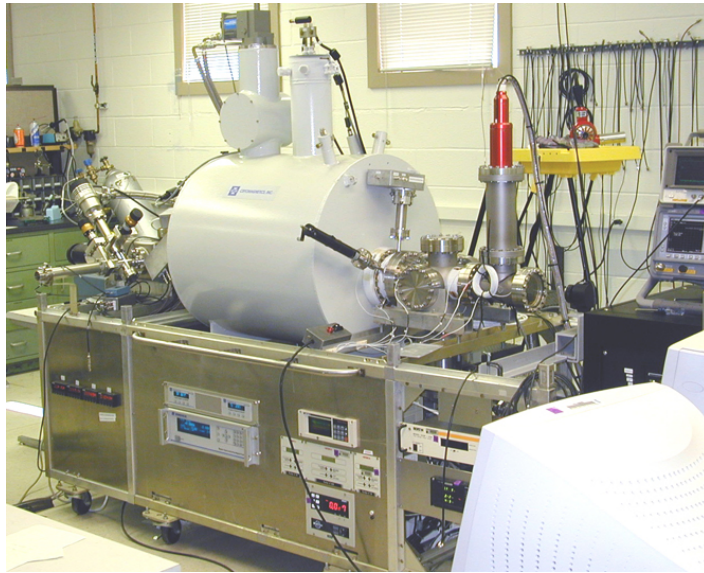
Equal cell mortality for tumour cells with less radiation dose (= damage to healthy cells)
Comparison with Carbon ion therapy

In-situ production of PET isotopes using antiprotons ?

Only 40 isotope production centres
(2 M\$/cyclotron) for 1700 PETs in US

Monte Carlo: 10^{10} - 10^{11} antiprotons for
15 mCi source (F-18)

Problem: short half-life
(C-11: 20', N-13: 10', O-15: 2', F-18: 110')



Plan: $\sim 10^{11}$ antiprotons trapped transport
to hospital and isotope production *in situ*

*Problem: Security! Risk of antiproton loss
requires severe shielding (several tons!)*

NASA/Penn State trap designed for 10^{12} antiprotons

N.b. Trapped electrons have been transported 5000 km; Gabrielse+Tseng, *Hyperfine Interactions* 76, 381 (1993)

The price of 1 g antiprotons*

The only 'official' financial estimate:

A company (Pbar-Tech, now out-of-business) offered to buy 1% of 2002 Fermilab antiprotons ($5 \cdot 10^{14}$ antiprotons/year): 274 K\$ for $5 \cdot 10^{12}$ antiprotons

1g at Fermilab: 32,000,000 billion \$

Production is dominated by power consumption:

CERN (PS Complex) ~ 30 MW. For a bunch of $5 \cdot 10^7$ antiprotons, need 2.4 s of operation. At a price of 0.1 € per kWh, this costs 2€ per bunch.

Energy consumption for 1 g antiprotons: $9 \cdot 10^{23}$ J

1g at CERN*: 24,000,000 billion €

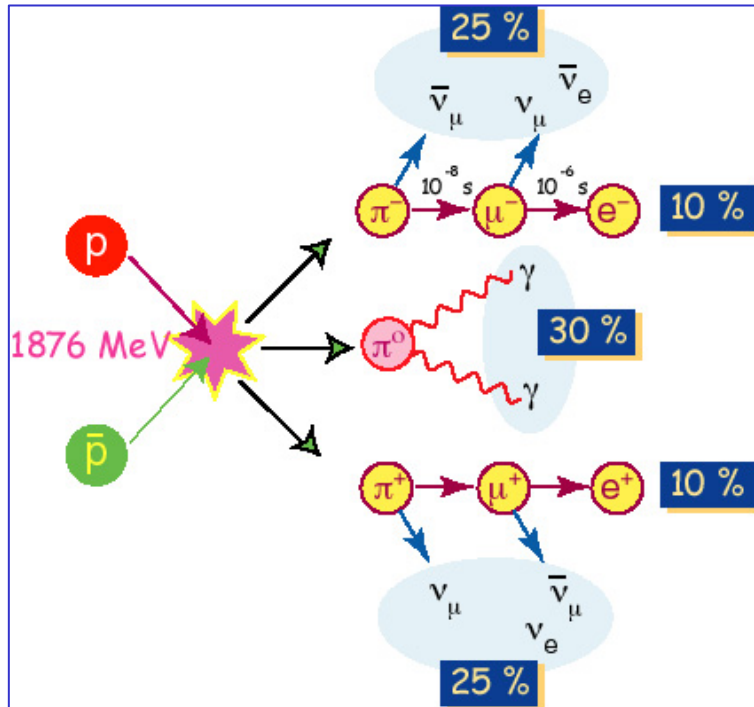
For comparison:

US debt (total, 2006):	8,400 billion \$	(~ 1/4000 g antimatter)
US defense budget:	540 billion \$	(~ 17 μ g)
World energy consumption (2003):	$4 \cdot 10^{20}$ J	(2250 years for 1 g)
1 μ g antiprotons	24 billion \$	

*allegedly stolen from CERN; see D. Brown "Angels+Demons"

Antimatter for space propulsion ??

Ambitious long-term plans: "Mars express", Oort cloud, Proxima Centauri, ...



Energy balance for pbar-p annihilation

Annihilation: energy release/mass =
 $10^{10} \times \text{H}_2/\text{O}_2$ combustion;
100 x fusion!

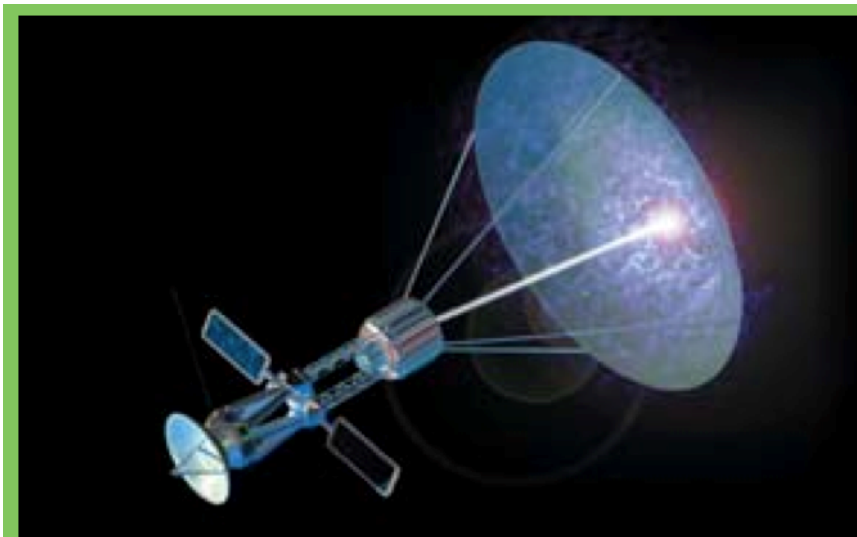
PROBLEM

Conversion of reaction energy into propulsive thrust
-> quick decay into gamma rays and neutrinos

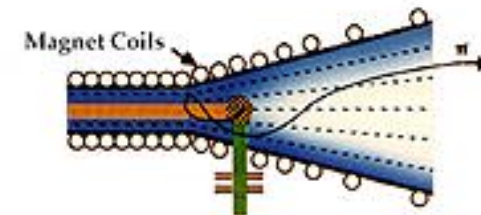
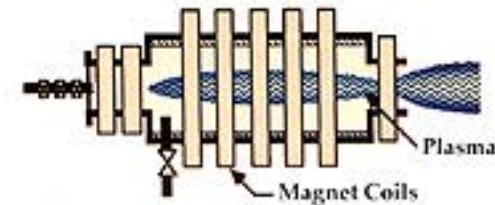
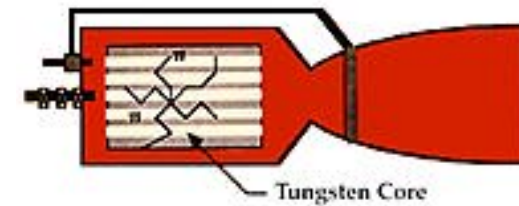
<http://www.msfc.nasa.gov/STD/propulsion/research/anti/>

Antimatter driven space engines?

- 1) solid and gas-core thermal rockets (pion heating)
- 2) antiproton-initiated fission driving fusion rockets (needs 1- 100 μg)



PROPOSED ANTIMATTER PROPULSION system uses antimatter pellets to trigger fission explosions on a uranium-coated sail.



PROBLEMS

- antihydrogen storage
- production cost
- Concept of antiproton-induced fission

Concluding remarks

- Antiparticles discovered 74 years ago
- Very important for the development of particle physics and cosmology
- Antiproton production, deceleration and cooling techniques well understood
- Cold antihydrogen atoms have been produced (> 1 mio)
- Next big challenge: trapping and cooling
- Antihydrogen atoms ideal for studying CPT invariance and antimatter gravity with very high precision
- Applications in daily life exist, but high production costs of antiprotons are (almost) prohibitive

Thank you for your attention
