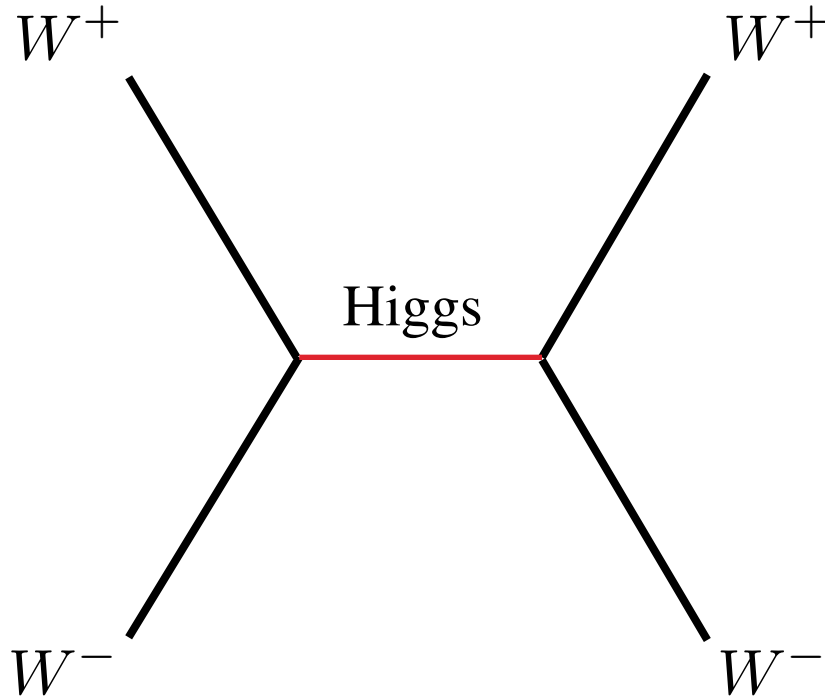


# The Higgs Completes the Standard Model



$$\lim_{E \rightarrow \infty} \mathcal{A} \propto \text{const.}$$

With the inclusion of the Higgs particle, the theory remains predictive.

Theory requires a Higgs mass  $< 1$  TeV

IV

Phenomenology of  
Supersymmetry:  
What Your Mother Never  
Told You

# Biggest Problem for the MSSM

We didn't see the *Higgs*.  
(not superpartners)

# The Higgs Potential

$$\lambda|h|^4 \rightarrow \frac{g^2}{8} [ |H_1|^2 - |H_2|^2 ]^2 \quad m_h = M_Z |\cos 2\beta|$$

SUSY-breaking loop required - same size as tree.

$$(m_h^2)_{tree} + \delta m_h^2 > (114 \text{ GeV})^2 \quad (\text{Big Susy-breaking in top sector})$$

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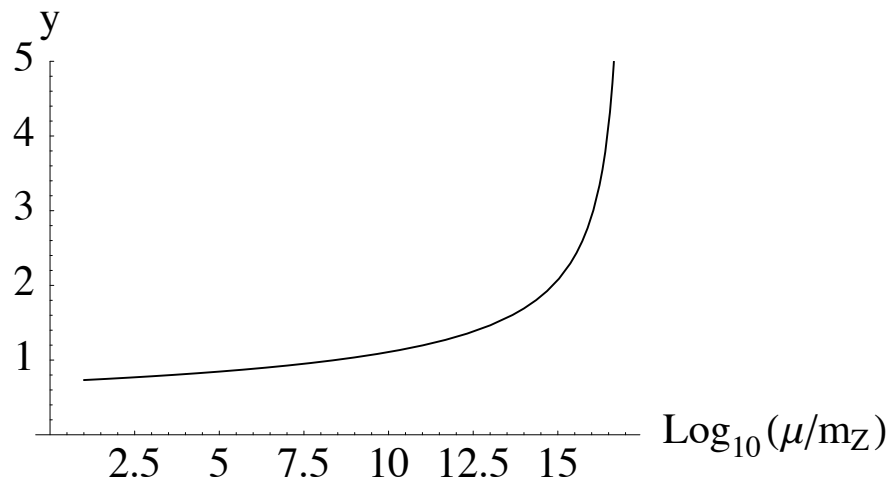
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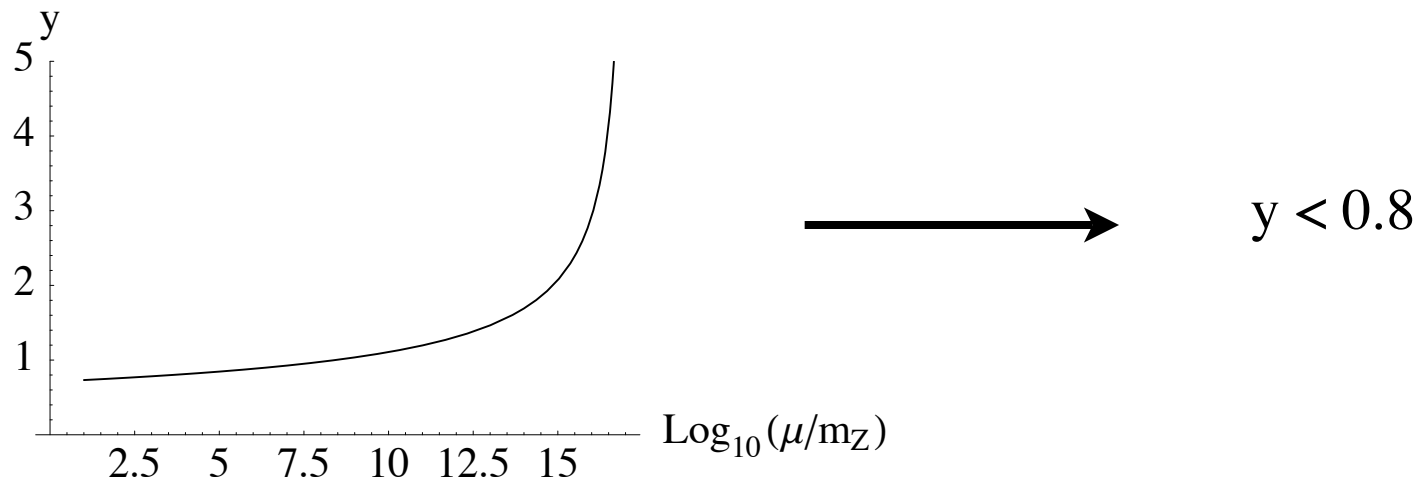
New tree-level contribution requires new fields:

$$W = ySH_1H_2 \rightarrow y^2 |H_1H_2|^2$$

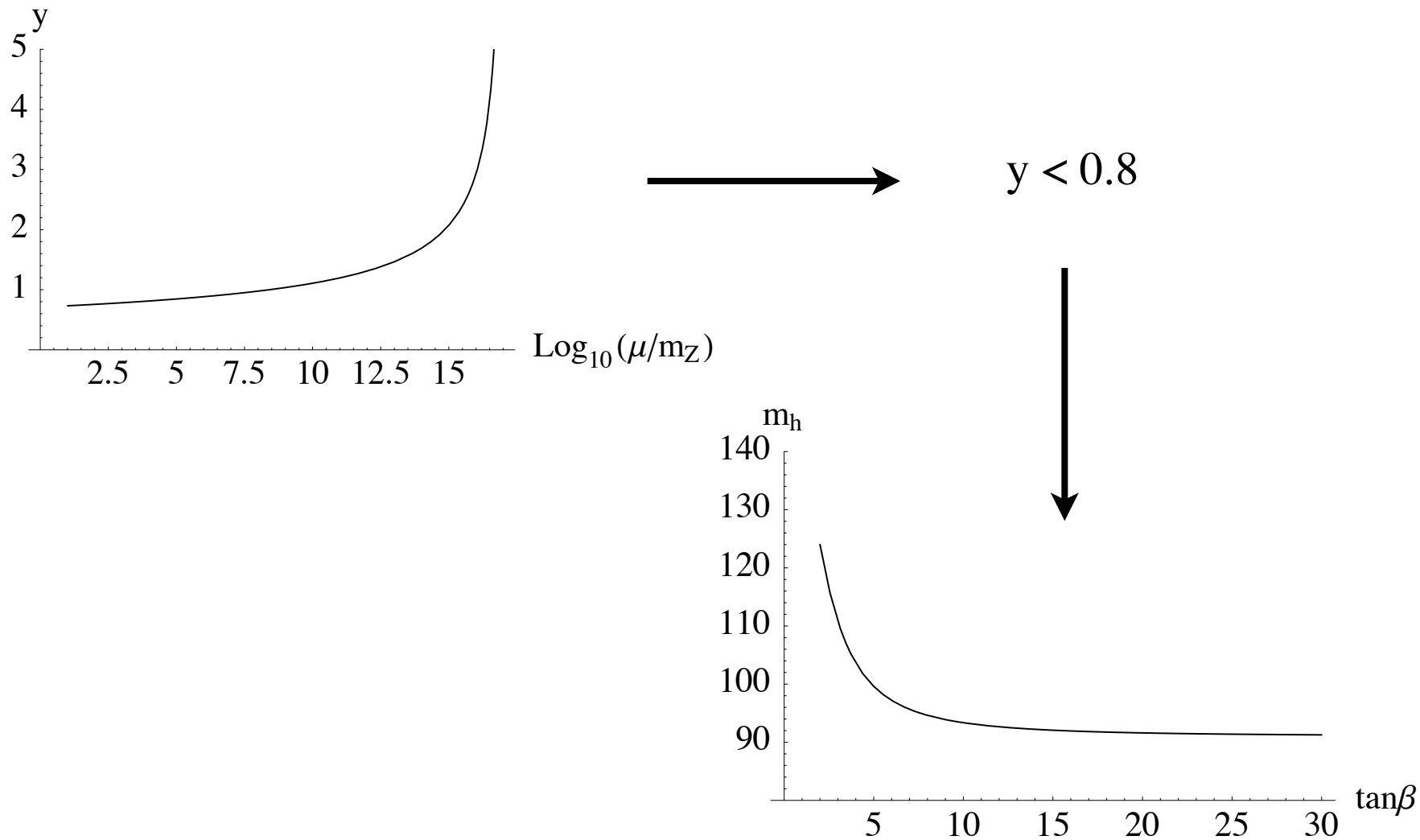
# Limits on the New Quartic



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# More Higgses

$$W = \lambda \hat{H}_u \hat{H}_d \hat{S} + \frac{\kappa}{3} \hat{S}^3$$

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$$W = \lambda \hat{H}_u \hat{H}_d \hat{S} + \frac{\kappa}{3} \hat{S}^3$$

$$M_{H_d}^2 = -\frac{\lambda^2}{2} (s^2 + v^2 \sin^2 \beta) + \frac{\lambda \kappa}{2} s^2 \tan \beta - \frac{M_Z^2}{2} \cos 2\beta + m_\lambda s \tan \beta ,$$

$$M_{H_u}^2 = -\frac{\lambda^2}{2} (s^2 + v^2 \cos^2 \beta) + \frac{\lambda \kappa s^2}{2 \tan \beta} + \frac{M_Z^2}{2} \cos 2\beta + \frac{m_\lambda s}{\tan \beta} ,$$

$$M_S^2 = -\frac{\lambda^2}{2} v^2 + \frac{\lambda \kappa}{2} v^2 \sin 2\beta - \kappa^2 s^2 + \frac{m_\lambda v^2}{2s} \sin 2\beta + m_\kappa s .$$

# More Higgses

$$W = \lambda \hat{H}_u \hat{H}_d \hat{S} + \frac{\kappa}{3} \hat{S}^3$$

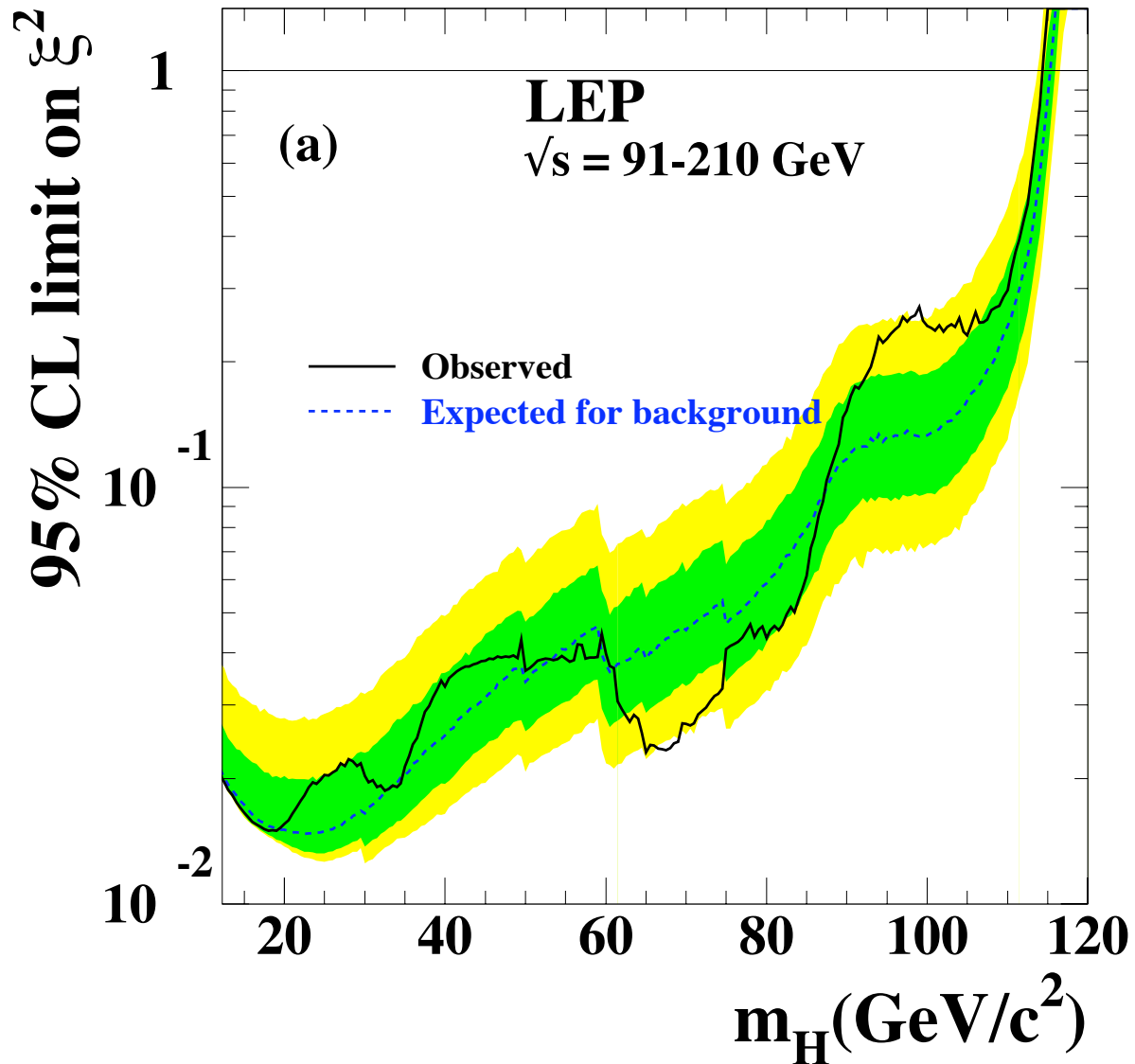
$$h_v^0, H_v^0, h_s^0 \quad A_v^0, A_s^0$$

# New Higgs Decay

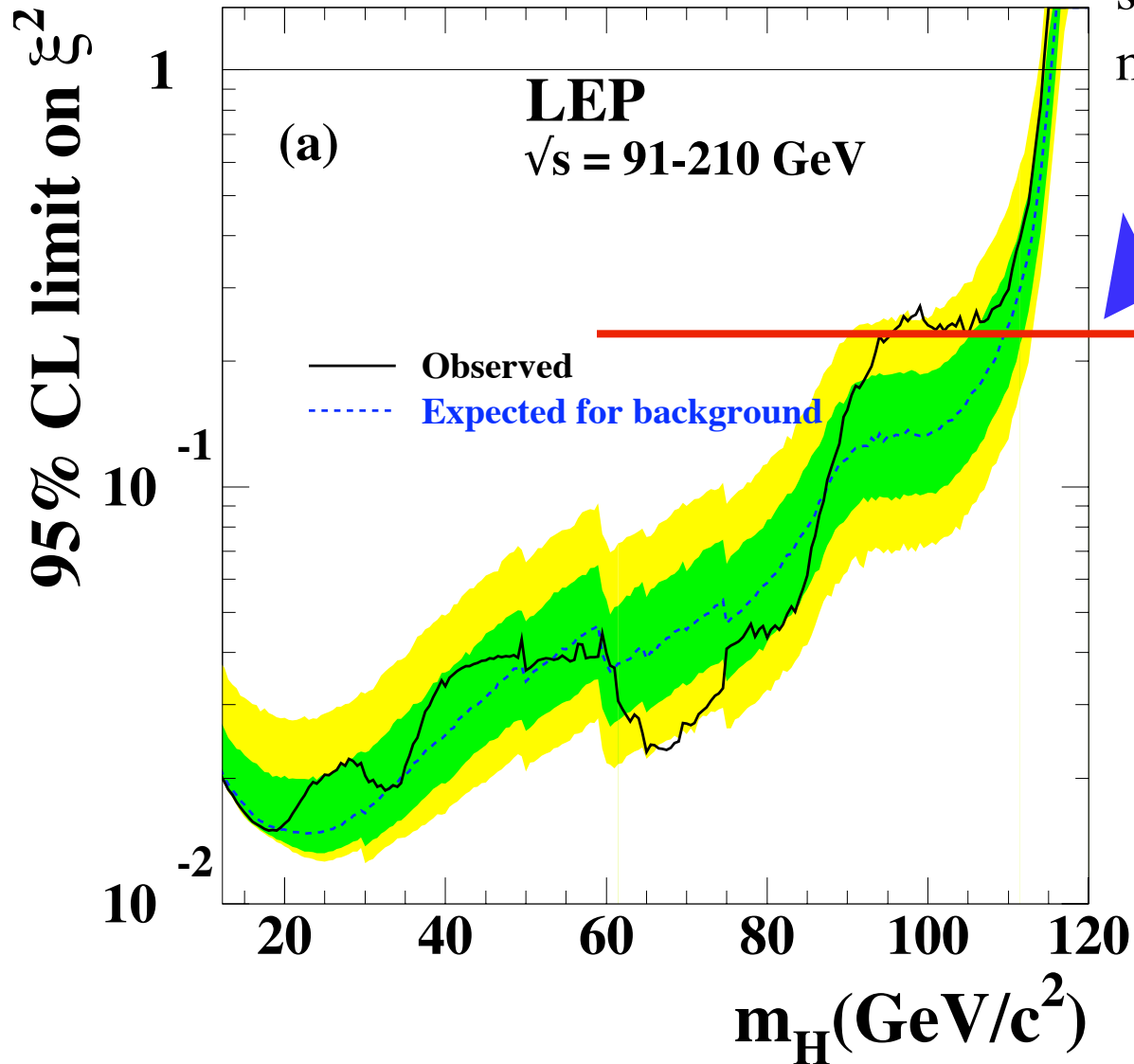
$$h^0 \longrightarrow a^0 a^0$$

And then the pseudo-scalars decay.

# SM Higgs

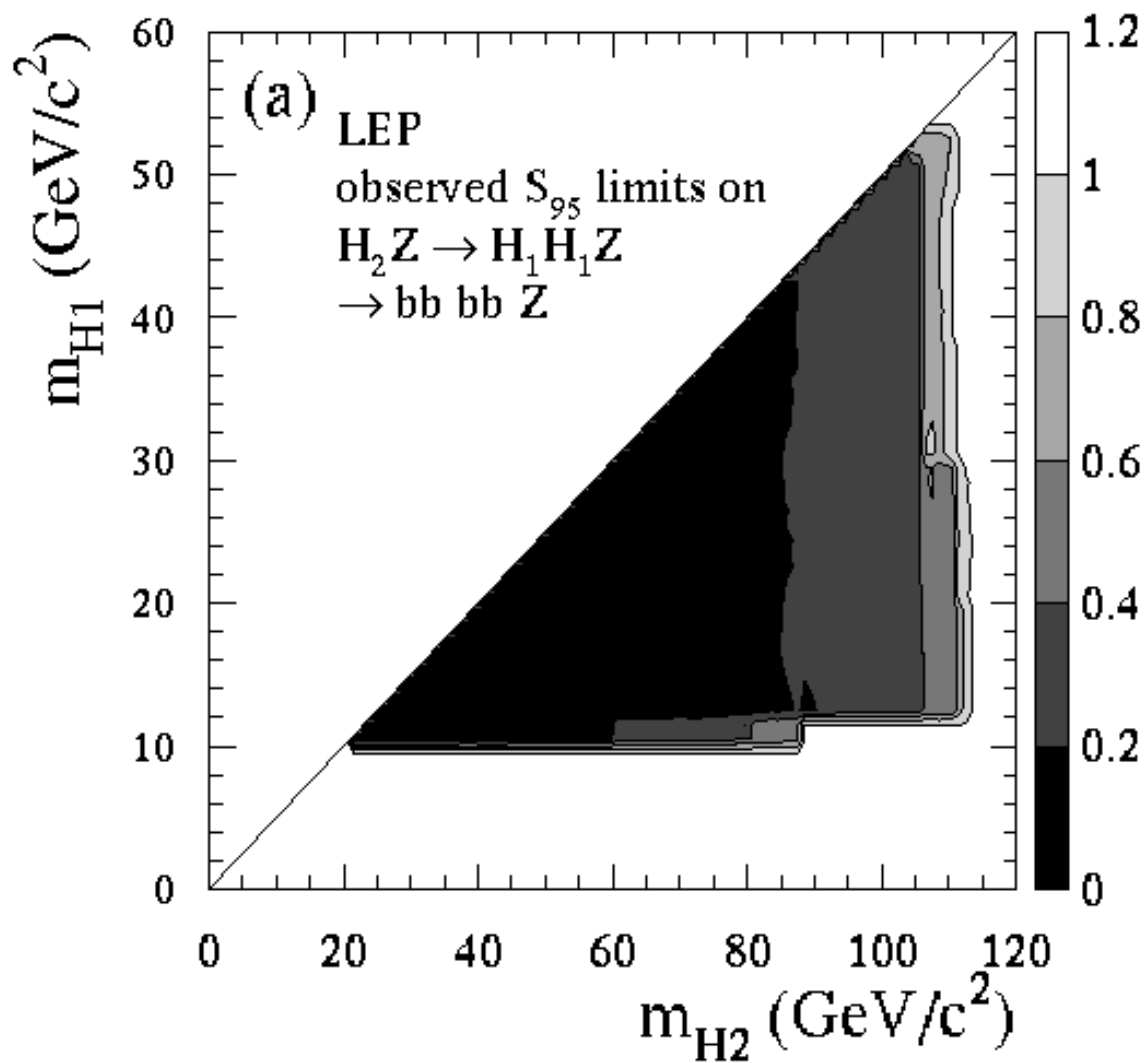


# SM Higgs



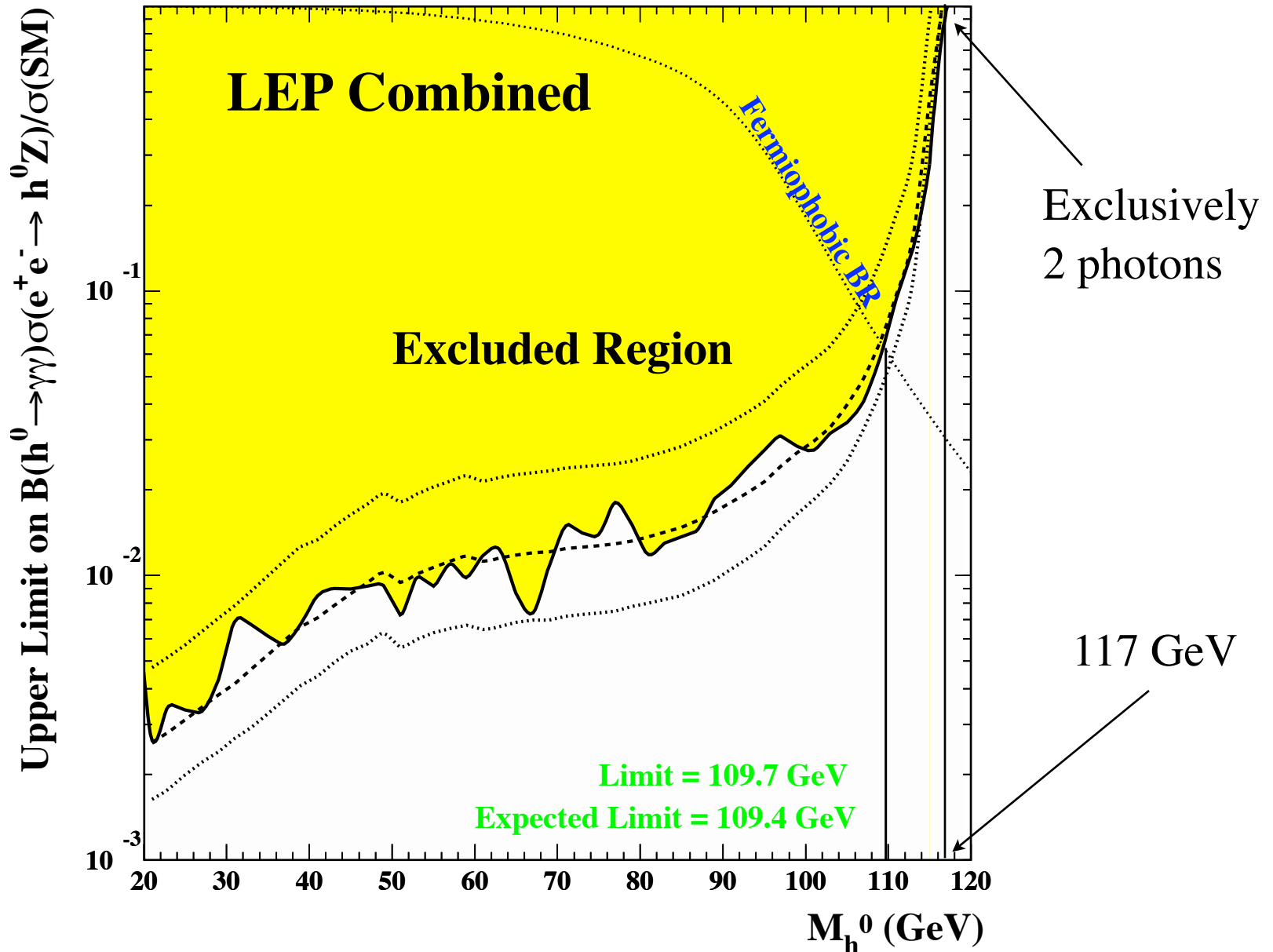
Suppress cross section or BR to standard model modes to  $\sim 20\%$

# Higgs to 4b's



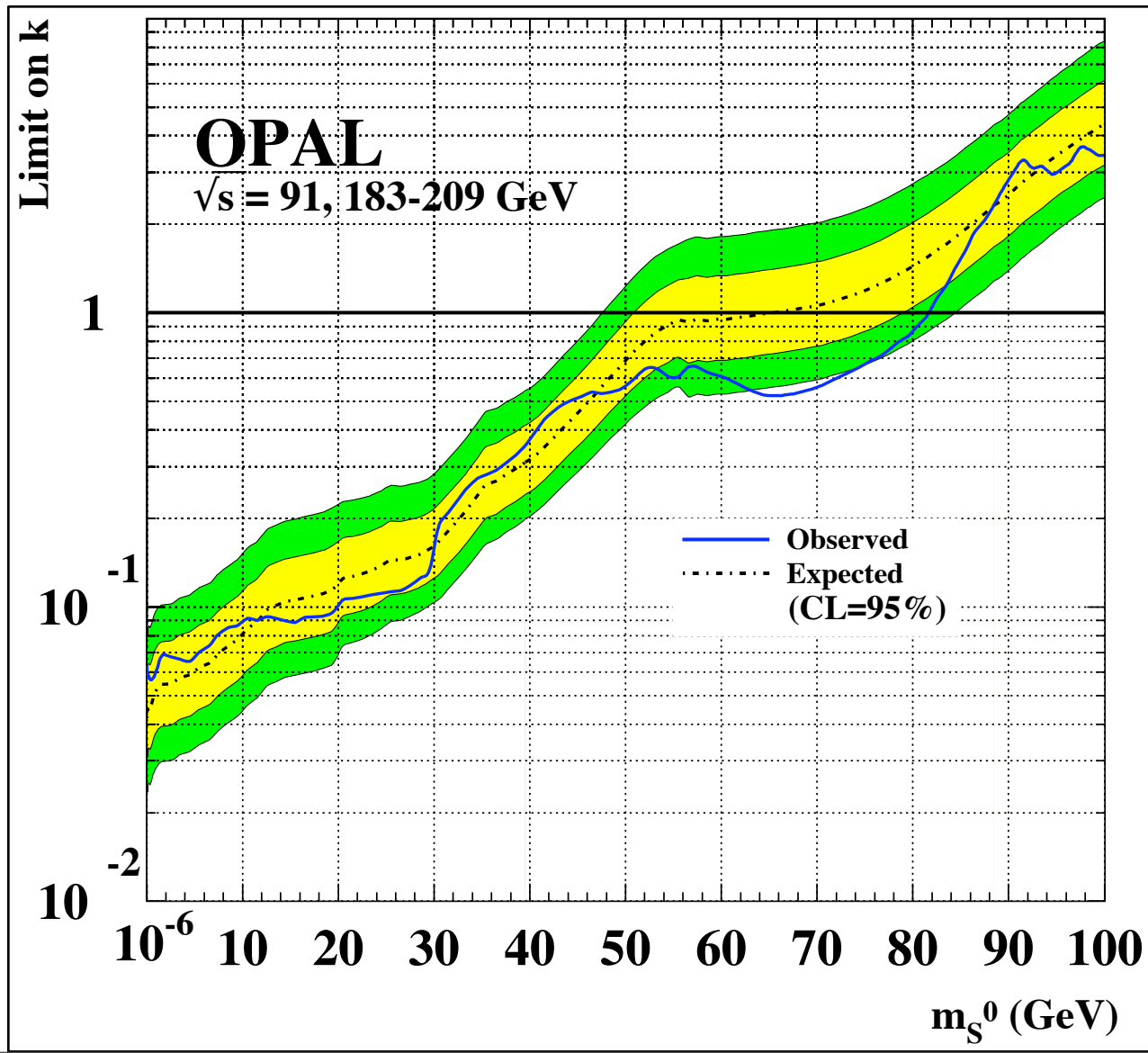
But NOT 4 taus!

# Fermiophobic





# Non-standard Higgs Decays



# New Higgs Decays

$$h \rightarrow a^0 a^0 \rightarrow b\bar{b}b\bar{b} \quad m_h > 110 \text{ GeV}$$

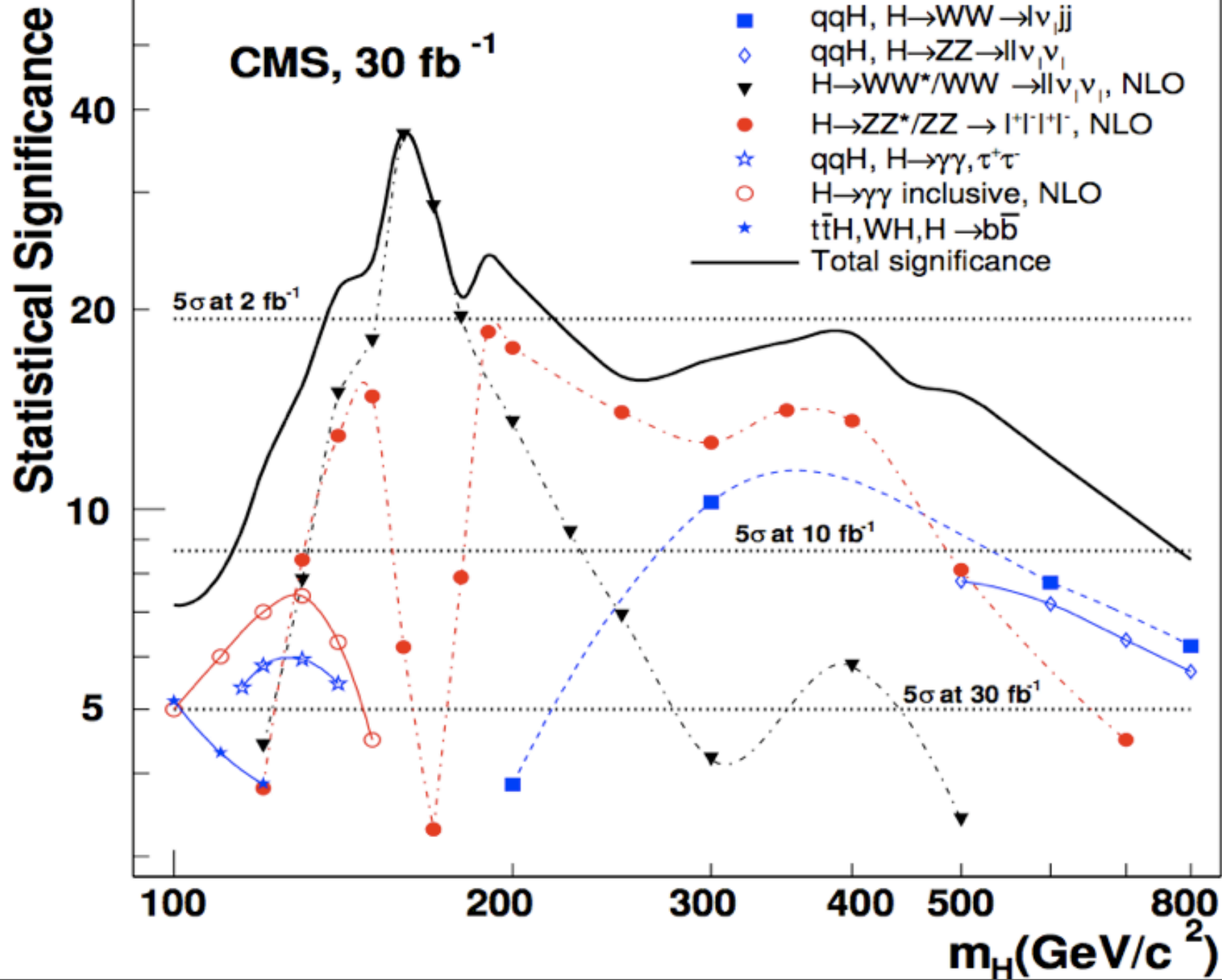
$$h \rightarrow a^0 a^0 \rightarrow \tau\bar{\tau}\tau\bar{\tau} \quad m_h > 86 \text{ GeV}$$

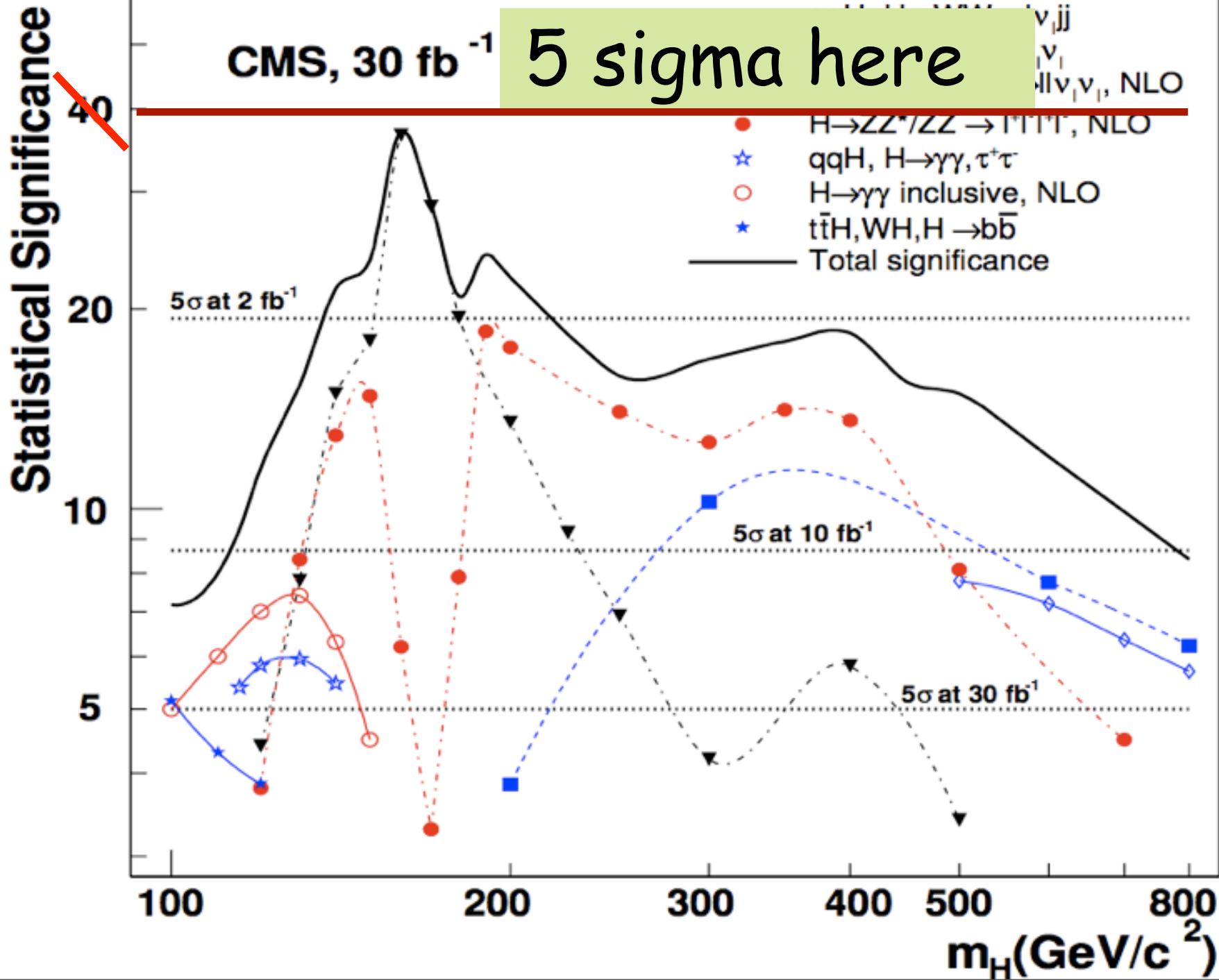
$$h \rightarrow a^0 a^0 \rightarrow gggg \quad m_h > 86 - 100 \text{ GeV?}$$

$$h \rightarrow a^0 a^0 \rightarrow 6\pi^0 \quad m_h > 117 \text{ GeV}$$

$$h \rightarrow a^0 a^0 \rightarrow 4\gamma \quad m_h > 117 \text{ GeV}$$

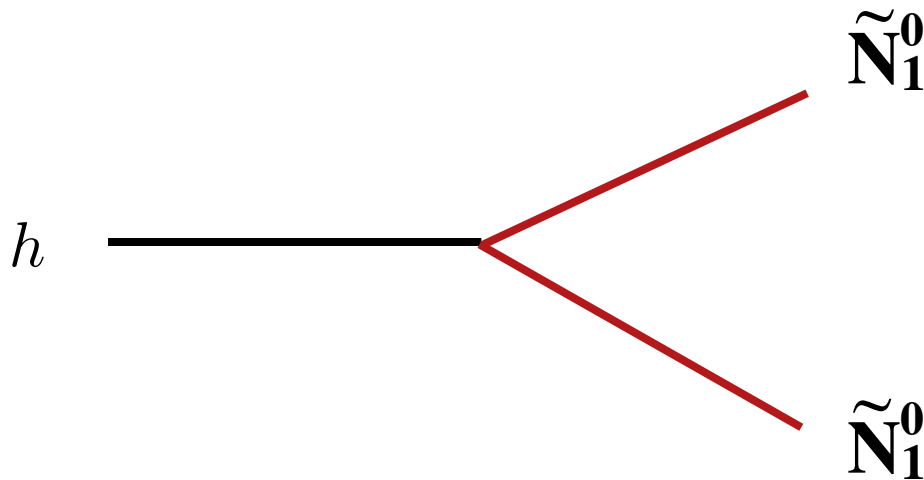
$$h \rightarrow ss \rightarrow a^0 a^0 a^0 a^0 \\ \rightarrow b\bar{b}b\bar{b}b\bar{b}b\bar{b} \quad m_h > 82 \text{ GeV???$$





# Higgs Decays to Superpartners

Most robust possibility:

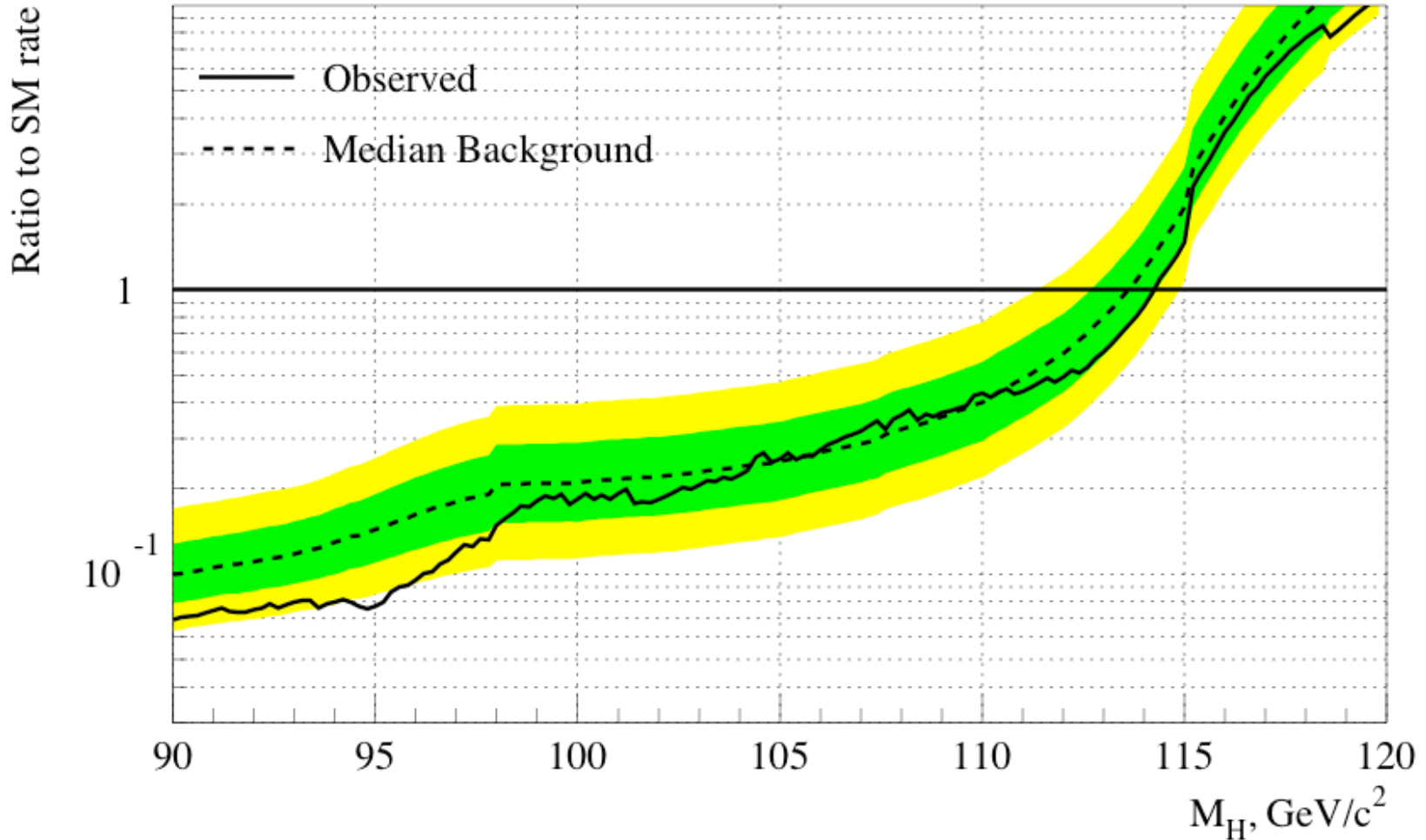


Neutralinos (like neutrinos) are invisible.

$$\tilde{N}_1^0 = \tilde{\chi}_1^0$$

# Invisible Higgs

114 GeV - same  
as SM Higgs



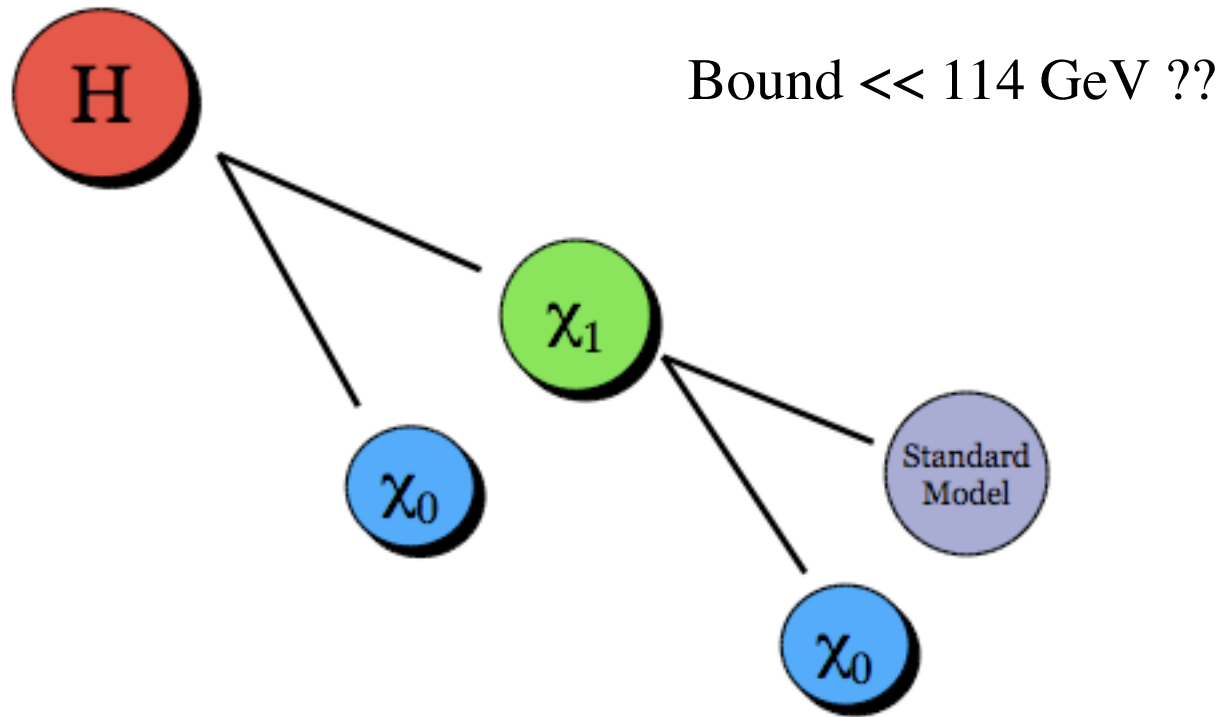
LEP Combined - '01

# Extra Singlet $\rightarrow$ Extra Neutralino

$$W = \lambda \hat{H}_u \hat{H}_d \hat{S} + \frac{\kappa}{3} \hat{S}^3$$

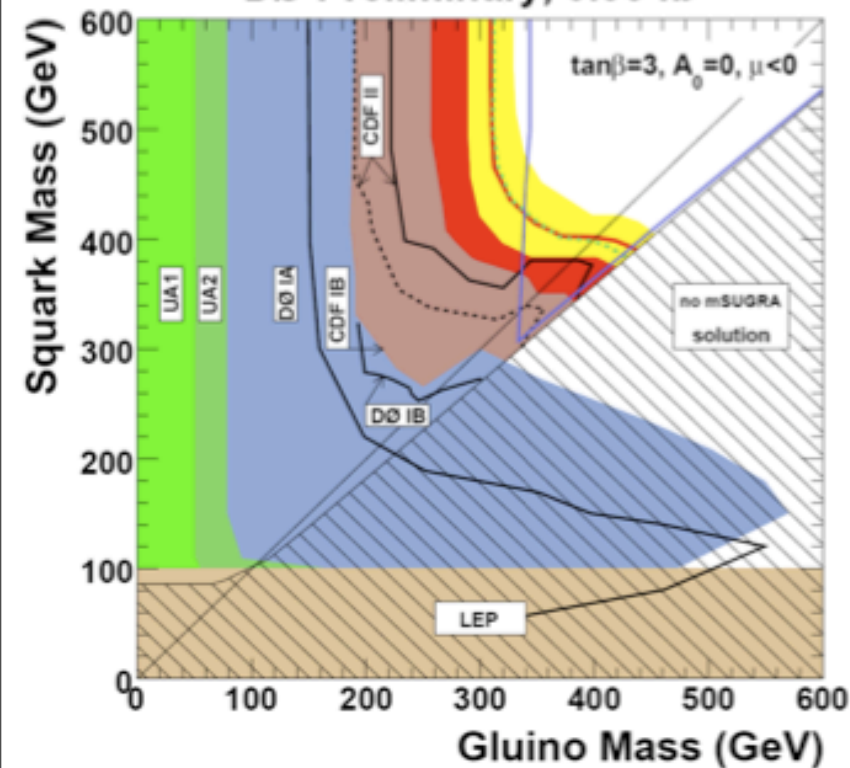
$$\mathbf{M}_{\tilde{N}} = \begin{pmatrix} M_1 & 0 & -g'v_d/\sqrt{2} & g'v_u/\sqrt{2} & 0 \\ 0 & M_2 & gv_d/\sqrt{2} & -gv_u/\sqrt{2} & 0 \\ -g'v_d/\sqrt{2} & gv_d/\sqrt{2} & 0 & -\lambda s & -\lambda v_u \\ g'v_u/\sqrt{2} & -gv_u/\sqrt{2} & -\lambda s & 0 & -\lambda v_d \\ 0 & 0 & -\lambda v_u & -\lambda v_d & 2\kappa s \end{pmatrix}$$

# Higgs to Neutralino Cascade

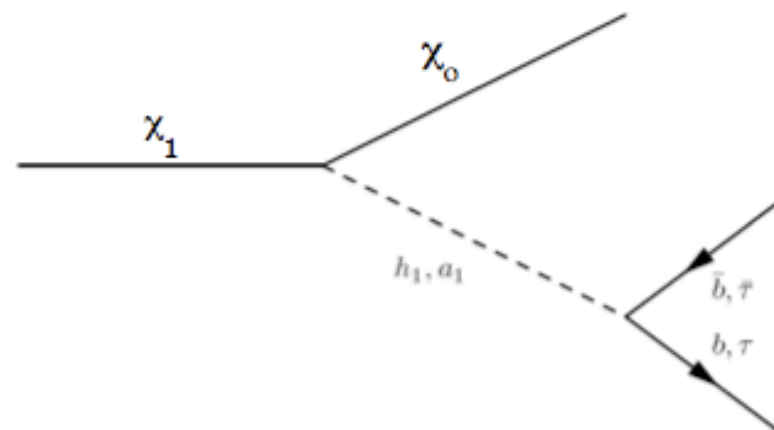




DØ Preliminary,  $0.96 \text{ fb}^{-1}$



Missing Energy signature **suppressed**,  
e.g.



As  $m_a$  approaches  $m_{\chi_1}$ , missing energy is reduced

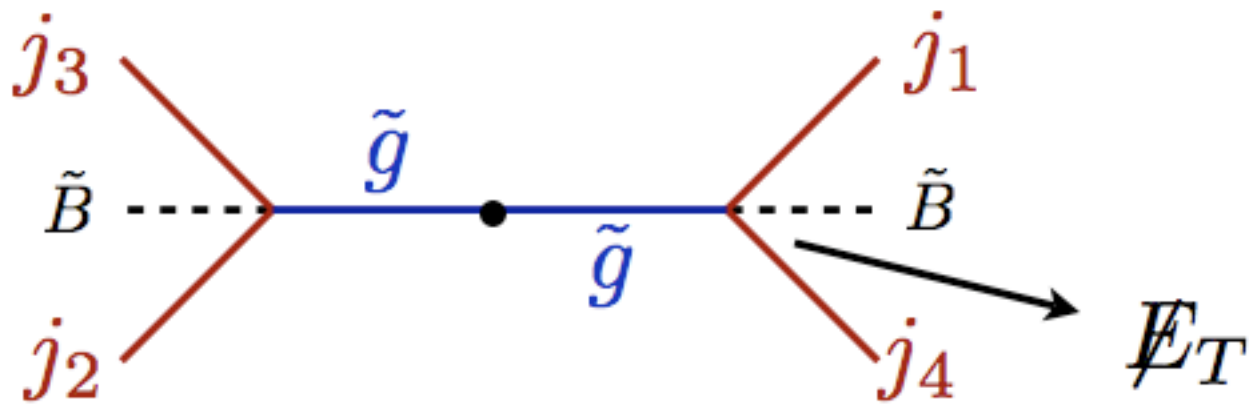
# Generalizing This Ratio

$$M_3 : M_2 : M_1 \simeq 7 : 2 : 1$$

What if, for example,  $M_2 > M_3 \gtrsim M_1$ ?

# Generalizing This Ratio

What if, for example,  $M_2 > M_3 \gtrsim M_1$ ?



$$Q = M_{\tilde{g}} - M_{\tilde{B}}$$

If  $Q < M_{\tilde{B}}$

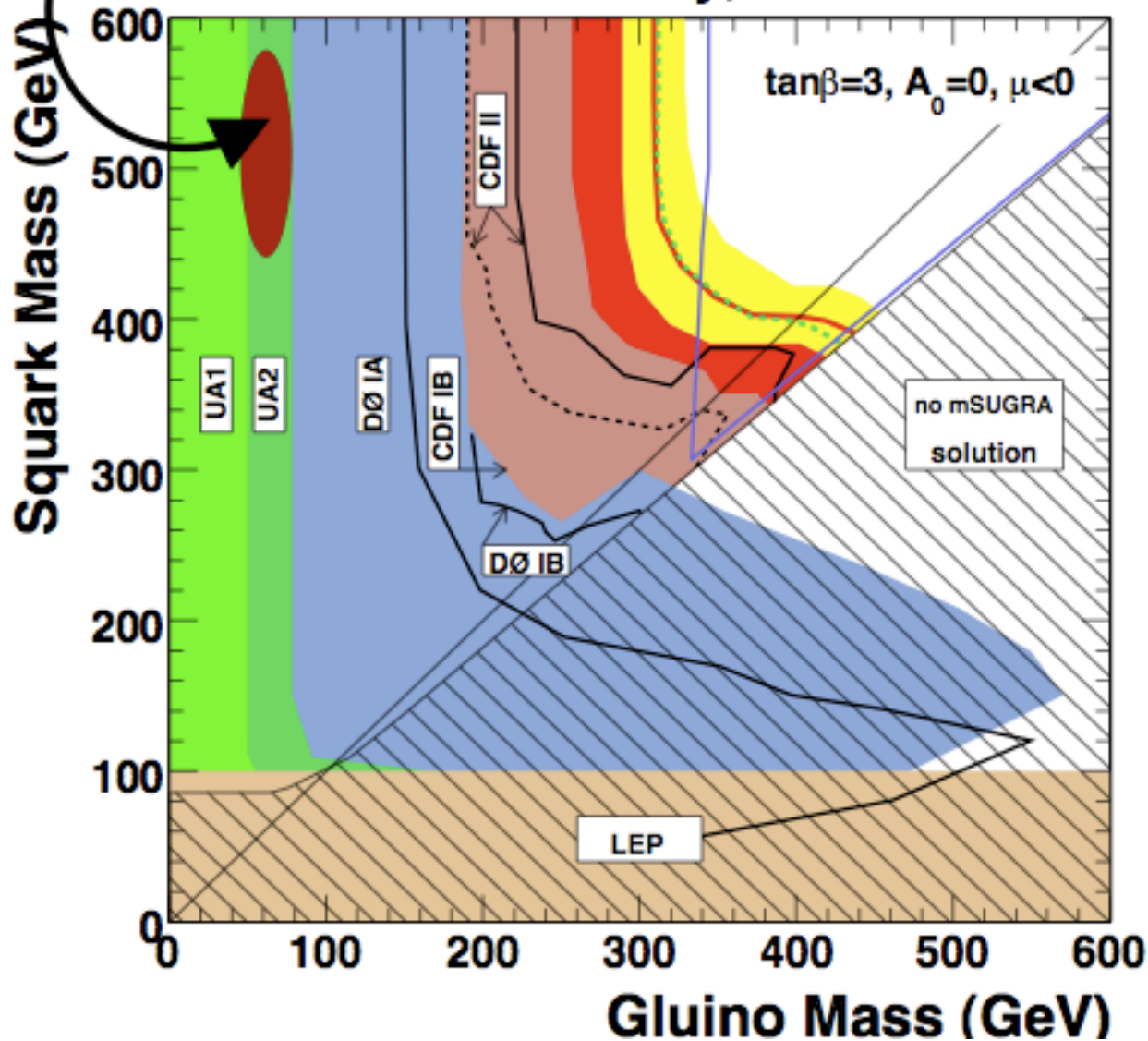
Bino carries away energy but not momentum

$$\cancel{E}_T \sim \frac{Q^2}{M_{\tilde{B}}}$$

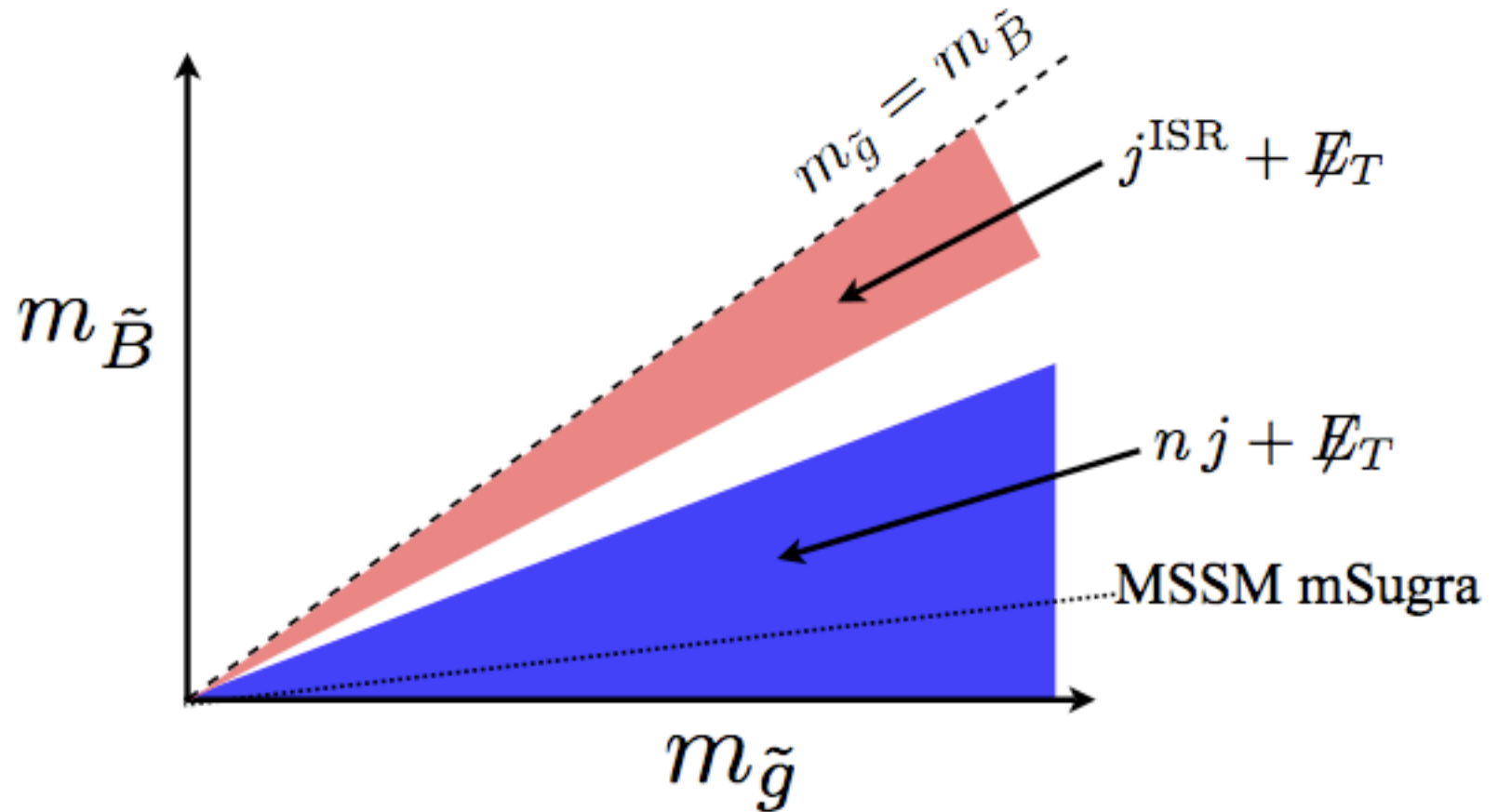
Take  $M_{\tilde{B}} = 60 \text{ GeV} \dots$

# 70 GeV Gluinos!??

DØ Preliminary,  $0.96 \text{ fb}^{-1}$



# Motivates a Gluino-Bino Plot



Wacker, et. al.

# Ratios of 'inos'

$$M_3 > M_1 > M_2$$

AMSB - cascades with leptons.  
LL charged particle

$$M_2 > M_3 > M_1$$

Only jets + missing  $E_T$

$$M_1 > M_3 > M_2$$

Jets plus missing  $E_T$  and/or LL  
charged particle

$$M_2 > M_1 > M_3$$

Glino may be long lived - at  
least 4-body decay

# Violation of R Parity

$$W = H_1 Q D^c + H_2 Q U^c + H_1 L E^c + \mu H_1 H_2 \\ + L Q D^c + U^c D^c D^c + L L E^c + \mu_L L H_2$$

# Violation of R Parity

$$W = H_1 Q D^c + H_2 Q U^c + H_1 L E^c + \mu H_1 H_2 \\ + L Q D^c + \cancel{U^c D^c D^c} + L L E^c + \mu_L L H_2$$

Could have “Baryon Parity”.



# Violation of R Parity

$$W = H_1 Q D^c + H_2 Q U^c + H_1 L E^c + \mu H_1 H_2$$

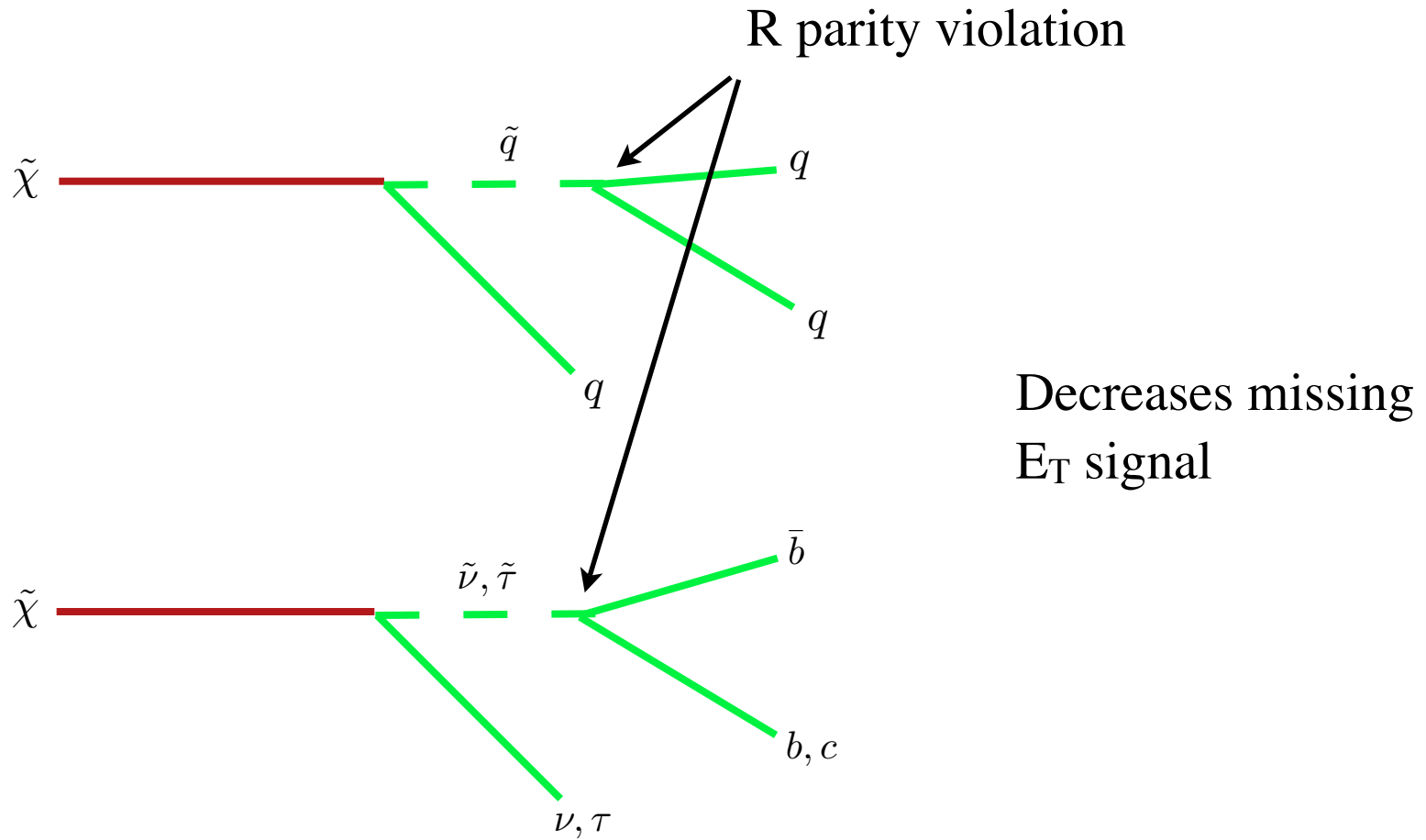
~~$+ L Q D^c + U^c D^c D^c + L L E^c + \mu_L L H_2$~~

or “Lepton Parity”

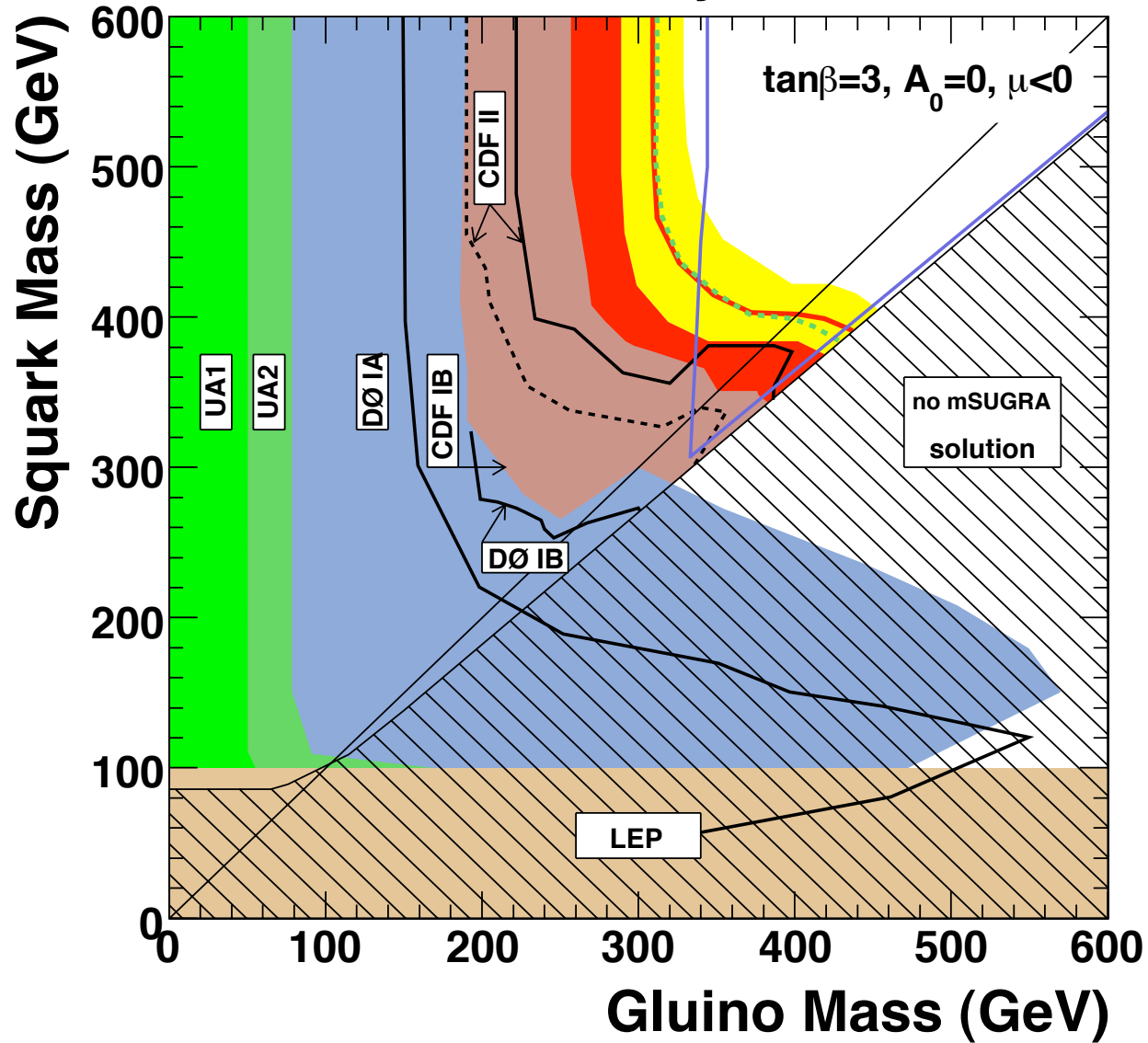
$$p \not\rightarrow X$$

$X$  must have an odd number of fermions,  
thus its lepton number  $L = 2n + 1$ . No lepton  
parity violation, no decay.

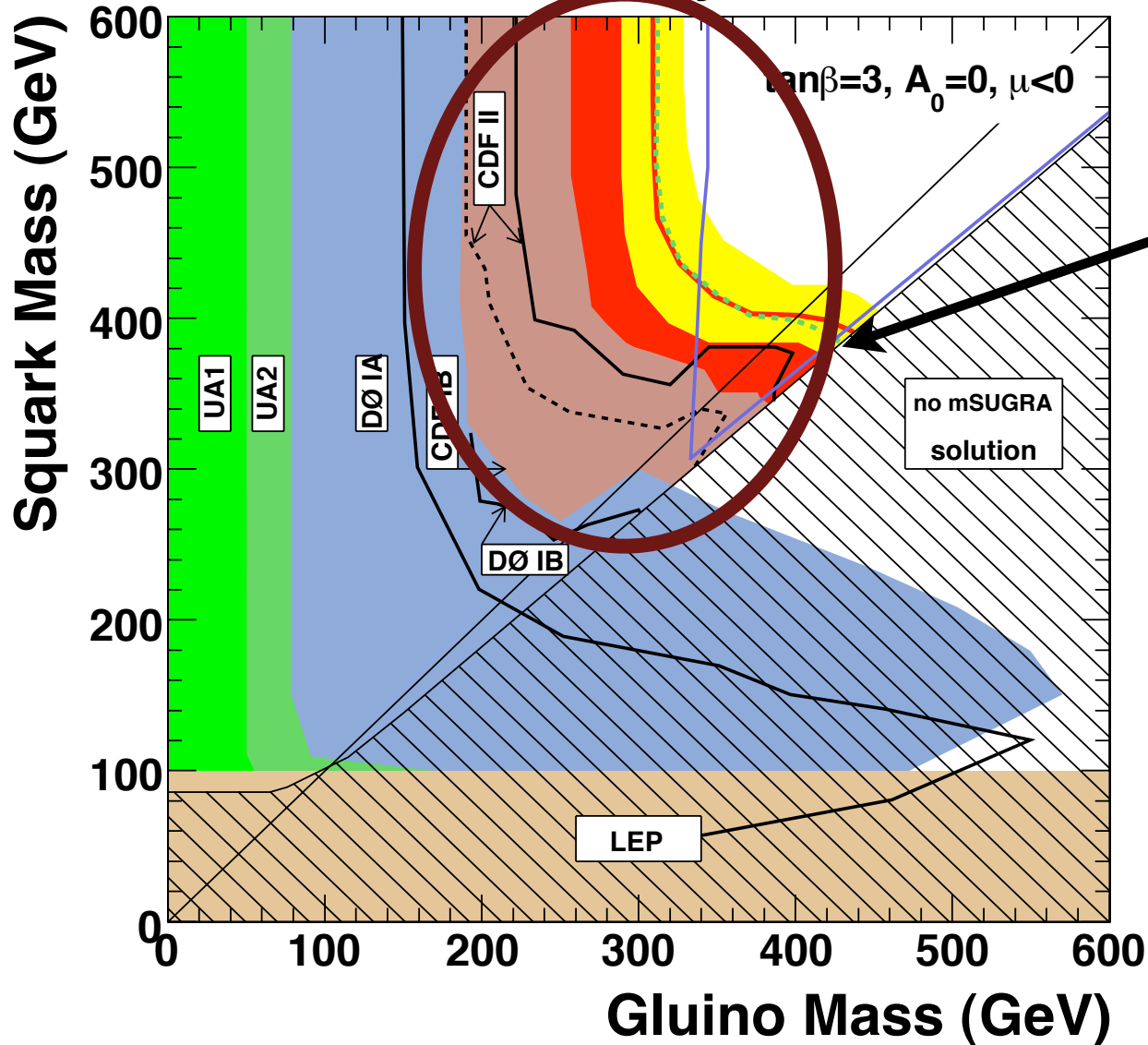
# LSP decays



# DØ Preliminary, 0.96 fb<sup>-1</sup>



# DØ Preliminary, 0.96 fb<sup>-1</sup>



Required  
missing  $E_T$   
> 100 GeV

no mSUGRA  
solution

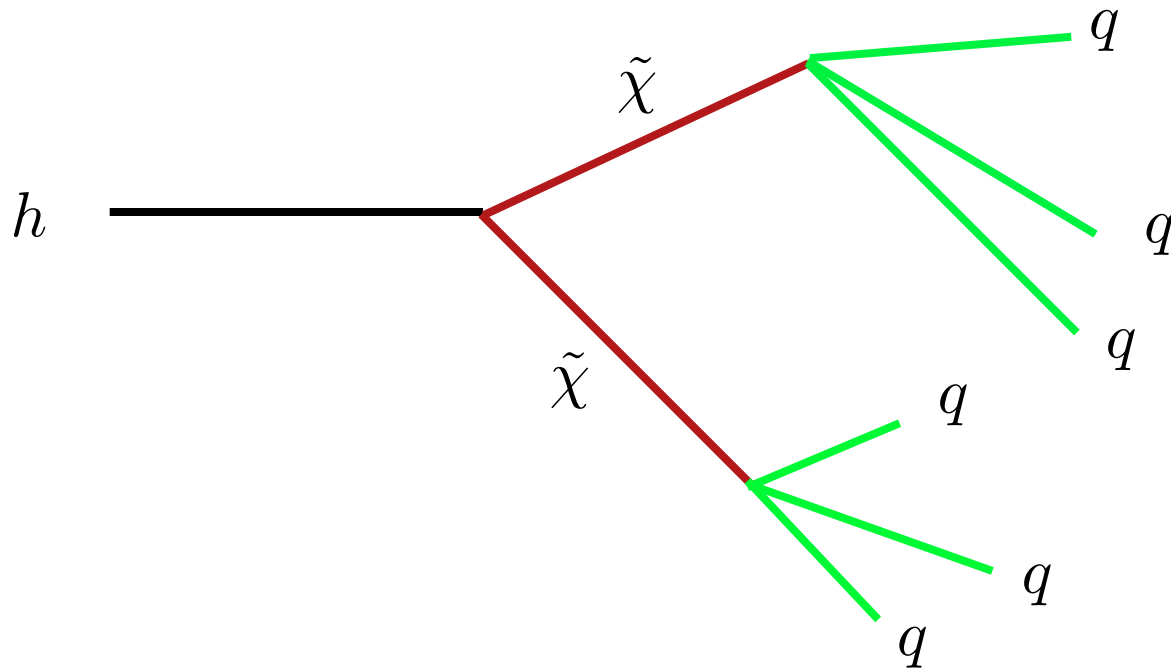
LEP

# Typical Bounds on Superpartners with B Violation

Sleptons (R)	94,85,70 GeV (A)
Sneutrinos	88,65,65 GeV (A)
Squarks ( $u_{L/R}, d_{L/R}$ )	87,80,86,56 GeV (L)
Stop	77 GeV (O,D,L)
Sbottom	7.5 (>55, <30) GeV (L)
Gluino	80 GeV ? (UA2)

Only Chargino bound roughly the same (102.5 GeV)

# Higgs decays with B-violation



higgs  $\rightarrow$  6 jets!

$m_h > 87 \text{ GeV}$

Could be a source of new displaced vertices.

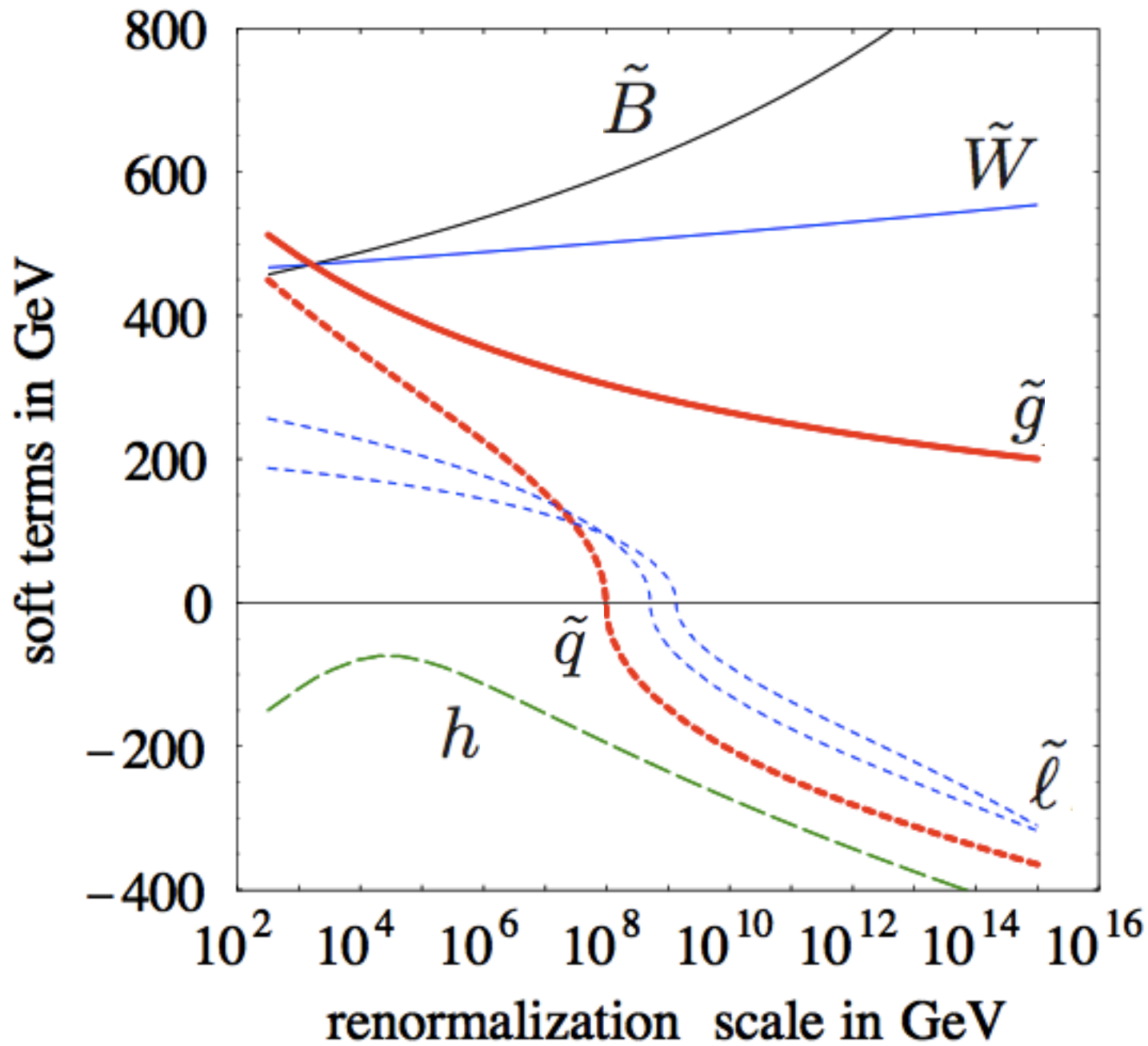
# Supersymmetry Breaking with 'Squashed' Spectrum

Deflected Anomaly Mediation

Mirage Mediation

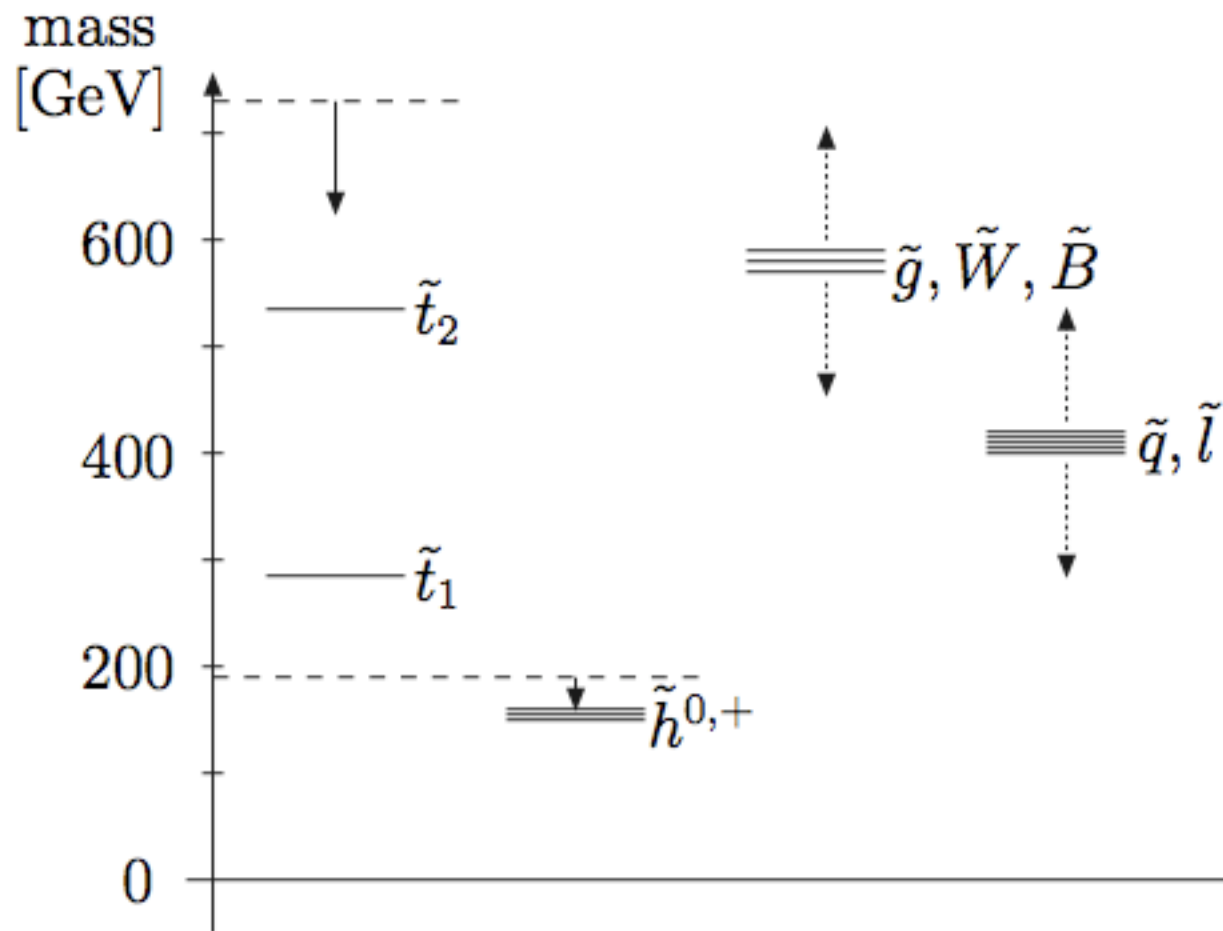
Dirac Gauginos

# DAMSB





# Mirage Mediation



# Supersymmetry Breaking with 'Squashed' Spectrum

Deflected Anomaly Mediation

Mirage Mediation

Dirac Gauginos vs. Majorana Gauginos

# "New" Parameter Set At the Weak Scale

$$\begin{aligned}
 \mathcal{L}_{\text{soft}}^{\text{MSSM}} = & -\frac{1}{2} (M_3 \tilde{g}\tilde{g} + M_2 \tilde{W}\tilde{W} + M_1 \tilde{B}\tilde{B} + \text{c.c.}) \\
 & - (\tilde{u} \mathbf{a}_u \tilde{Q} H_u - \tilde{d} \mathbf{a}_d \tilde{Q} H_d - \tilde{e} \mathbf{a}_e \tilde{L} H_d + \text{c.c.}) \\
 & - \tilde{Q}^\dagger \mathbf{m}_Q^2 \tilde{Q} - \tilde{L}^\dagger \mathbf{m}_L^2 \tilde{L} - \tilde{u} \mathbf{m}_u^2 \tilde{u}^\dagger - \tilde{d} \mathbf{m}_d^2 \tilde{d}^\dagger - \tilde{e} \mathbf{m}_e^2 \tilde{e}^\dagger \\
 & - m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (\mu H_u H_d + \text{c.c.}) .
 \end{aligned}$$

parameter x  
identity →  
matrix

Use this  
to fix the  
Z mass.

$\mu$  sets higgsino mass

$\tan \beta$

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 & - \tilde{Q}^\dagger \mathbf{m}_Q^2 \tilde{Q} - \tilde{L}^\dagger \mathbf{m}_L^2 \tilde{L} - \tilde{u} \mathbf{m}_u^2 \tilde{u}^\dagger - \tilde{d} \mathbf{m}_d^2 \tilde{d}^\dagger - \tilde{e} \mathbf{m}_e^2 \tilde{e}^\dagger \\
 & - m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (\mu H_u H_d + \text{c.c.}).
 \end{aligned}$$

$$\mathbf{a}_u = \mathbf{y}_u A_u$$

parameter x  
identity  $\rightarrow$   
matrix

Use this  
to fix the  
Z mass.

$\tan \beta$

$\mu$  sets higgsino mass

# Should the world be "Natural"?

$$(m_h^{phys})^2 = M_{planck}^2 + M_{qtm\ corr}^2$$

Cancellation of one part in  $10^{34}$

This doesn't happen in condensed matter system unless we force it to happen.

# The Cosmological Constant

Fine tuning of one part in  $10^{120}$ !

(one part in  $10^{60}$  with SUSY)

In 1987 Weinberg suggested that if our universe was one of a large selection, we would live in one where the cosmological constant was *just small enough* to allow for structure to form, and if we don't measure one, we can rule out the anthropic principle.

# The Cosmological Constant

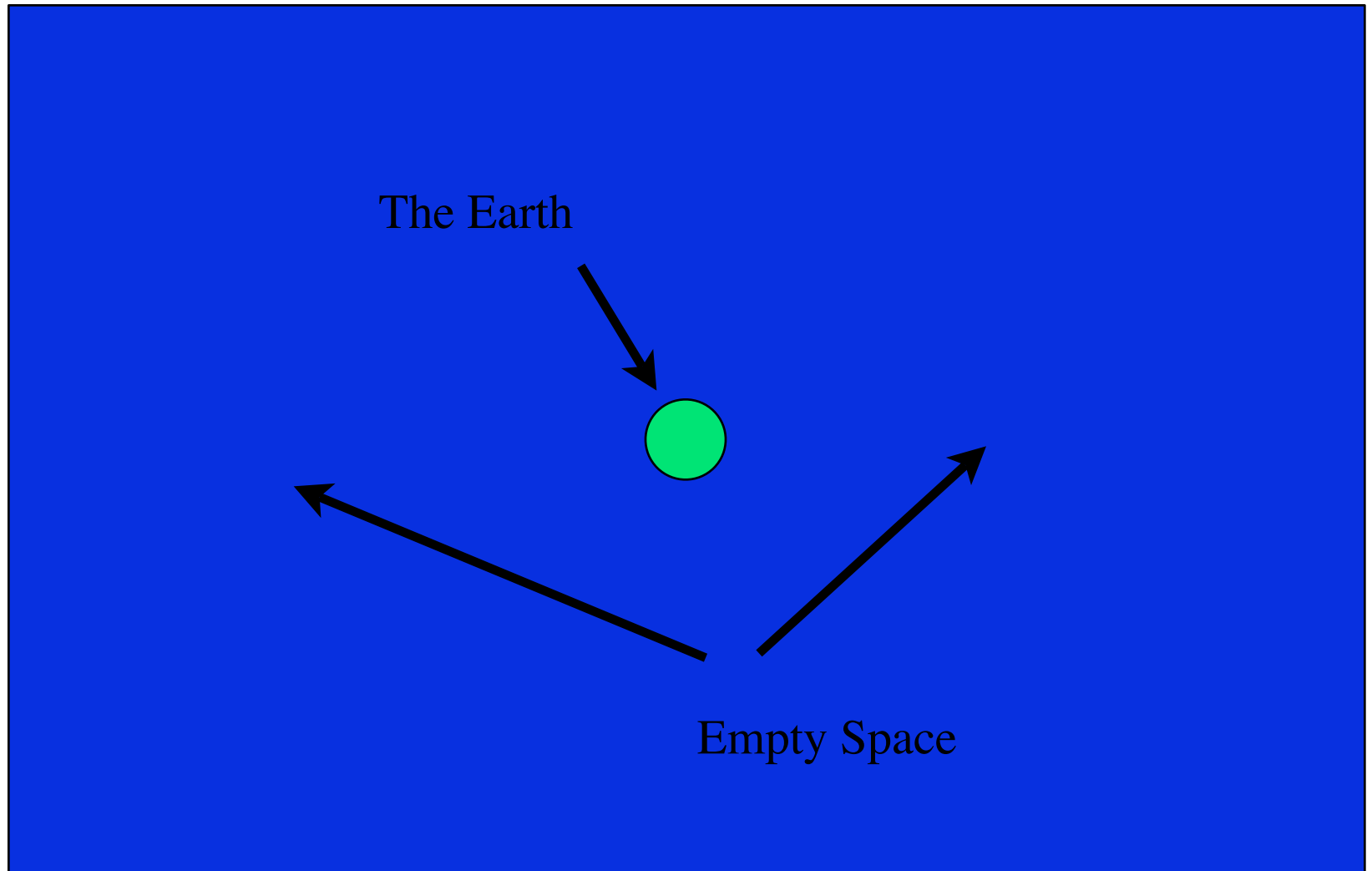
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But then we did measure one...

# Environmental Selection

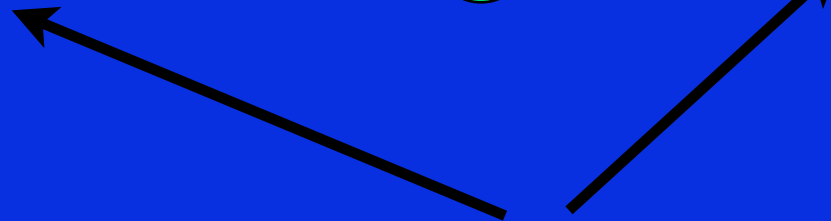
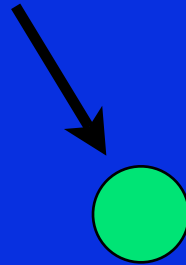




# Environmental Selection

Fine tuning of one part in  $10^{57}$ ?

The Earth



Empty Space

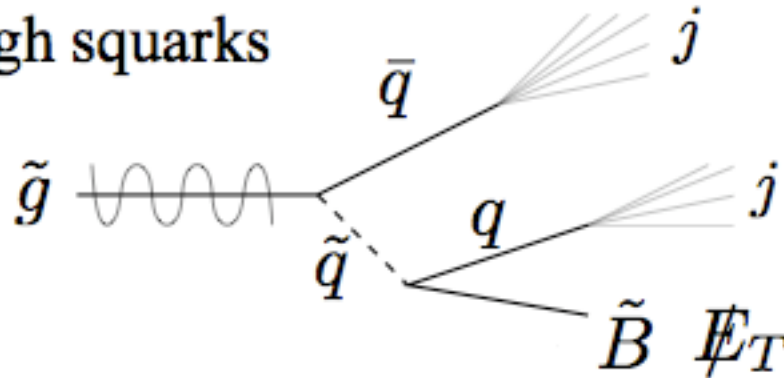
# Split Supersymmetry

	$M_{\text{Pl.}}$	$10^{16}$ TeV
Scalars (Squarks, sleptons, ...)	$M_{\text{susy}}$ ? {	$10^{15}$ TeV 10 TeV
Fermions (Higgsinos, gauginos) +SM Higgs	$M_{\text{weak}}$	$\sim 1$ TeV
	$M_{\text{CC}}$	$10^{-15}$ TeV

*Preserves Unification  
and Dark Matter*

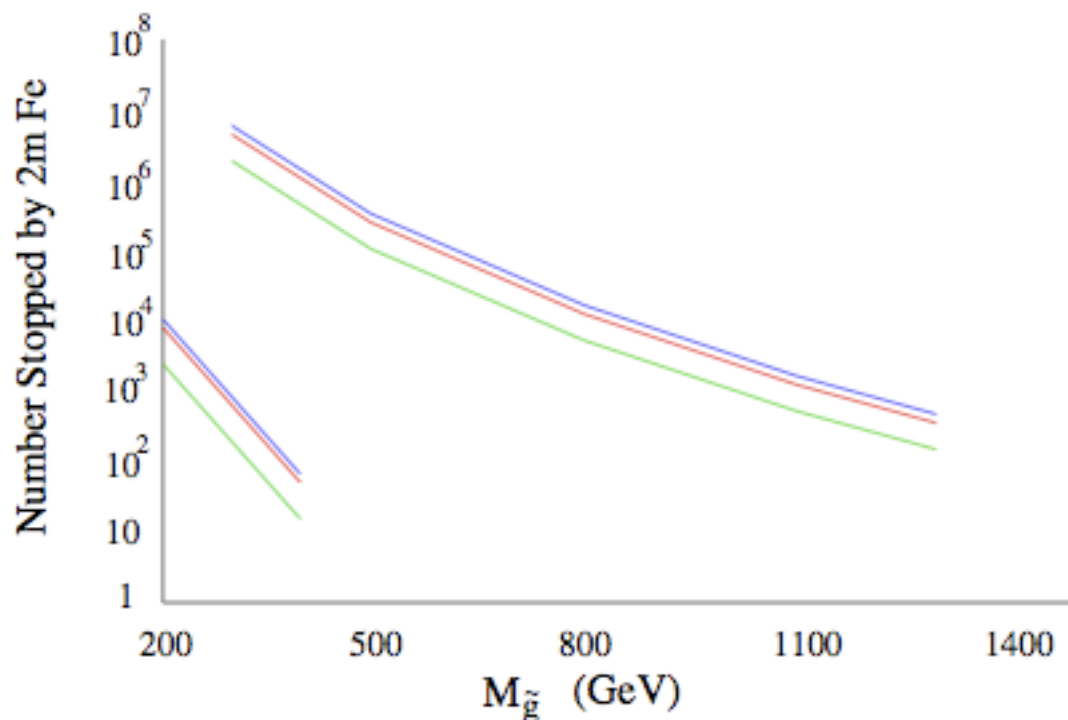
# Gluinos are Long-Lived

Must decay through squarks



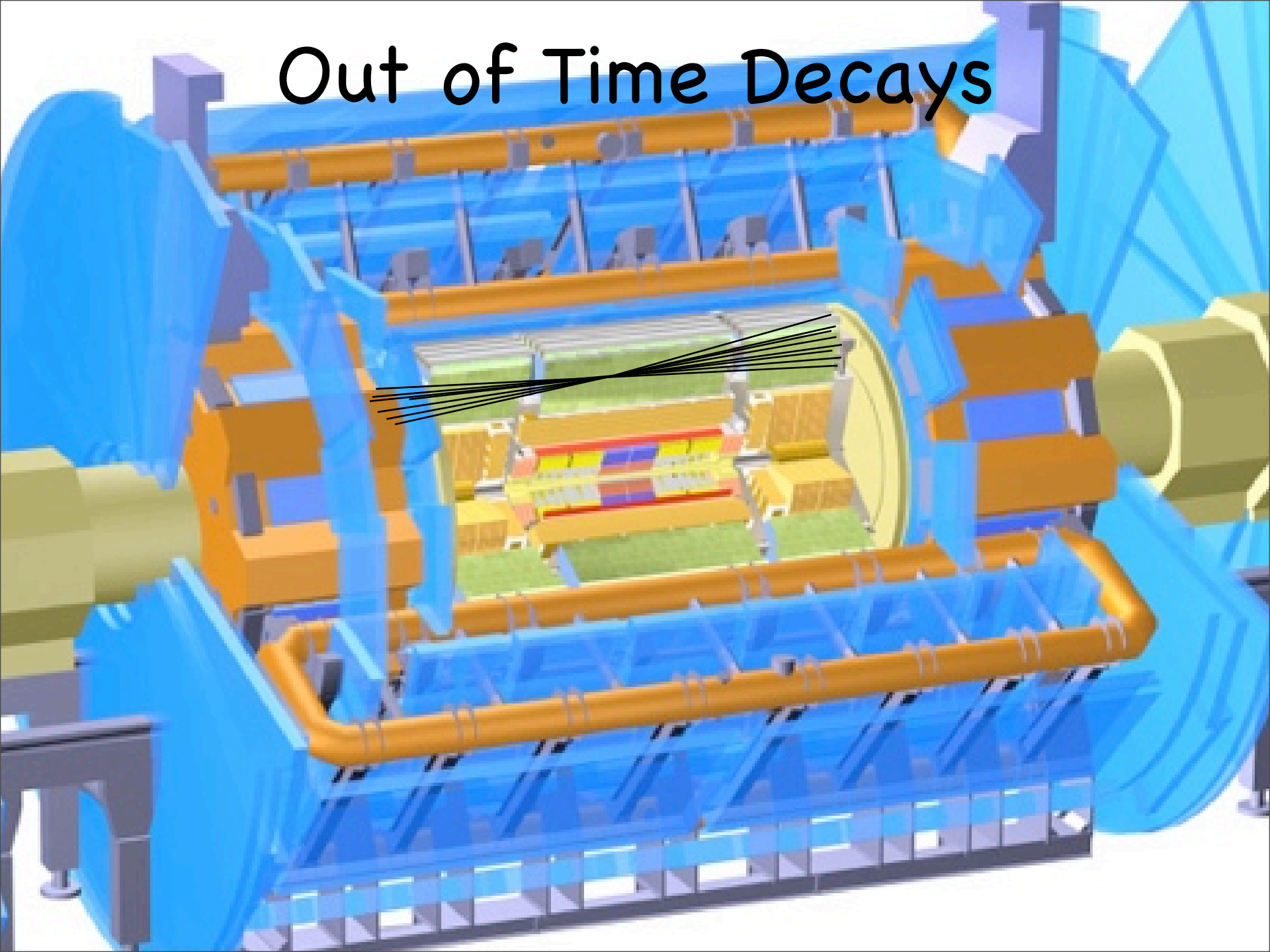
$$\tau_{\tilde{g}} \simeq 2 \text{ sec.} \left( \frac{350 \text{ GeV}}{m_{\tilde{g}}} \right)^5 \left( \frac{M_{\text{Susy}}}{10^6 \text{ TeV}} \right)^4$$

# Stopped Gluinos



$2 \text{ fb}^{-1}$	200 GeV	300 GeV	400 GeV
CDF	$4.1 \times 10^3$	$3.1 \times 10^2$	$3.3 \times 10^1$
D0	$4.5 \times 10^3$	$3.3 \times 10^2$	$3.4 \times 10^1$
$100 \text{ fb}^{-1}$	300 GeV	800 GeV	1300 GeV
ATLAS	$5.8 \times 10^6$	$1.8 \times 10^4$	$6.2 \times 10^2$
CMS	$3.7 \times 10^6$	$1.2 \times 10^4$	$3.9 \times 10^2$

# Out of Time Decays



Even if you HATE use of the  
anthropic principle...

Is pointed us to a new signal to look for -  
very long lived stopped particles.