

# MMHT2014 PDFs and connection to Xfitter

Robert Thorne

March 20th, 2016



University College London

In collaboration with Lucian Harland-Lang, Ricky Nathvani and Alan Martin

and thanks to Patrick Motylinski, Ben Watt, Graeme Watt and James Stirling

I will cover a number of topics.

– The impact of fitting new **LHC** and **Tevatron** data. → clear improvements in some PDF uncertainties.

Specific issues with **ATLAS 7 TeV** jet data. Some discussion of **NNLO** effects.

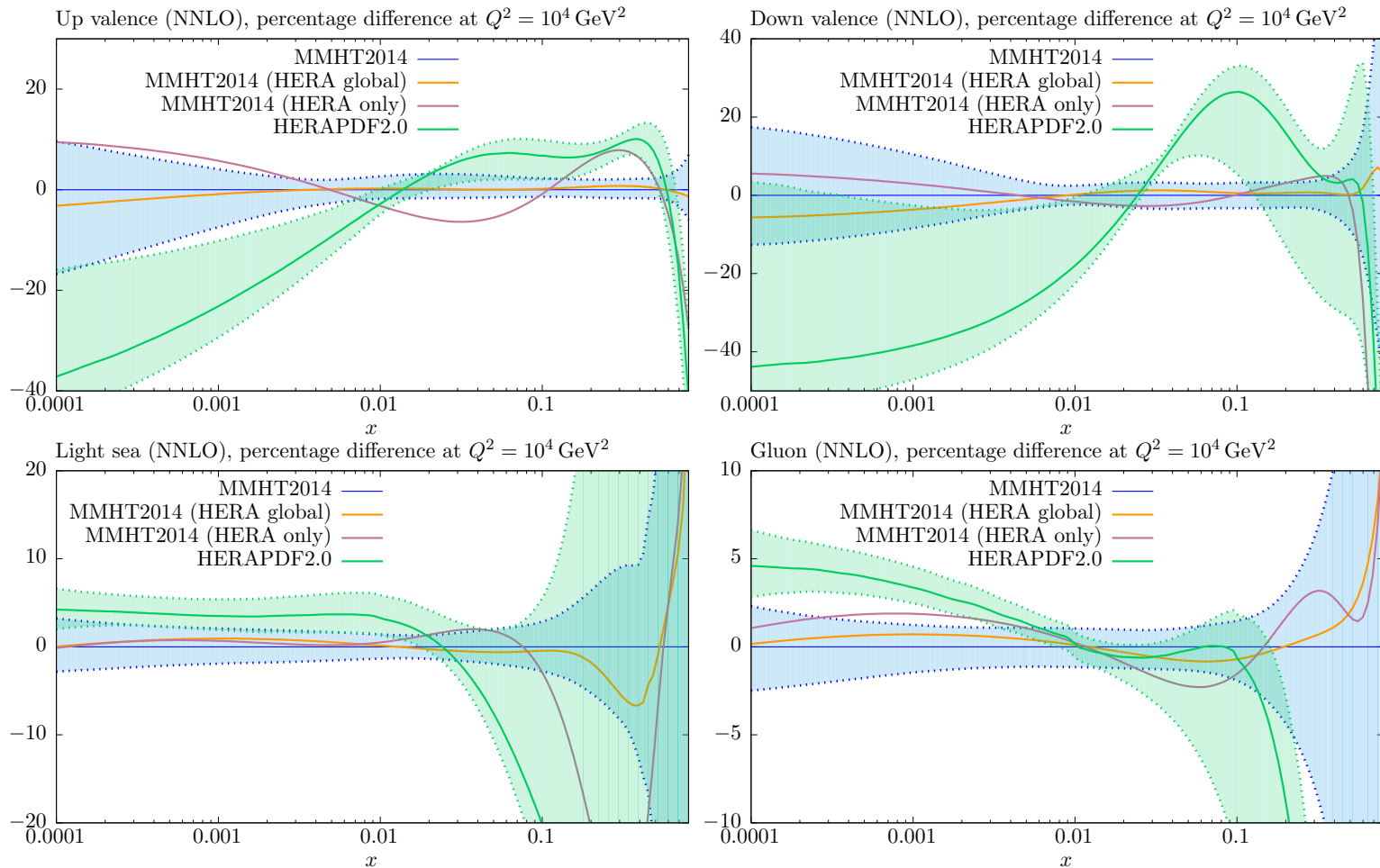
Preliminary results with inclusion of final **ATLAS 7 TeV  $W, Z$**  rapidity dependent data.

A brief intro to topics just starting investigation. Extension of parameterisation and **QED** corrections.

Connection of **MSTW/MMHT** to **xFitter**

# HERA II Combined data

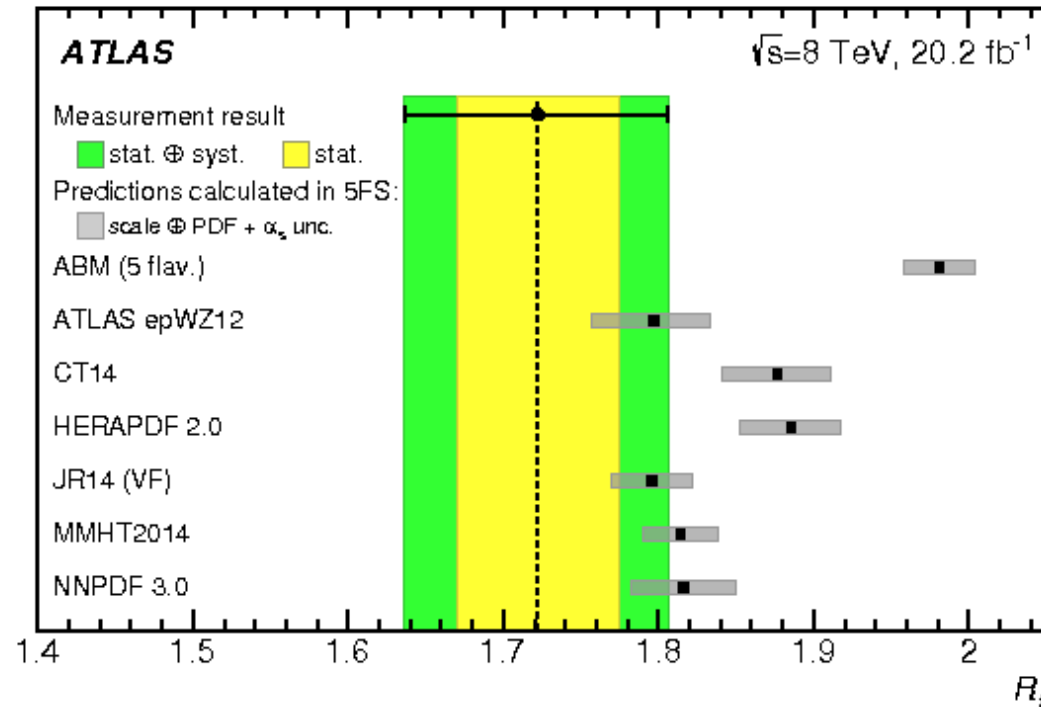
Fit quality in global fit at NNLO  $Q^2 > 2.5 \sim 1435/1168$



HERA II modified PDFs very well within MMHT2014 uncertainties. PDFs from HERA II data only fit in some ways similar to HERAPDF2.0.

Up to 10% reduction in uncertainties. Small changes in central values.

Note that highish- $x$  enhancement in up quark preferred by HERA  $e^-$  charged-current data in tension with recent most accurate measurement of single top ratio.



Also disfavors any other reason for enhanced  $u(x)/d(x)$  for  $x \sim 0.1$ .

## Breakdown of fit quality to new hadron collider data

We now also fit to high rapidity  $W, Z$  data from LHCb at 7 and 8 TeV,  $W + c$  jets from CMS, which constrains strange quarks, high precision CMS data on  $W^{+,-}$  rapidity distributions which can also be interpreted as an asymmetry measurement, and also the final  $e$  asymmetry data from D0 (lepton, not  $W$  asymmetry).

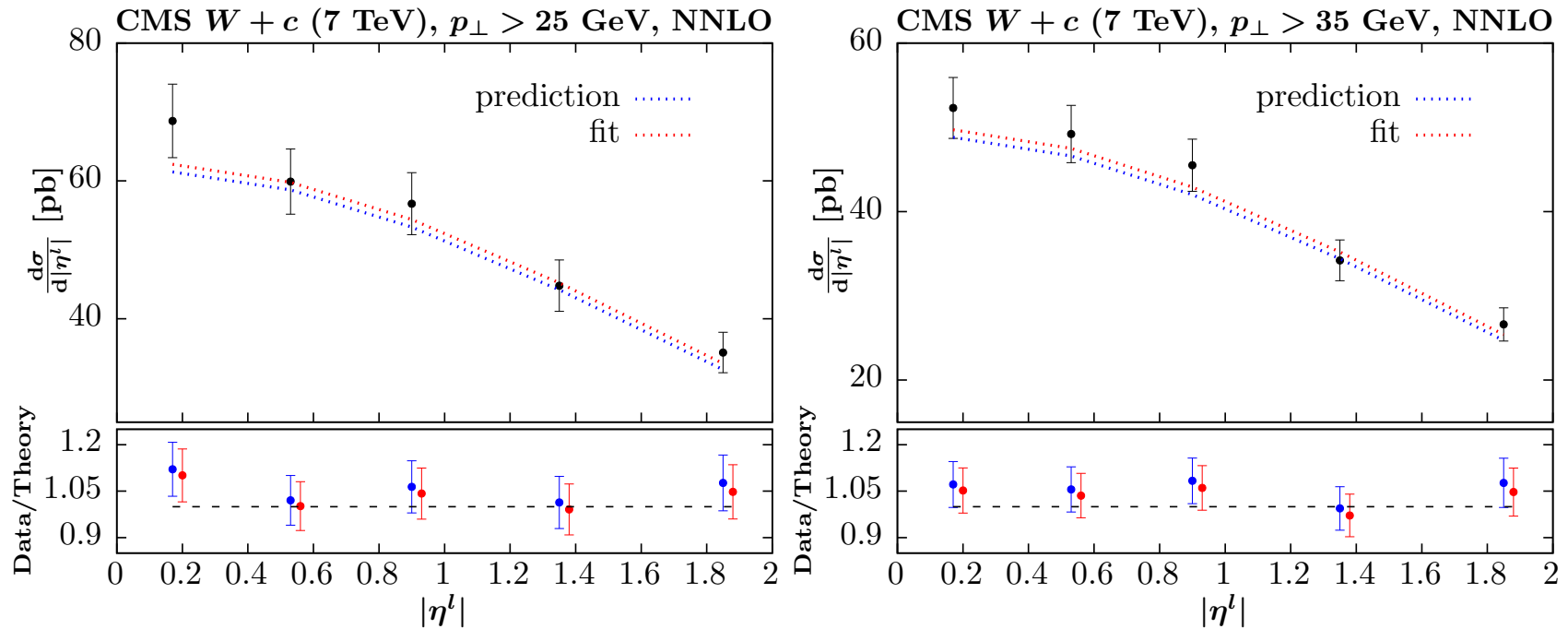
	no. points	NLO $\chi_{pred}^2$	NLO $\chi_{new}^2$	NNLO $\chi_{pred}^2$	NNLO $\chi_{new}^2$
$\sigma_{t\bar{t}}$ Tevatron +CMS+ATLAS	18	19.6	20.5	14.7	15.5
LHCb 7 TeV $W + Z$	33	50.1	45.4	46.5	42.9
LHCb 8 TeV $W + Z$	34	77.0	58.9	62.6	59.0
LHCb 8TeV $e$	17	37.4	33.4	30.3	28.9
CMS 8 TeV $W$	22	32.6	18.6	34.9	20.5
CMS 7 TeV $W + c$	10	8.5	10.0	8.7	8.0
D0 $e$ asymmetry	13	22.2	21.5	27.3	25.8
total	3738/3405	4375.9	4336.1	3741.5	3723.7

Predictions good, and no real tension with other data when refitting, i.e. changes in PDFs relatively small. Slightly ( $\sim 10$  units) better than previous report due to improvements (and one correction) in K-factors.

At NLO  $\Delta\chi^2 = 9$  for the remainder of the data and at NNLO  $\Delta\chi^2 = 8$ .

When couplings left free at NLO  $\alpha_S(M_Z^2)$  stays very close to 0.120 but at NNLO  $\alpha_S(M_Z^2)$  marginally above 0.118, higher than MMHT2014.

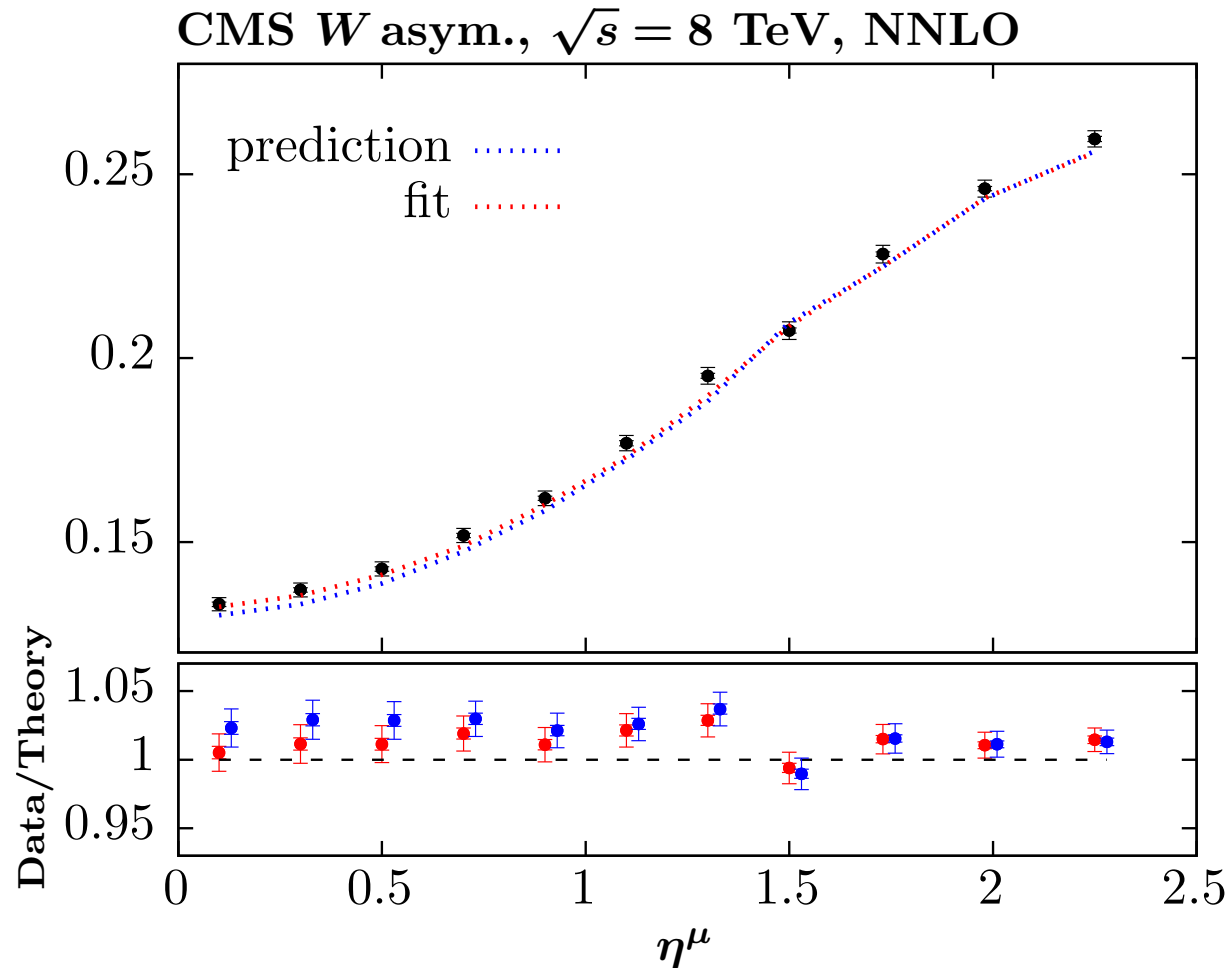
# New data sets for fit – $W + c$ differential distributions.



Data on plot use uncertainties added in quadrature.

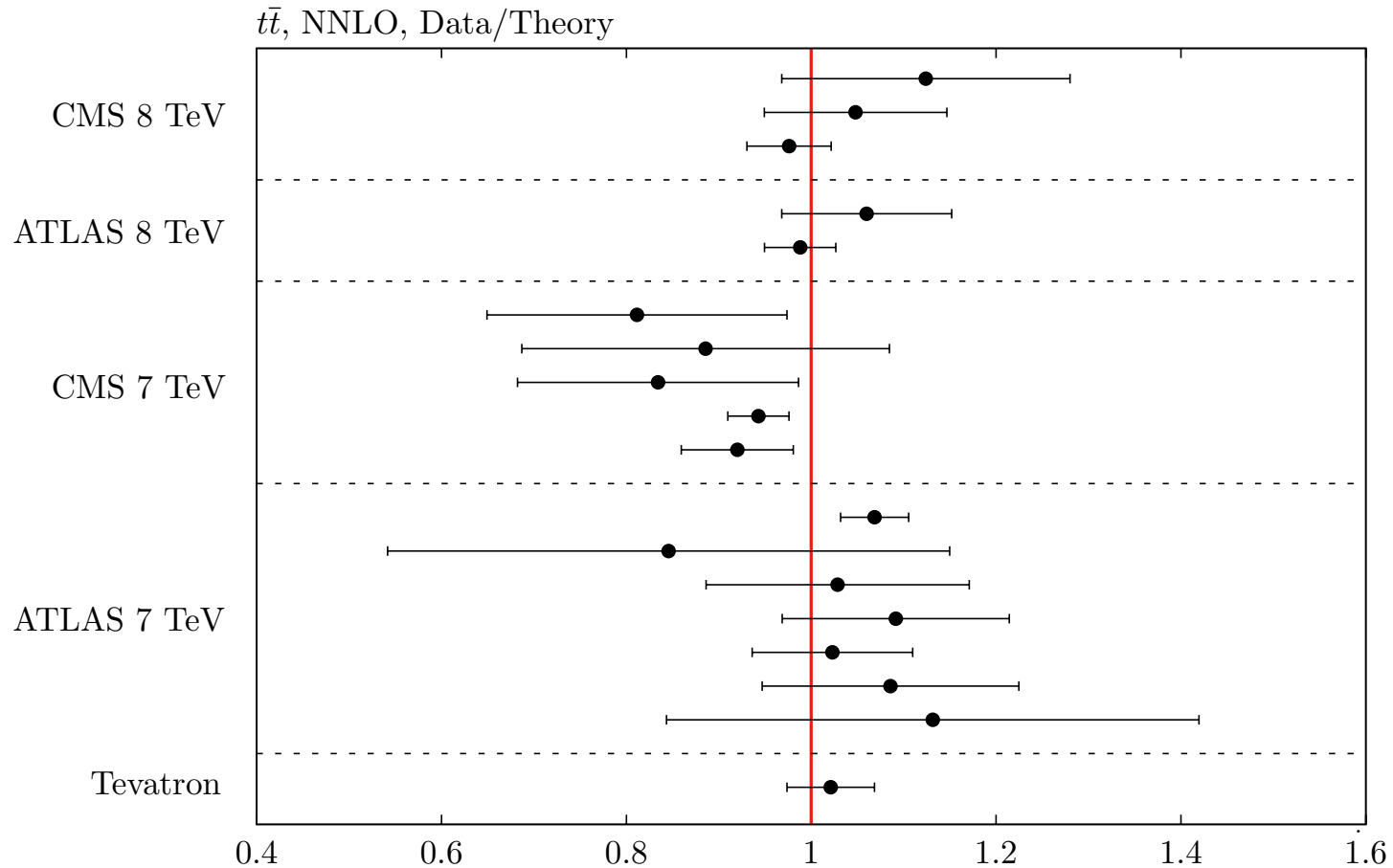
Very little change after fit. By eye comparison looks worse, but slightly better when covariance matrix used (as in fit).

**Good agreement with new 8 TeV CMS  $W^\pm$  rapidity** and asymmetry data (shown). (Fit to individual distributions not asymmetry.)



Small- $x$  valence quarks require some modification of order the size of uncertainty. Scope for reduced uncertainty with new data inclusion.

# Included some more up-to-date results on $\sigma_{t\bar{t}}$ .



Fit very good and with  $\alpha_S(M_Z^2) = 0.118$  the fitted  $m_t^{pole} = 173.4$  GeV.  
 At NLO  $m_t^{pole} = 170.2$  GeV. MMHT values  $m_t^{pole} = 174.2$  GeV and  $m_t^{pole} = 171.7$  GeV

Helps drive slight increase in  $\alpha_S(M_Z^2)$



## PDF sets generated

We generate a preliminary (not for distribution) central set at **NLO** and **NNLO** for fit to new data – labelled **MMHT (2016 fit)**.

Also generate PDF eigenvector sets for uncertainties at **NNLO**.

Use same basis of **25** free PDF parameters as in **MMHT2014** (this is subject to possible change in the future).

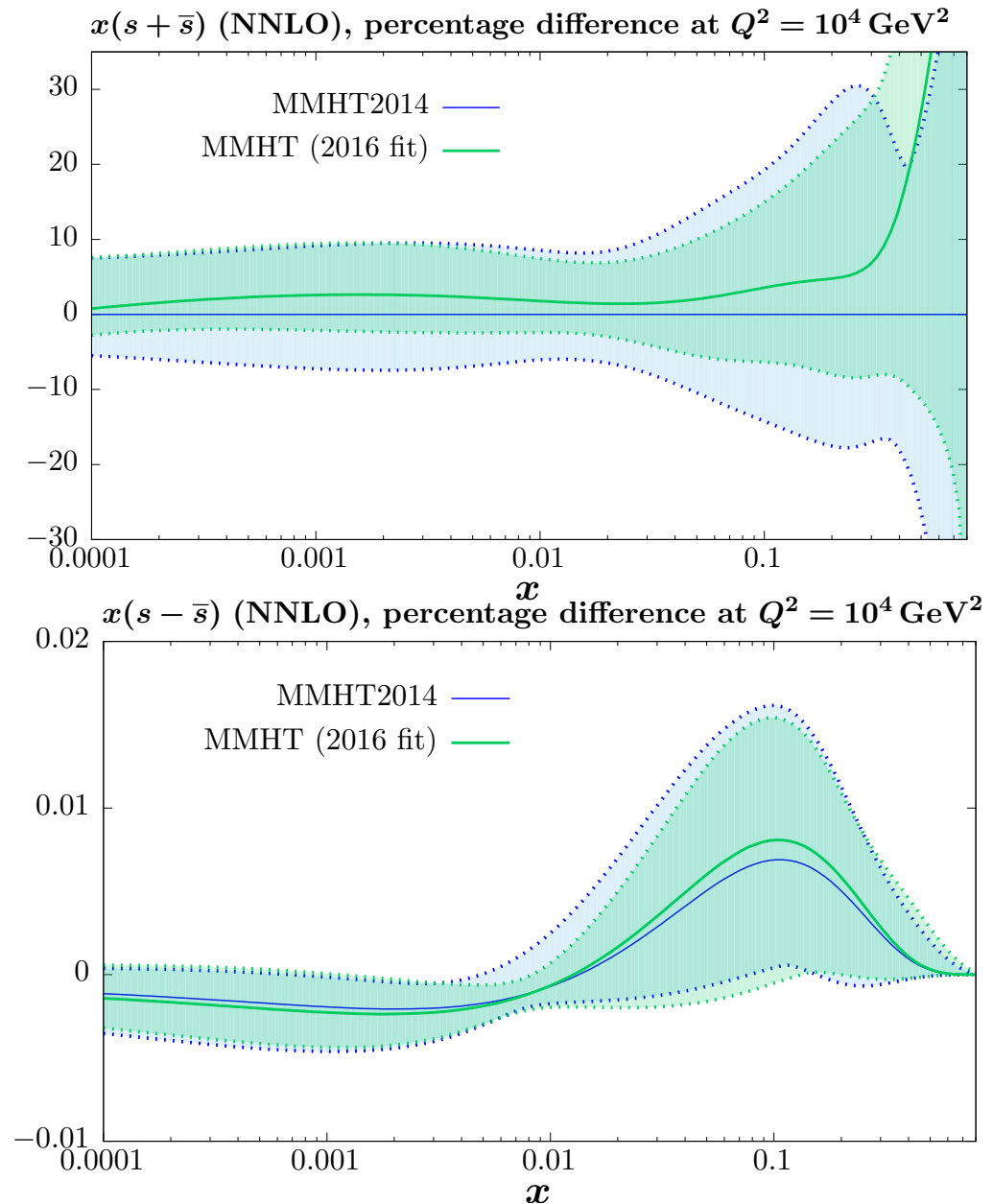
Hence, **50** eigenvector directions.

**14** of these are best constrained by one of the new (**LHC**) data sets, **CMS 8 TeV  $W$**  data and  **$W + c$  jets** and the new **LHCb** data.

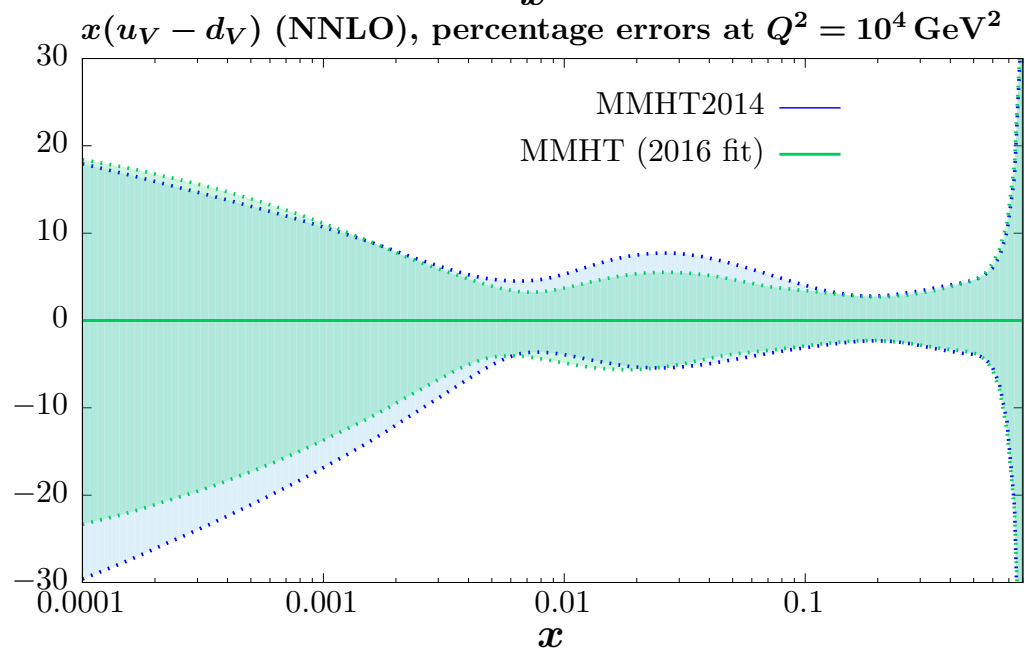
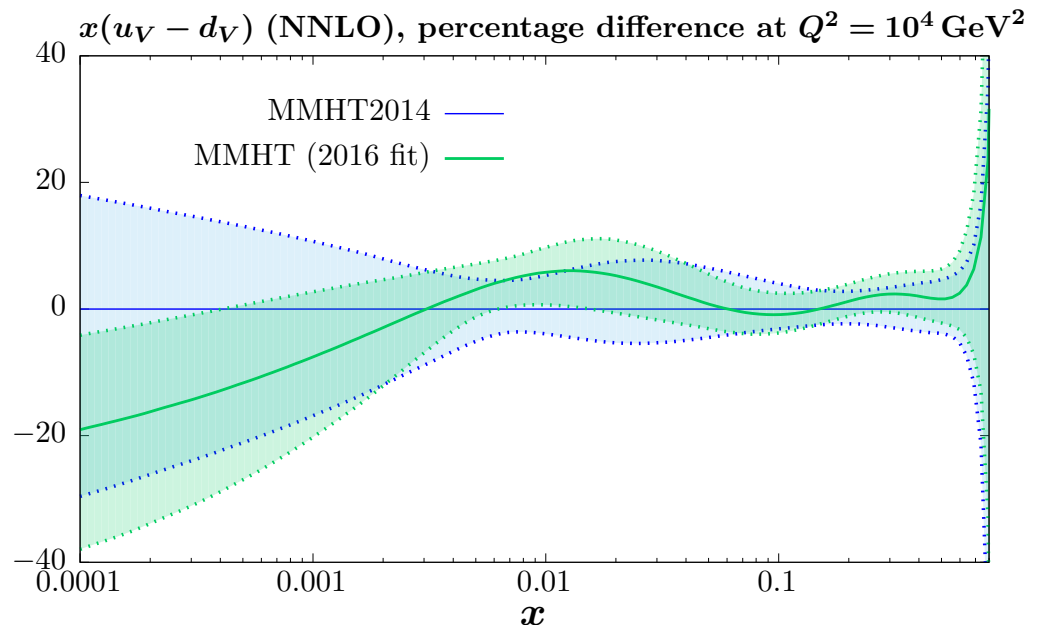
## Effect on PDFs

Large reduction in the  $s + \bar{s}$  uncertainty, but little change in central value. Due to  $W + c$  jets data.

There is some impact on the  $s - \bar{s}$  uncertainty, from (effective) asymmetry data.



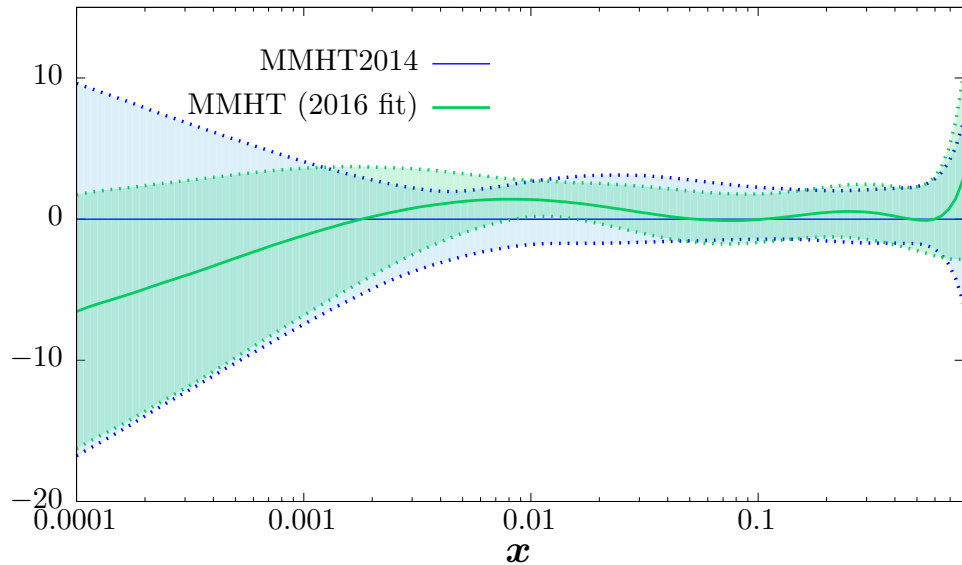
A significant change in  $u_v - d_v$ , and reduction in the uncertainty, from (effective) CMS asymmetry data.



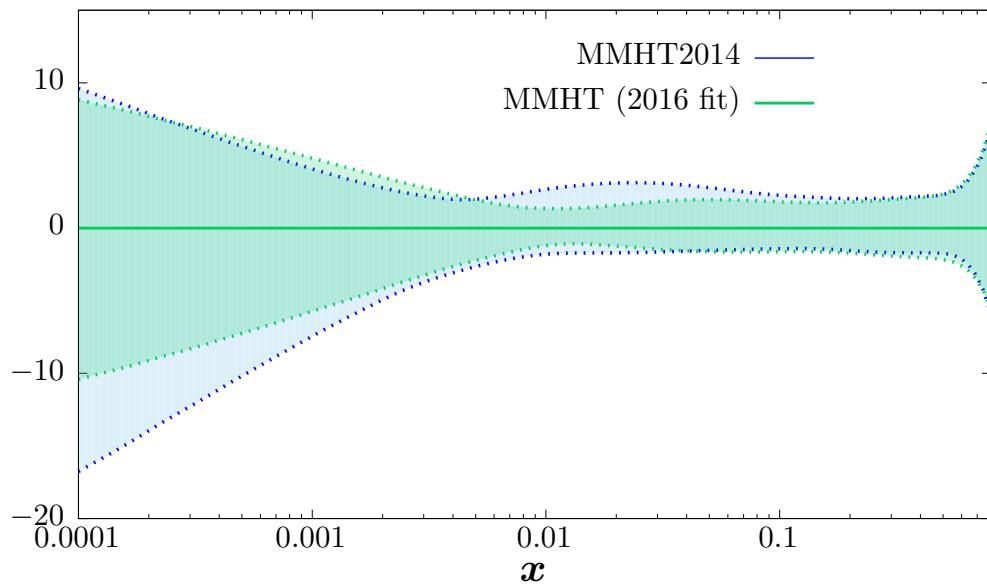
Main change and uncertainty reduction in  $u_v$  rather than  $d_v$ .

Mainly **CMS** data, but some impact of new **HERA** data.

Up valence (NNLO), percentage difference at  $Q^2 = 10^4 \text{ GeV}^2$



Up valence (NNLO), percentage errors at  $Q^2 = 10^4 \text{ GeV}^2$



## Attempted fit to high luminosity **ATLAS 7 TeV** inclusive jet data (**JHEP 02 (2015) 153**)

Take as default  $R = 0.4$  and  $\mu = p_{T,1}$  and work at **NLO**.

Prediction at **NLO** gives  $\chi^2/N_{pts} = 413.1/140$ .

Refit gives improvement only to  $\chi^2/N_{pts} = 400.4/140$ .

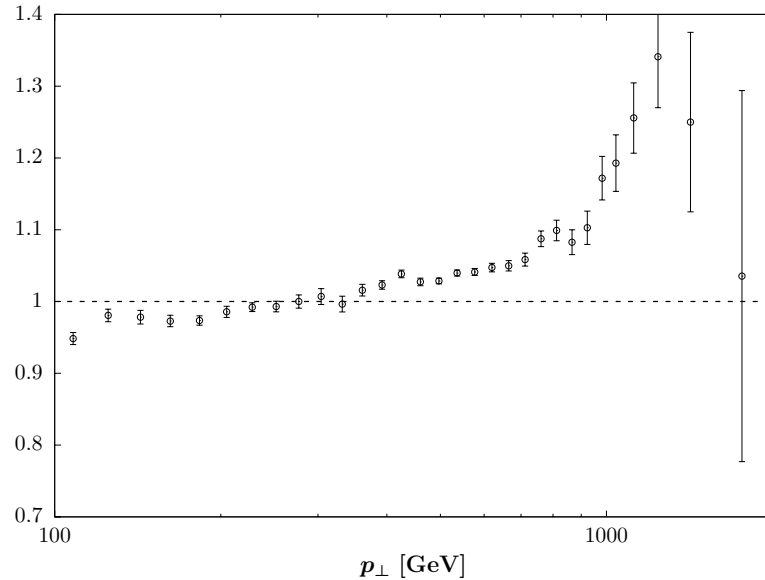
Deterioration in other data only  $\Delta\chi^2 \sim 3$ , so failure not due to strong tensions.

Cannot simultaneously fit data in all bins. Mismatch in one rapidity bin different in form to neighbouring bins probing PDFs of similar flavour,  $x$  and  $Q^2$ .

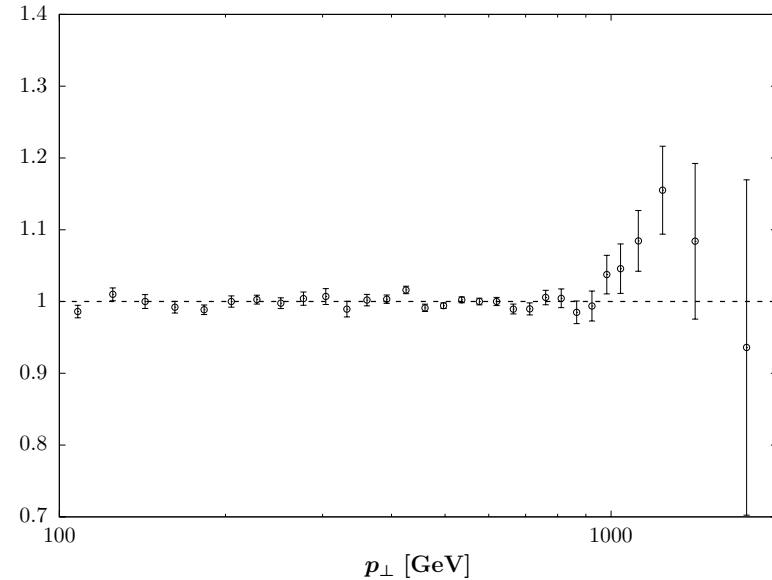
Similar results also seen by other groups.

# Large correlated systematics. Dominate over uncorrelated uncertainties.

Data/Theory,  $0.0 < |y| < 0.5$ , No shift, stat. errors only



Data/Theory,  $0.0 < |y| < 0.5$ , Shift, stat. errors only

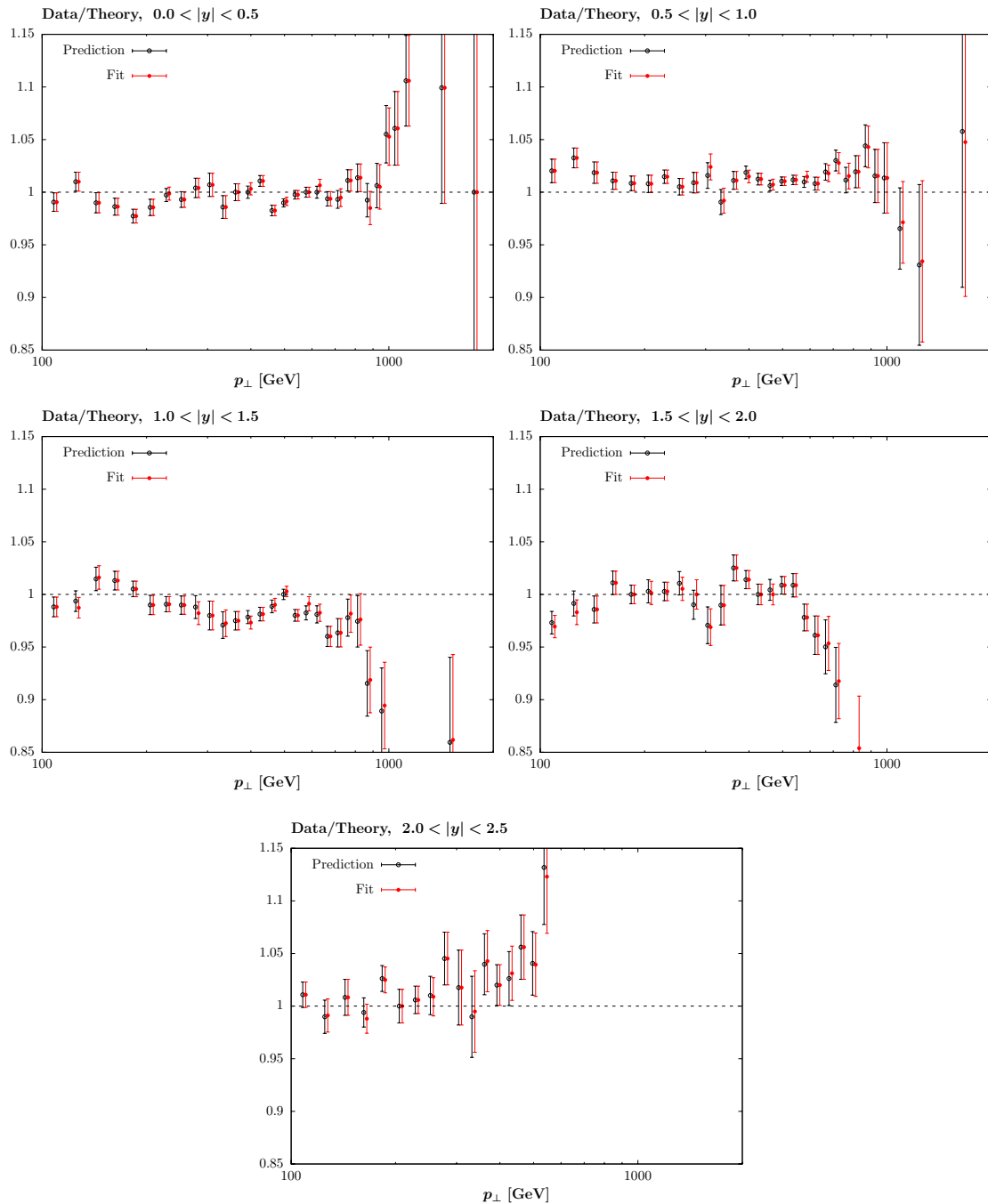


For “weak” assumption about correlations there are **71** correlated systematic uncertainties, 65 related to jet energy scale.

Best fit requires shift in data against theory. Highly correlated from one rapidity bin to another.

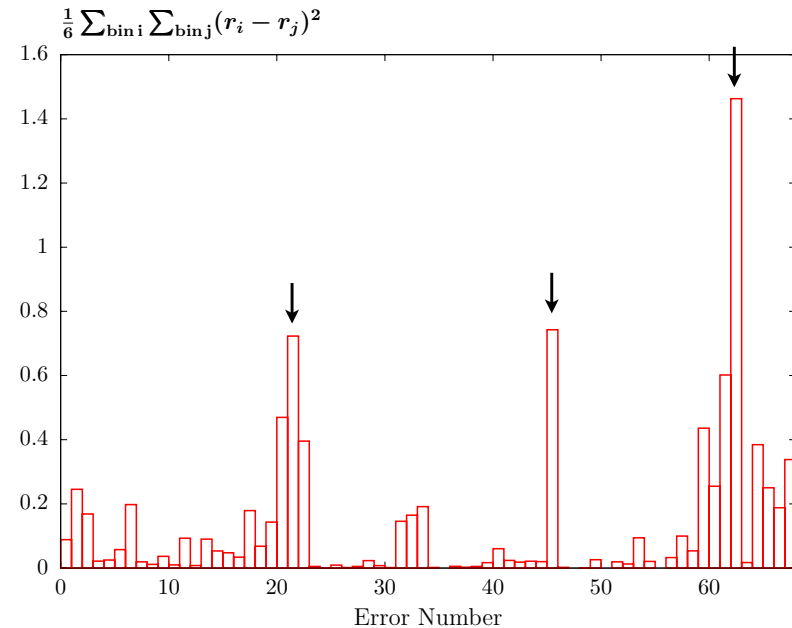
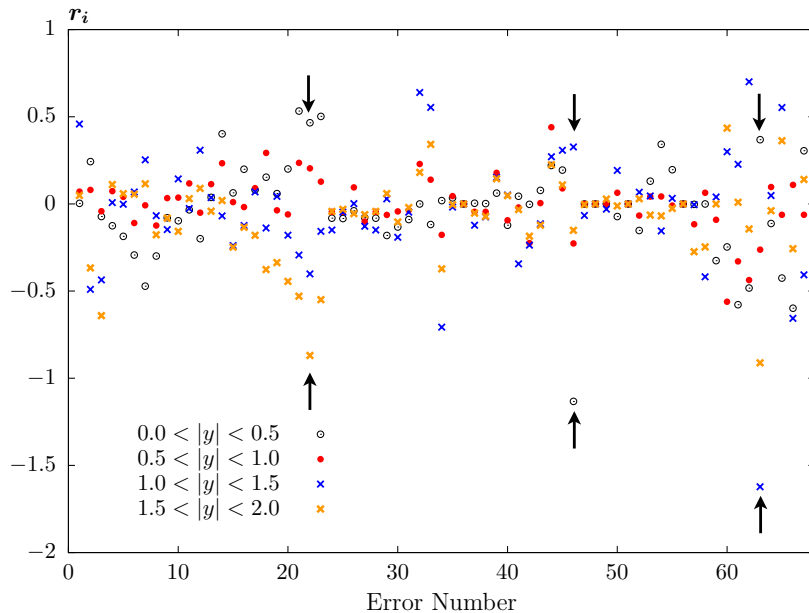
Cannot simultaneously fit data in all bins. Mismatch in one rapidity bin different to others probing PDFs of similar flavour,  $x$  and  $Q^2$ .

Similar results also seen by other groups.



Good fit ( $\chi^2/N_{pts} \sim 1$ ) possible when fitting all individual rapidity bins.

Look at shifts due to each source of correlated uncertainty, i.e. preferred  $r_k$  when fitting each separate rapidity bin.



A small number of sources prefer very different values when fits to different bins performed.

These are [jes21](#), [jes45](#) and [jes62](#) (Multi-jet balance asym., JES pile-up and JES close by jets [1406.0076](#)).



## Exercise on decorrelating uncertainties

We consider the effect on the  $\chi^2$  of the simultaneous fit to all data of decorrelating these three error sources, i.e. making them independent between the 6 rapidity bins.

Compared to the original  $\chi^2/N_{pts} = 400/140$  we get instead

	21	45	62
$\chi^2$	221	316	330

	21,45	21,62	45,62
$\chi^2$	213	178	230

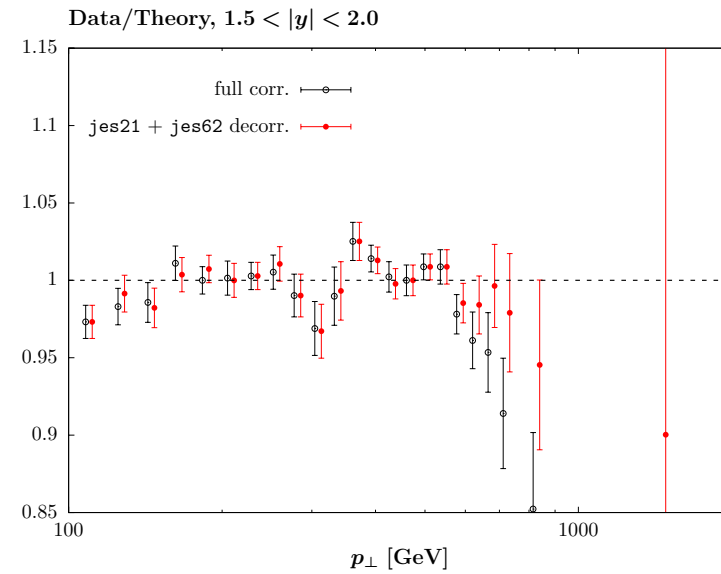
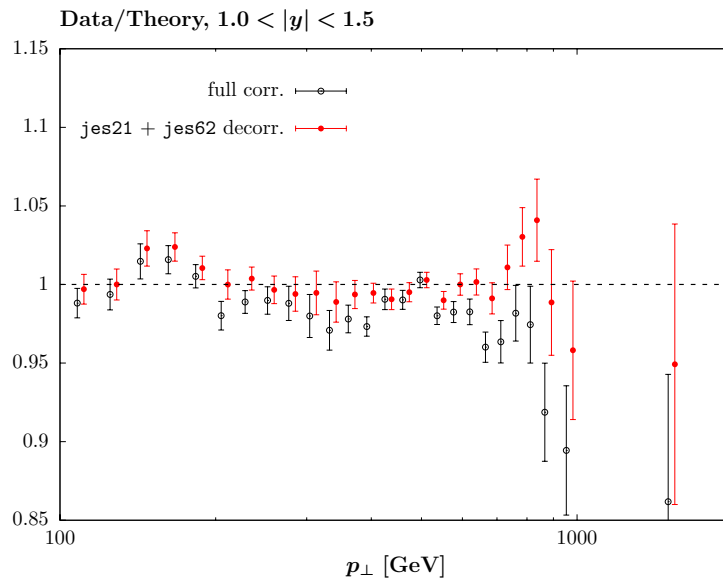
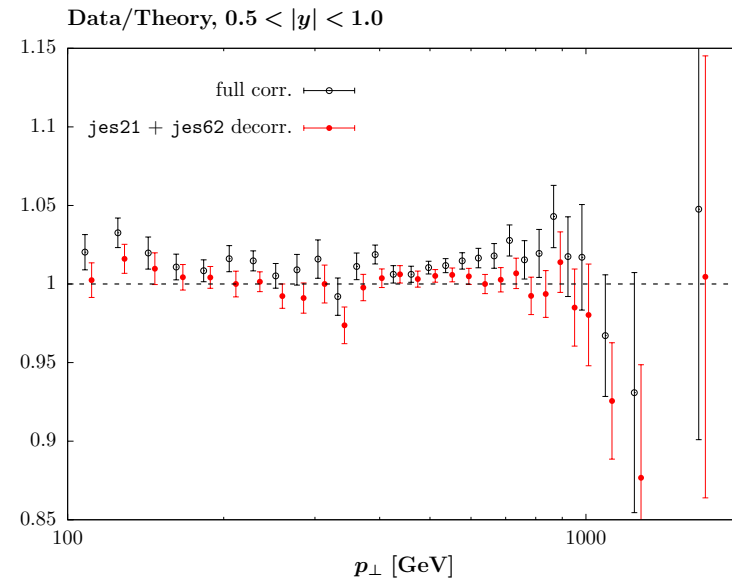
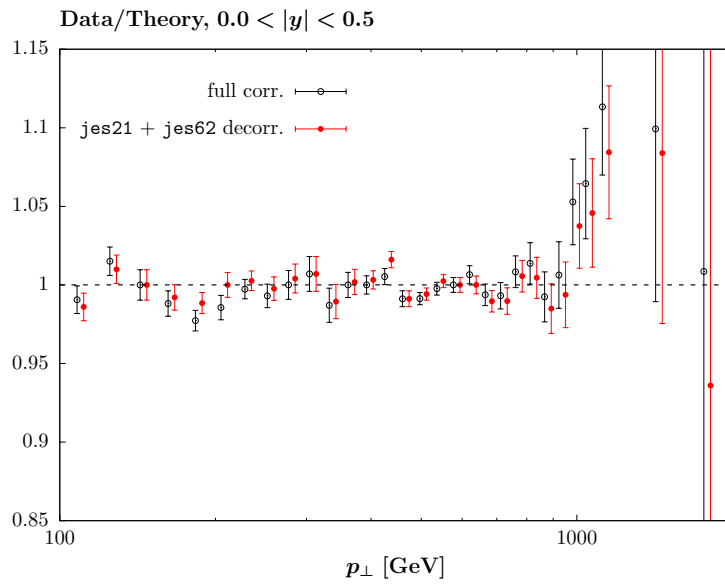
	21,45,62
$\chi^2$	172

Very significant improvement, particularly from decorrelating jes21.

Little improvement if jes45 decorrelated on top of jes21 and jes62.

With correlations between rapidity bins relaxed for just two sources of systematics  $\chi^2/N_{pts} = 178/140 = 1.27$ .

Anything other than just an interesting observation?



Fit quality evidently reasonable in all rapidity bins when jes21 and jes62 decorrelated between  $|y|$  bins.

## The effect of NNLO corrections.

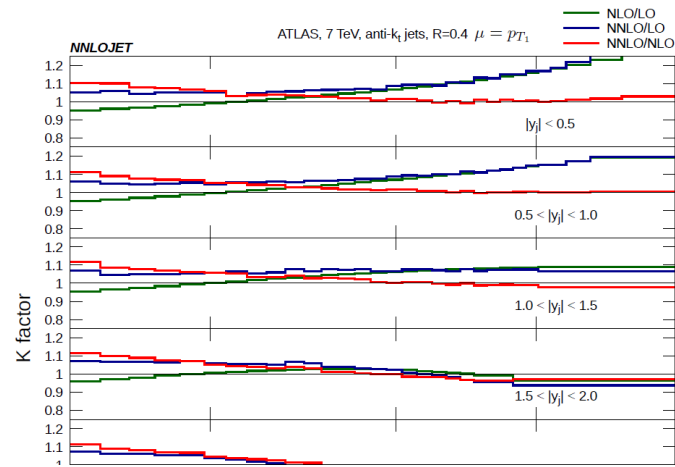
The NNLO corrections are now completed ([arXiv:1611.01460](https://arxiv.org/abs/1611.01460)) Currie *et al.*. Explicit K-factors available for ATLAS 7 TeV data.

Does this make a significant change to the conclusions just presented for the NLO fit?

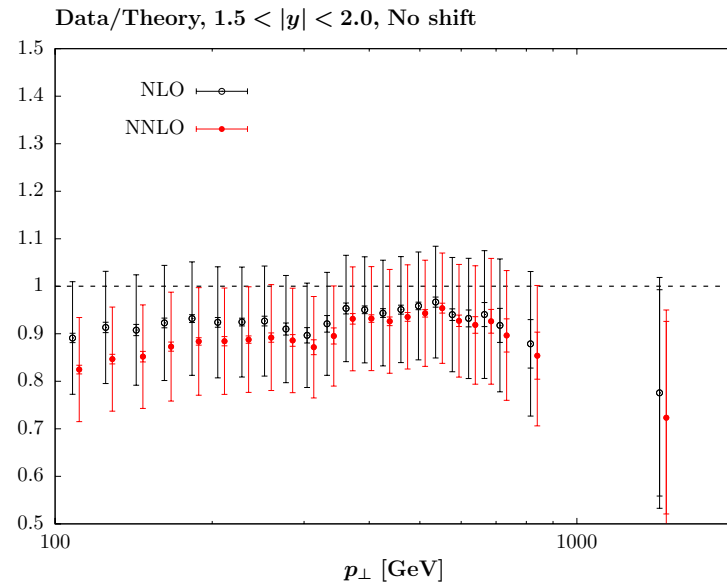
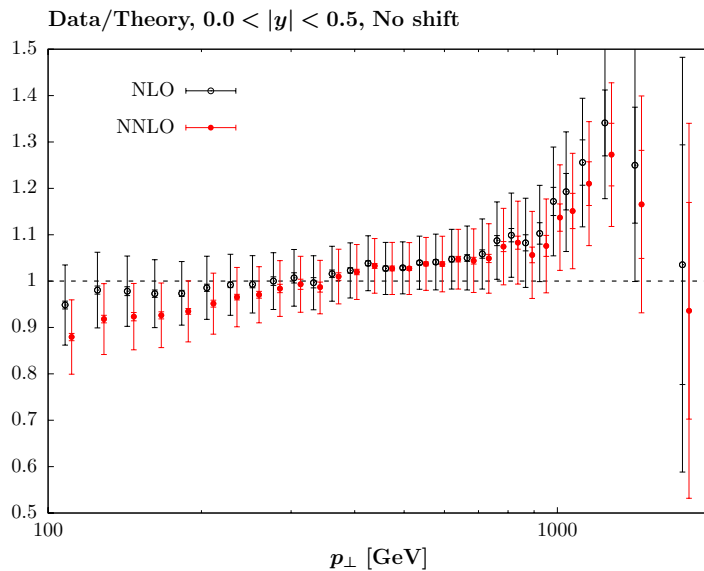
Fit to MMHT data set with HERA data updated to final combined set and Tevatron jet data excluded (usually fit using NNLO threshold approximation).

As at NLO take as default  $R = 0.4$  and  $\mu = p_{T,1}$ .

At low  $p_T$  NNLO correction significant ( $\mathcal{O}(10\%)$ ) and positive.



Moves unshifted data and theory further apart. Expect worse fit quality?



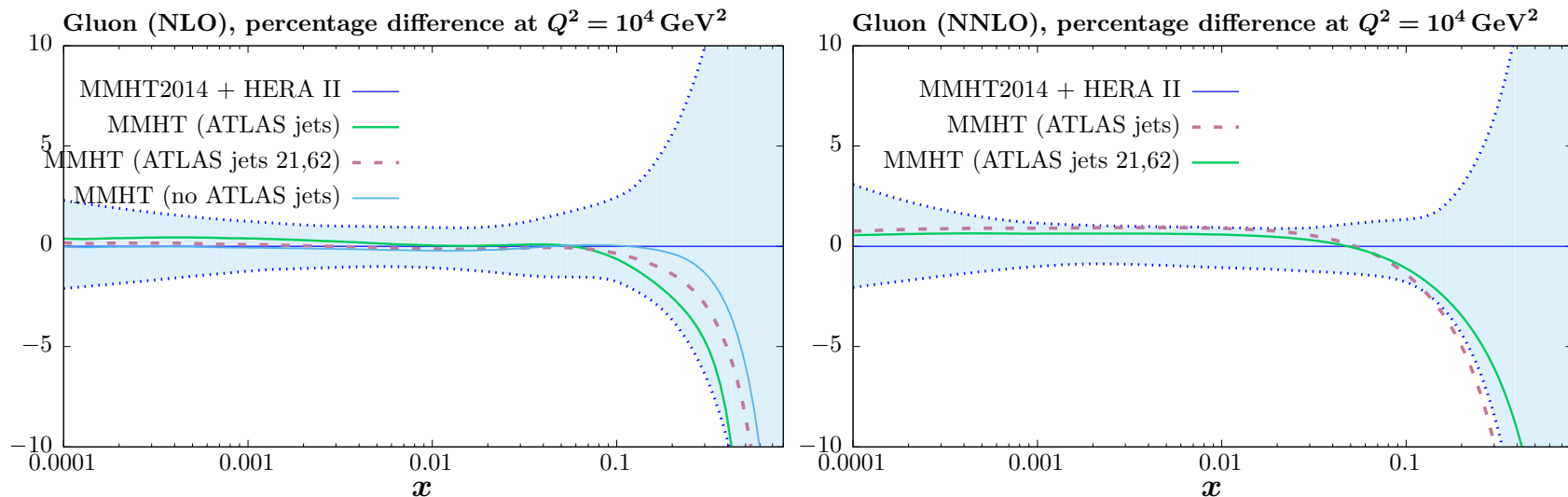
Values of  $\chi^2$  without (with) the **ATLAS** jets data in the fit

	Full Corr.	21,62 decorr.
$\chi^2$ , NLO	(413)400	(180)178
$\chi^2$ , NNLO	(443)427	(211)204

Find significant, if not dramatic, deterioration in fit quality in all cases.  
Not an issue of tension with other data.

By eye fit quality looks very similar to **NLO** but slightly higher penalty from shifts.

# Gluon including ATLAS jet data at NLO and NNLO



The effect on the best fit gluon is noticeable, but within (or at boundary) of uncertainties. Softer at very high  $x$ .

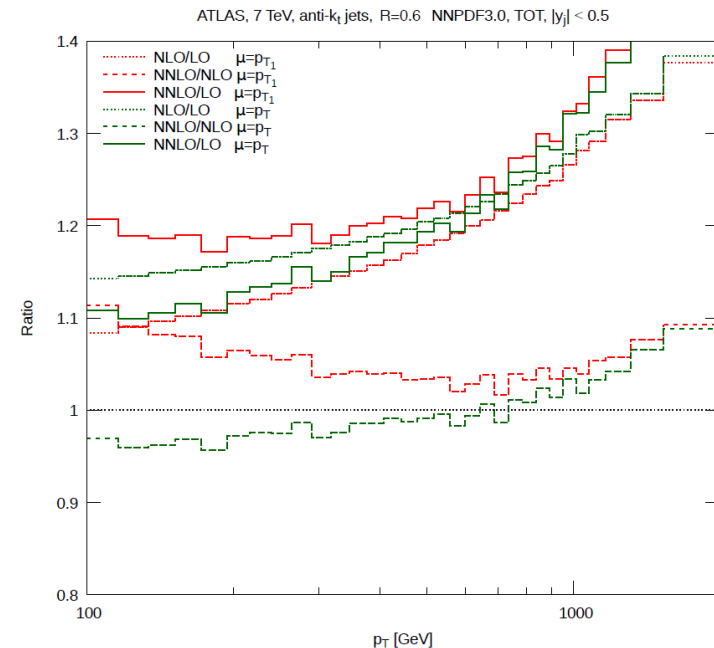
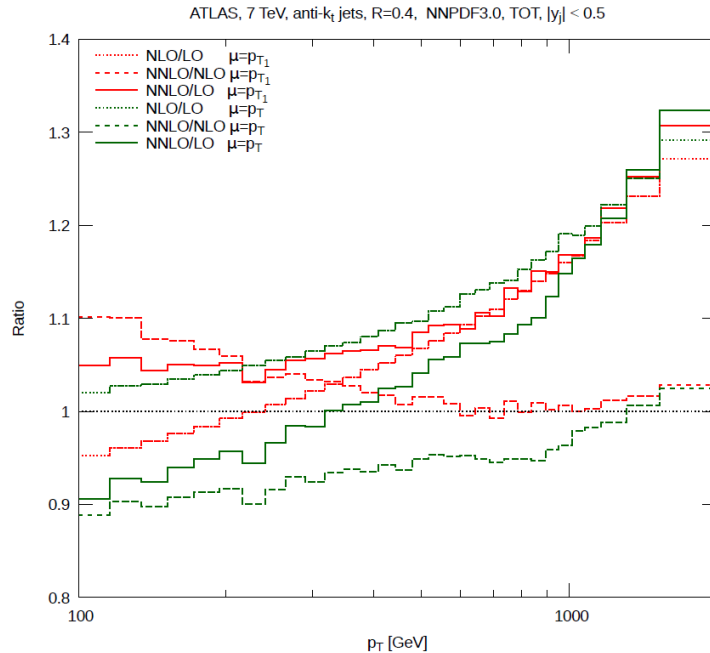
Slightly smaller effect at NLO than at NNLO.

Not very dependent on whether jes21 and jes 62 are decorrelated or not.

# CMS 7 TeV jets.

Here take as default  $R = 0.7$  and  $\mu = p_T$ .

Larger  $R$ , and  $\mu = p_T$  rather than  $\mu = p_{T1}$ , lead to more stable **NNLO** corrections.



Therefore good **NLO** fit ( $\chi^2/N_{pts} = 138/133$ ) very likely to be maintained at **NNLO** and little change in gluon expected.

More stability from **NLO** to **NNLO** expected for **ATLAS** jets if larger  $R$  and different scale chosen for fit?

## Inclusion of high precision **ATLAS W, Z** data [arXiv:1612.03016](https://arxiv.org/abs/1612.03016)

Confirm we obtain  $\chi^2/N_{pts} \sim 400/61$  from **MMHT14** PDFs at **NNLO**.

For slightly modified PDFs with final **HERA** combined data (and some more  $\sigma_{t\bar{t}}$  points) obtain  $\chi^2/N_{pts} \sim 387/61$ . Use this as our “baseline”. Essentially the same PDFs as from **MMHT** + final **HERA** data.

Including **ATLAS W, Z** data in fit goes to  $\chi^2/N_{pts} \sim 130/61$ , similar to **ATLAS** profiling. Use this as basis for study of effects on PDFs.

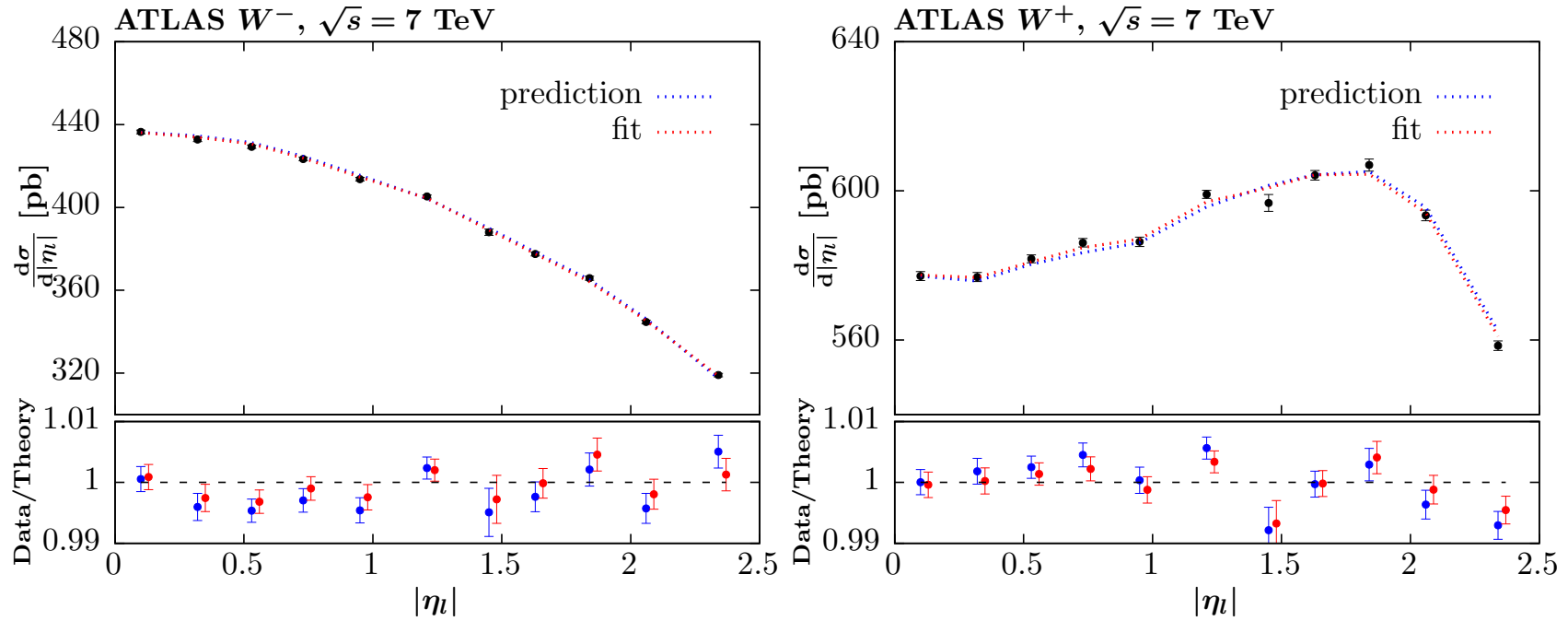
Deterioration in fit to other data  $\Delta\chi^2 = 54$ . Worst for **CMS double differential Z/ $\gamma$  data** ( $\Delta\chi^2 = 17$ ) and **CCFR/NuTeV dimuon data** ( $\Delta\chi^2 = 16$ ). For latter branching ratio requires 25% shift, but has uncertainty of 15%.

Other deterioration in fixed target DIS data, **E866** Drell-Yan asymmetry and **CDF W**-asymmetry.

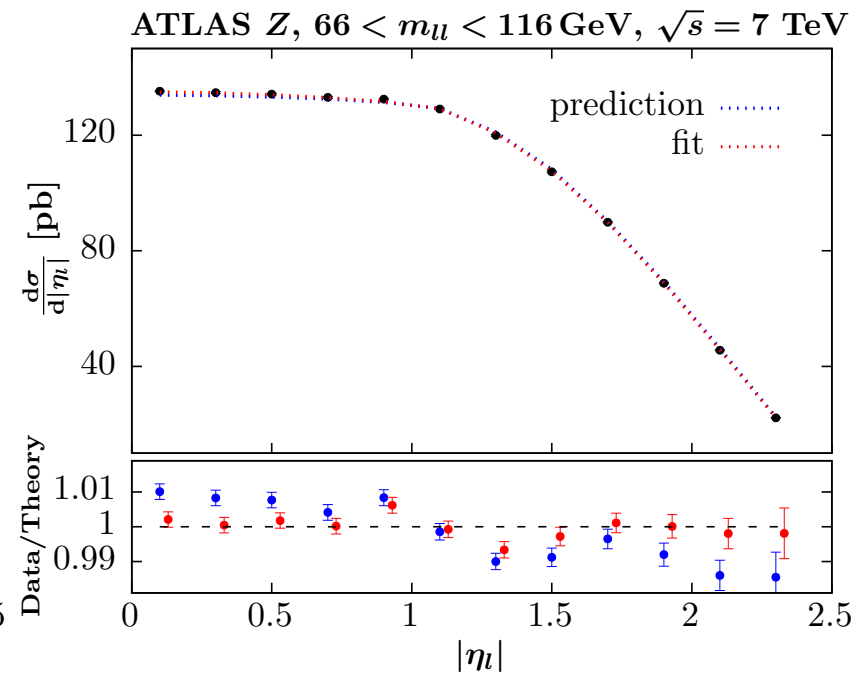
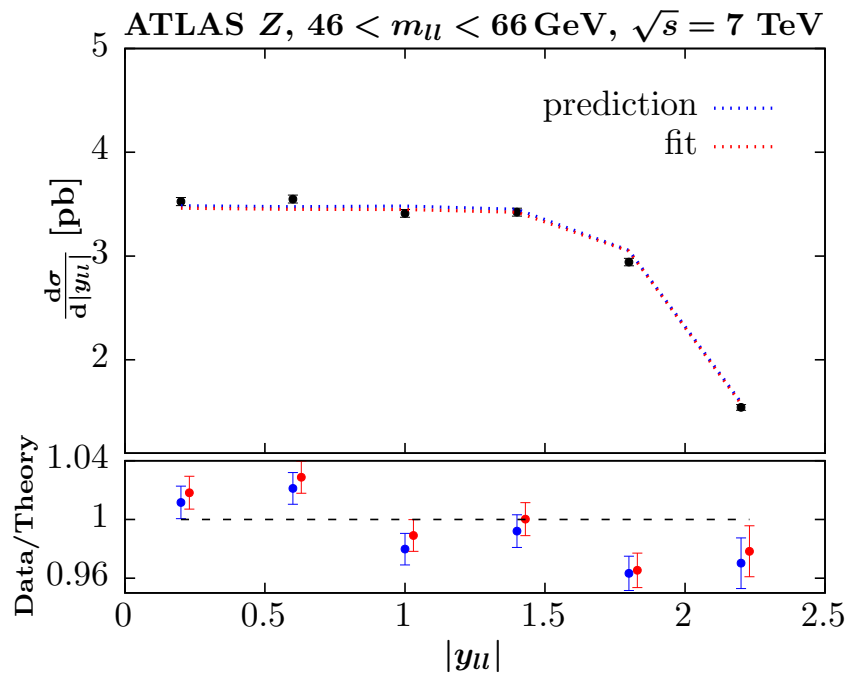
Generate PDFs with uncertainties including new **ATLAS W, Z** data, using same basis for 25 eigenvectors as for **MMHT2014**. **ATLAS W, Z** data constrains 5 eigenvector directions.



# Prediction and Fit to data



Slight reduction in lower  $|\eta|$   $W^-$  required and opposite for  $W^+$ .



Significant change in shape required for  $Z$  production, Higher at low  $|\eta|$  and lower at high  $|\eta|$

Even with fit difficulty in shape for lower mass data.

## Additional fits with high precision ATLAS W, Z data.

Increase weight of new ATLAS data by factor 10.

$\chi^2$  improves to  $\chi^2/N_{pts} \sim 121/61$ . Deterioration in fit to other data  $\Delta\chi^2 = 92$ .

Further increase in CMS double differential Z/ $\gamma$  data ( $\Delta\chi^2 = 24$ ) and E866 Drell-Yan asymmetry. Dimuon data not any worse.

Now also deterioration in HERA combined data, both NC and CC and CDF differential Z/ $\gamma$  data.

Also try fit where all other new LHC data from LHCb and CMS included. Compared to baseline plus ATLAS W, Z data very little change  $\Delta\chi^2 = 3$  in total, and essentially no change in ATLAS W, Z data.

However, inclusion of ATLAS W, Z data lowers  $\chi^2$  for new LHC (plus final D0) data by  $\Delta\chi^2 = -10$ .

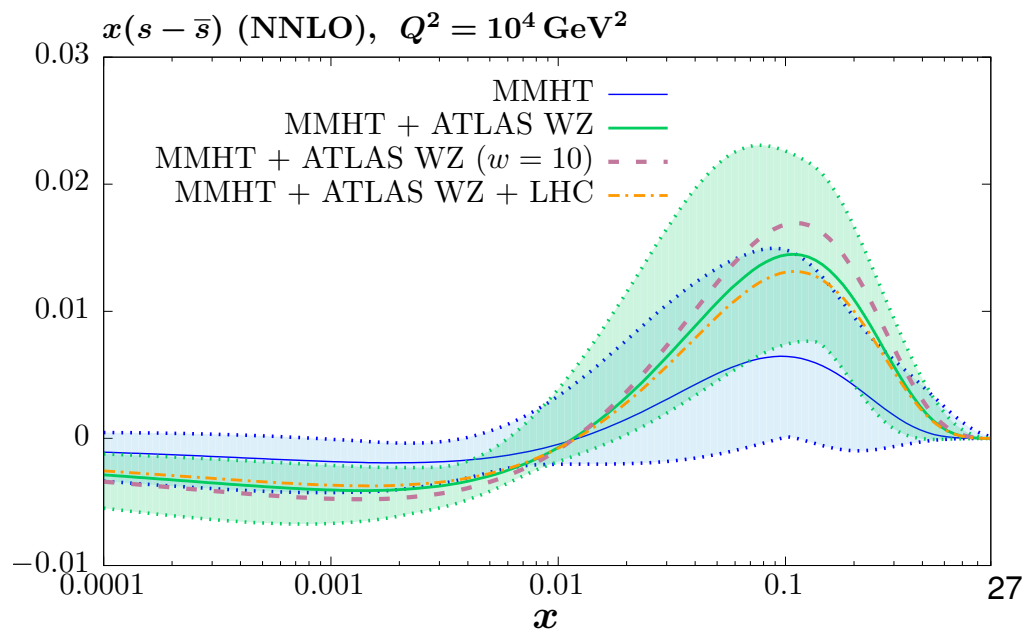
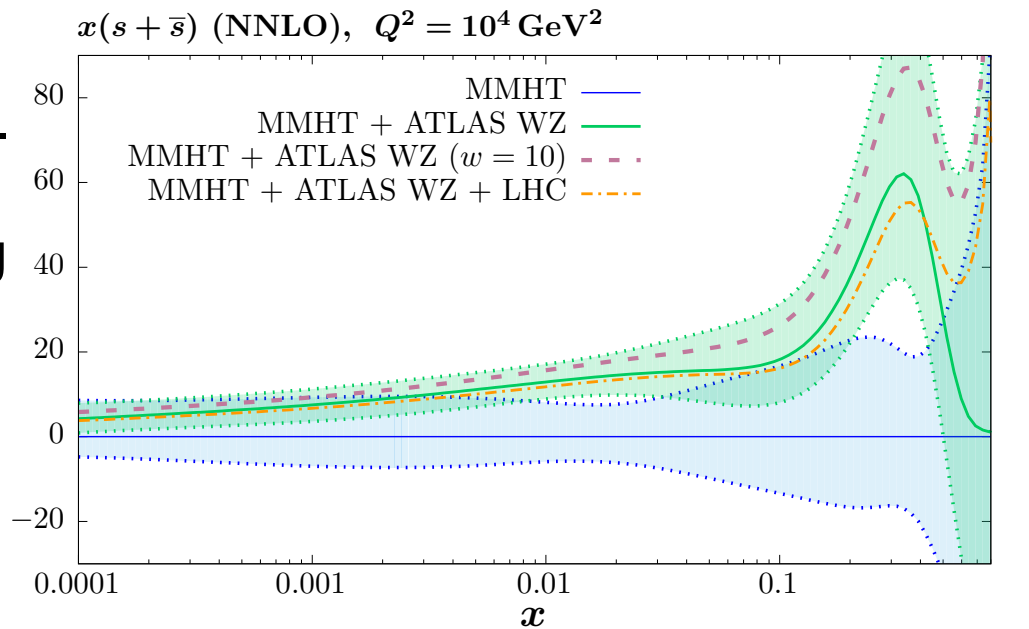
Hence ATLAS W, Z data and other new LHC data compatible and pull in same direction. Only CMS  $W + c$  deteriorates slightly.

## Effect on PDFs

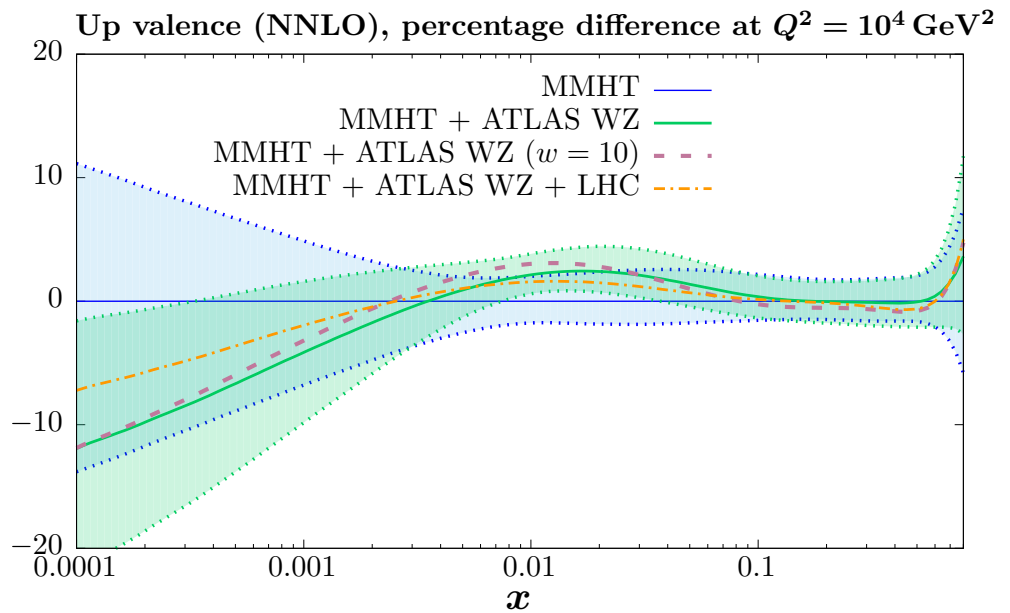
Large increase in  $s + \bar{s}$  and decrease in uncertainty. Correlation with fit to dimuon data (lower branching ratio) leads to increase at high  $x$ . (Note negative NNLO correction [Phys. Rev. Lett. 116 \(2016\), Berger \*et al.\*](#))

Larger for  $x > 0.1$  due to significant down quark contribution in this region despite Cabibbo suppression.

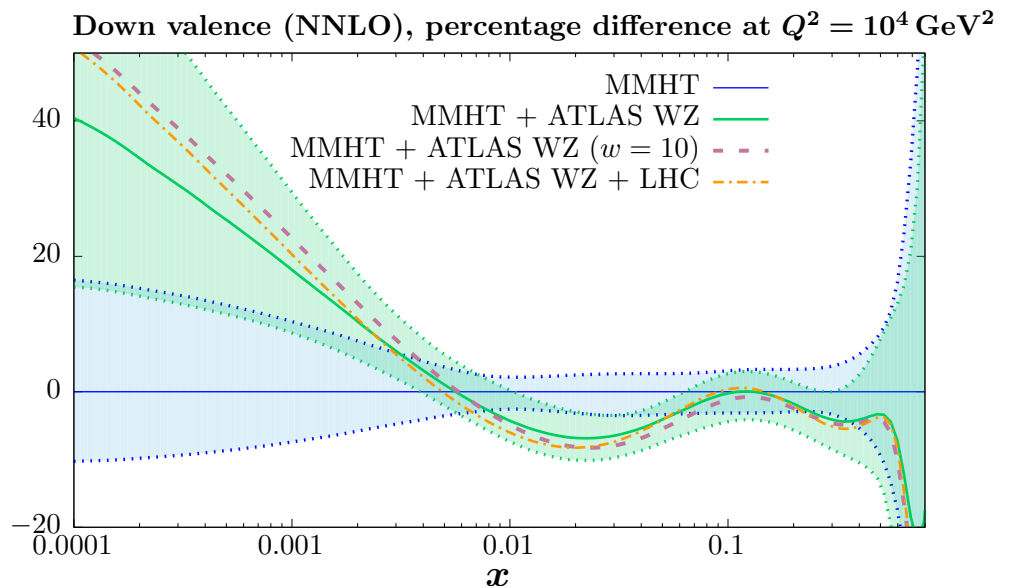
There is impact on  $s - \bar{s}$  uncertainty, from the change in branching ratio.



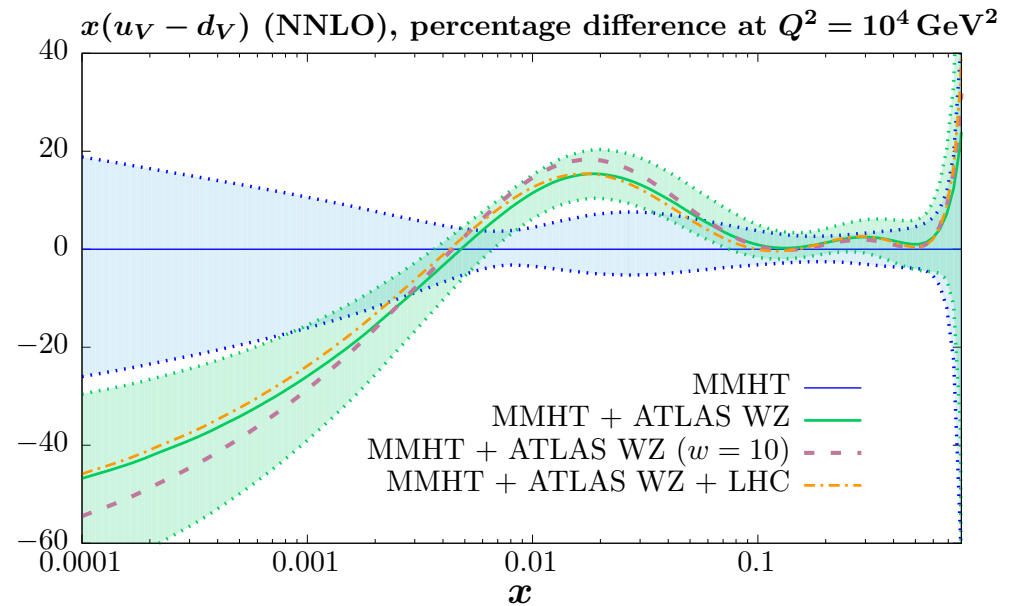
Significant impact on shape of valence quarks.



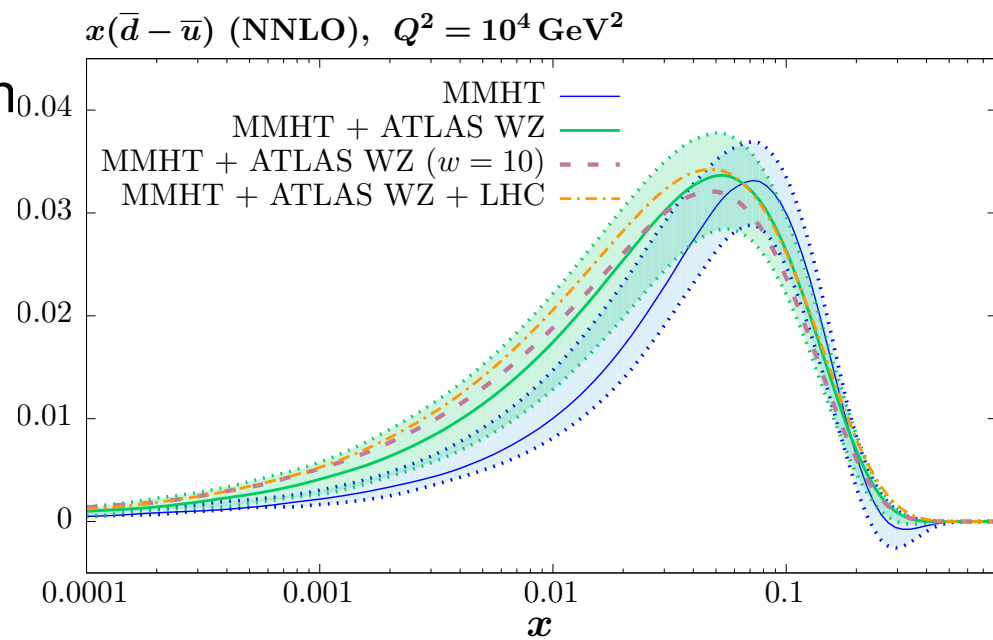
Same direction as impact of other new **LHC** data.



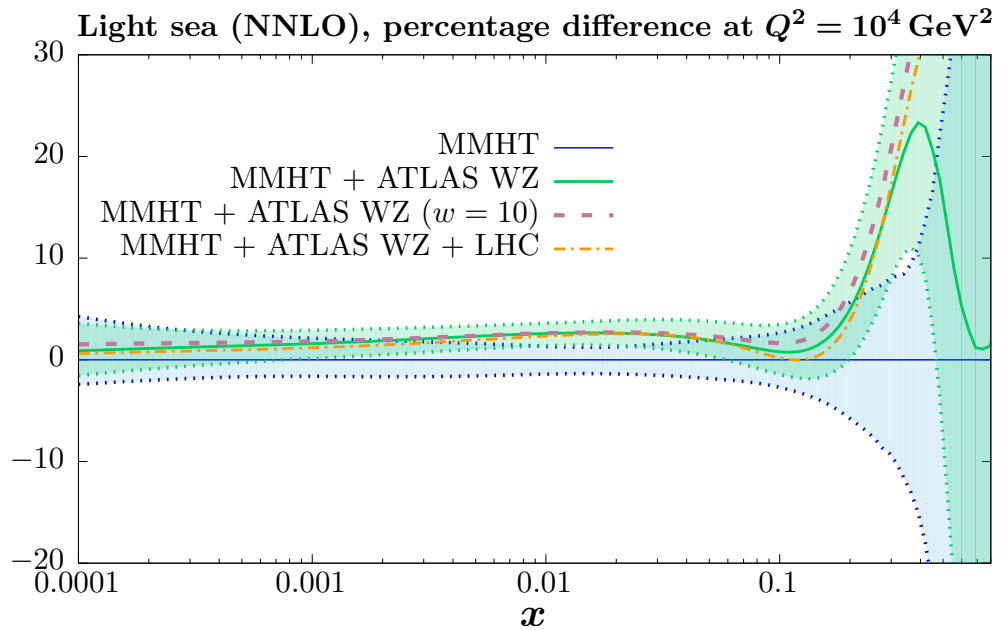
As implied by individual distributions, significant change in  $u_V - d_V$ .



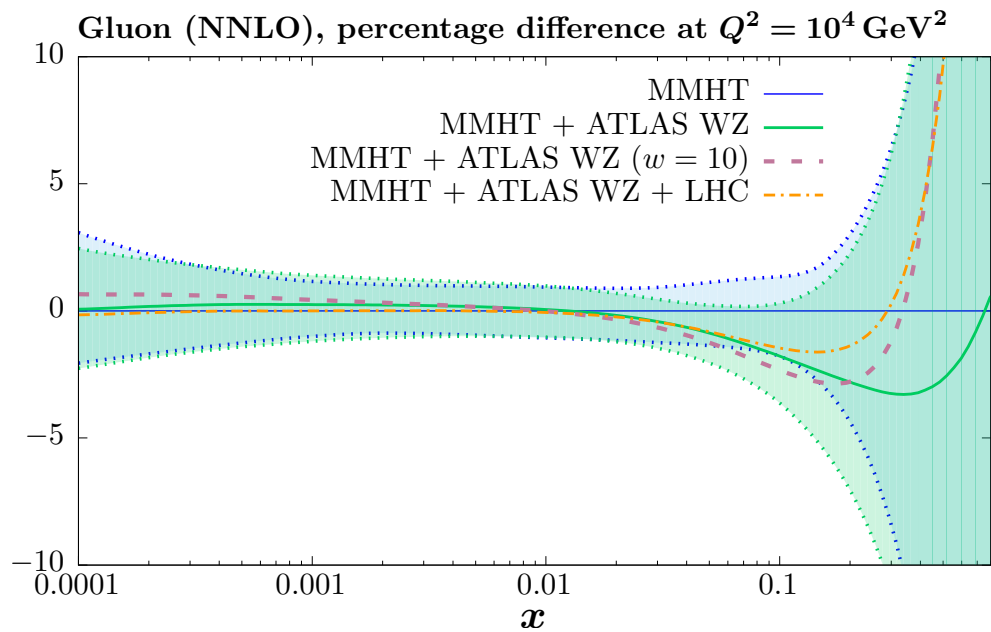
Shift in best fit  $\bar{d} - \bar{u}$  accompanying deterioration in fit to E866 Drell-Yan asymmetry.



Change in strange quark affects sea, making it generally larger.



Little impact on gluon distribution. .



## Extension of $\bar{d} - \bar{u}$ parameterisation.

Currently use **3** free parameters, i.e.

$$(\bar{d} - \bar{u})(x, Q_0^2) = A(1 - x)^{\eta_{sea}+2} x^\delta (1 + \gamma x + \Delta x^2),$$

Extend to

$$(\bar{d} - \bar{u})(x, Q_0^2) = A(1 - x)^{\eta_{sea}+2} x^\delta (1 + \sum_{i=1}^4 a_i T_i(1 - 2x^{\frac{1}{2}})),$$

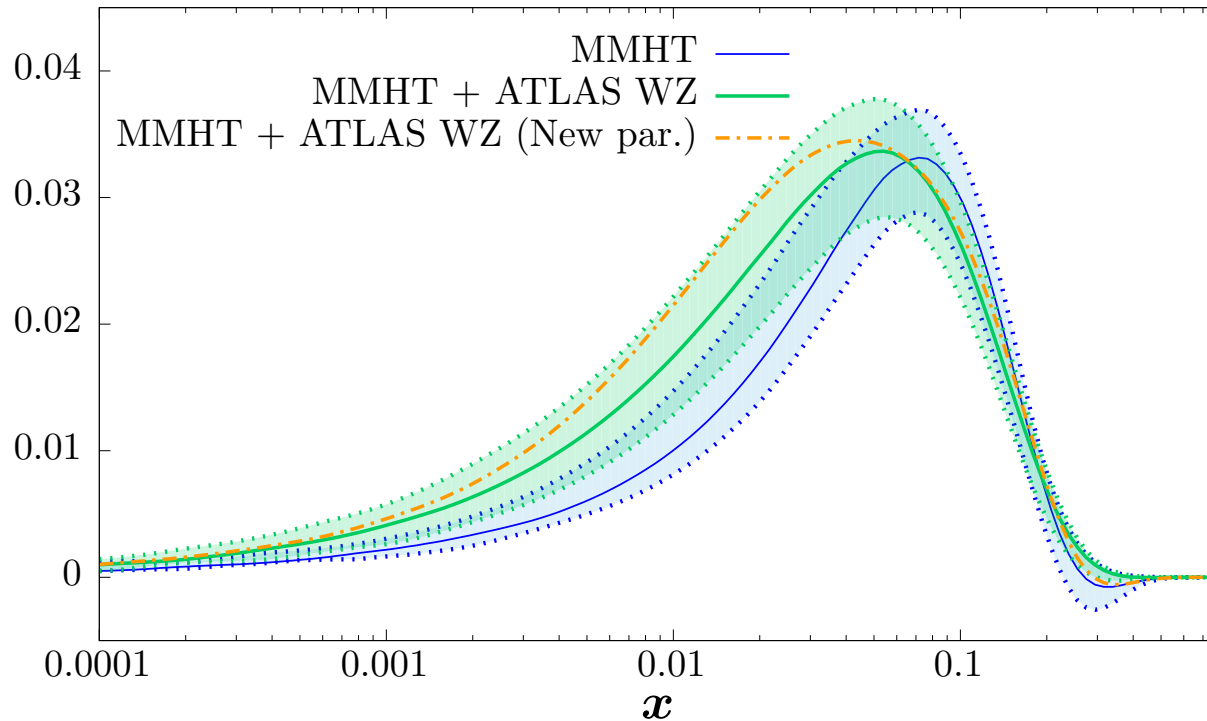
where  $T_i(1 - 2x^{\frac{1}{2}})$  are Chebyshev polynomials. So **5** free parameters. Easily allows multiple turning points (seen in first fit iteration).

Global fit including new **LHC** data and new **ATLAS  $W, Z$**  data improves by **10** units, but over **5** of this in **E866 Drell Yan asymmetry**.

Parameterisation alleviates some tension between **ATLAS** data and **Drell Yan asymmetry**.



$x(\bar{d} - \bar{u})$  (NNLO),  $Q^2 = 10^4 \text{ GeV}^2$



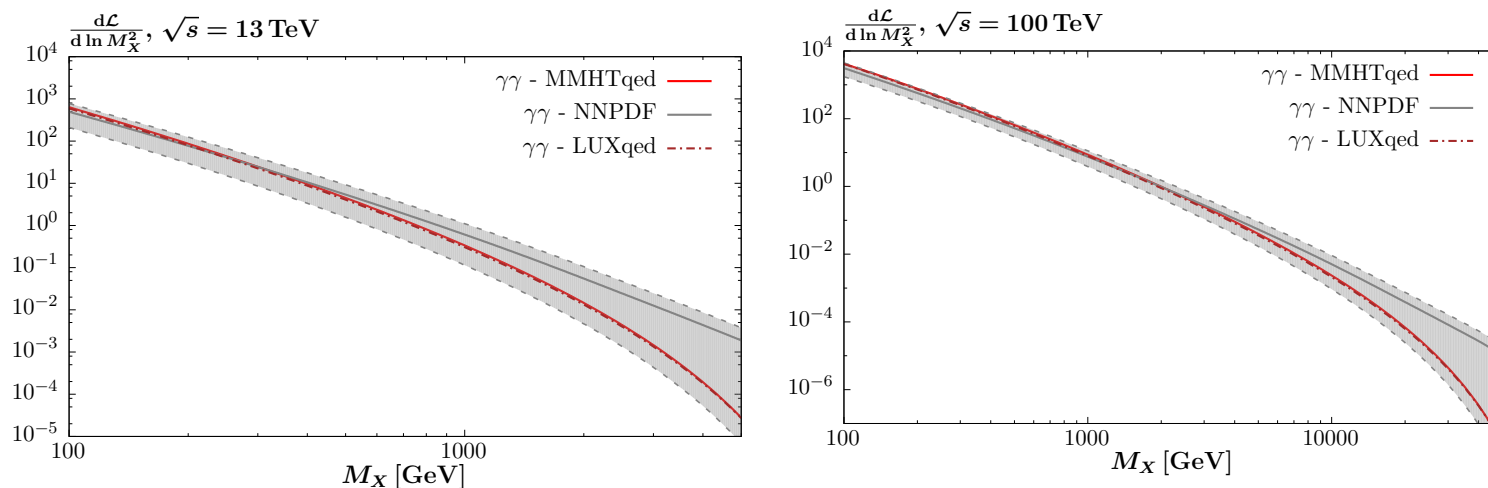
New  $(\bar{d} - \bar{u})$  distribution similar at high  $x$  to previous one. (Dips to negative values at low- $x$  allowed by, and seen using new parameterisation.)

Now a smaller decrease towards zero at low  $x$  beyond edge of previous uncertainty band.

## PDFs with QED corrections

We will base photon input for PDFs at low  $Q^2$  on the LUXqed prescription which demonstrated determination in terms of structure function, and hence precision of at worst a few percent.

Effect of photon evolution fully incorporated to couple with that of quarks and gluon for both proton and neutron.



Final details of transition from low  $Q^2$  to be finalised (e.g. elastic (coherent) contribution still important above  $Q_0^2 = 1\text{GeV}^2$ , but MMHT photon (Nathvani) very similar to LUXqed.

## Connection to xFitter

I will briefly discuss a number of connections working in both directions.

Main historical connection, the general mass variable flavour number scheme **GM-VFNS** used by **MSTW/MMHT** has been used for a long time by **HERAFitter/xFitter**.

Also discuss various possibilities for the future - very much a case of “to be discussed”.

The **GM-VFNS** can be defined by demanding equivalence of the  $n_f$  light flavour and  $n_f + 1$  light flavour descriptions at all orders – above transition point  $n_f \rightarrow n_f + 1$

$$F(x, Q^2) = C_k^{FF, n_f}(Q^2/m_H^2) \otimes f_k^{n_f}(Q^2) = C_j^{VF, n_f+1}(Q^2/m_H^2) \otimes f_j^{n_f+1}(Q^2) \\ \equiv C_j^{VF, n_f+1}(Q^2/m_H^2) \otimes A_{jk}(Q^2/m_H^2) \otimes f_k^{n_f}(Q^2).$$

Hence, the **VFNS** coefficient functions satisfy

$$C_k^{FF, n_f}(Q^2/m_H^2) = C_j^{VF, n_f+1}(Q^2/m_H^2) \otimes A_{jk}(Q^2/m_H^2),$$

which at  $\mathcal{O}(\alpha_S)$  gives (in  $\overline{MS}$  scheme)

$$C_{2, Hg}^{FF, n_f, (1)}\left(\frac{Q^2}{m_H^2}\right) = C_{2, HH}^{VF, n_f+1, (0)}\left(\frac{Q^2}{m_H^2}\right) \otimes P_{qg}^0 \ln(Q^2/m_H^2) + C_{2, Hg}^{VF, n_f+1, (1)}\left(\frac{Q^2}{m_H^2}\right),$$

The **VFNS** coefficient functions tend to the  $m=0$  limits as  $Q^2/m_H^2 \rightarrow \infty$ .

However,  $C_j^{VF}(Q^2/m_H^2)$  only uniquely defined in this limit.

Can swap  $\mathcal{O}(m_H^2/Q^2)$  terms between  $C_{2, HH}^{VF, 0}(Q^2/m_H^2)$  and  $C_{2, g}^{VF, 1}(Q^2/m_H^2)$ .

Have the freedom to modify the heavy quark coefficient function, by default

$$C_{2,HH}^{VF,0}(Q^2/m_H^2, z) = \delta(z - x_{\max}).$$

Appears in convolutions for higher order subtraction terms, so do not want complicated  $x$  dependence. Simple choice.

$$C_{2,HH}^{VF,0}(Q^2/m_H^2, z) \rightarrow (1 + b(m_H^2/Q^2)^c)\delta(z - x_{\max}),$$

where  $c$  really encompasses  $(m_H^2/Q^2)$  with logarithmic corrections.

Can also modify argument of  $\delta$ -function, as in Intermediate Mass (IM) scheme of **Nadolsky, Tung**. Let argument of heavy quark contribution change like

$$\xi = x/x_{\max} \rightarrow x(1 + (x(1 + 4m_H^2/Q^2))^d 4m_H^2/Q^2),$$

so kinematic limit stays the same, but if  $d > 0$  small  $x$  less suppressed, or if  $d < 0$  (must be  $> -1$ ) small  $x$  more suppressed.

Default  $a, b, c, d$  all zero. Limit either by fit quality or *sensible* choices.

6 extreme variations tried.

GMVFNS1 –  $b = -1, c = 1$ .

GMVFNS2 –  $b = -1, c = 0.5$ .

GMVFNS1 –  $a = 1$ .

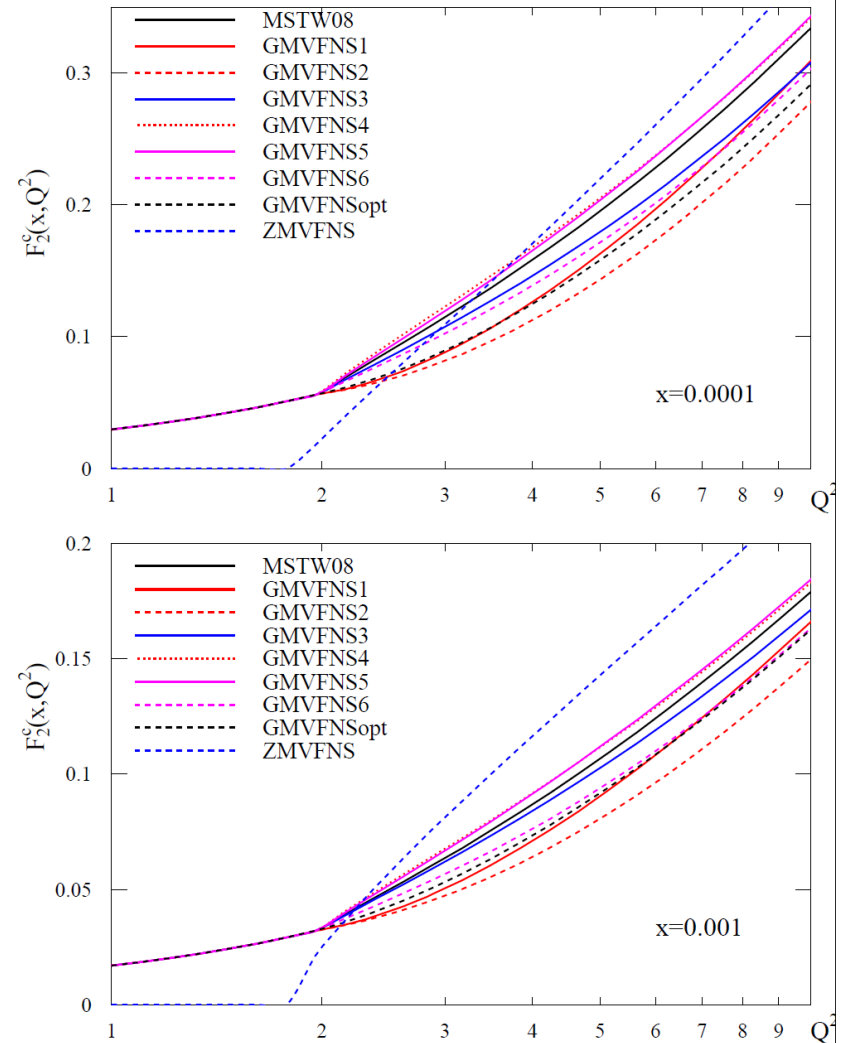
GMVFNS1 –  $b = +0.3, c = 1$  – fit.

GMVFNS1 –  $d = 0.1$  – fit.

GMVFNS1 –  $d = -0.2$  – fit.

Variations in  $F_2^c(x, Q^2)$  near the transition point at NLO due to different choices of GM-VFNS.

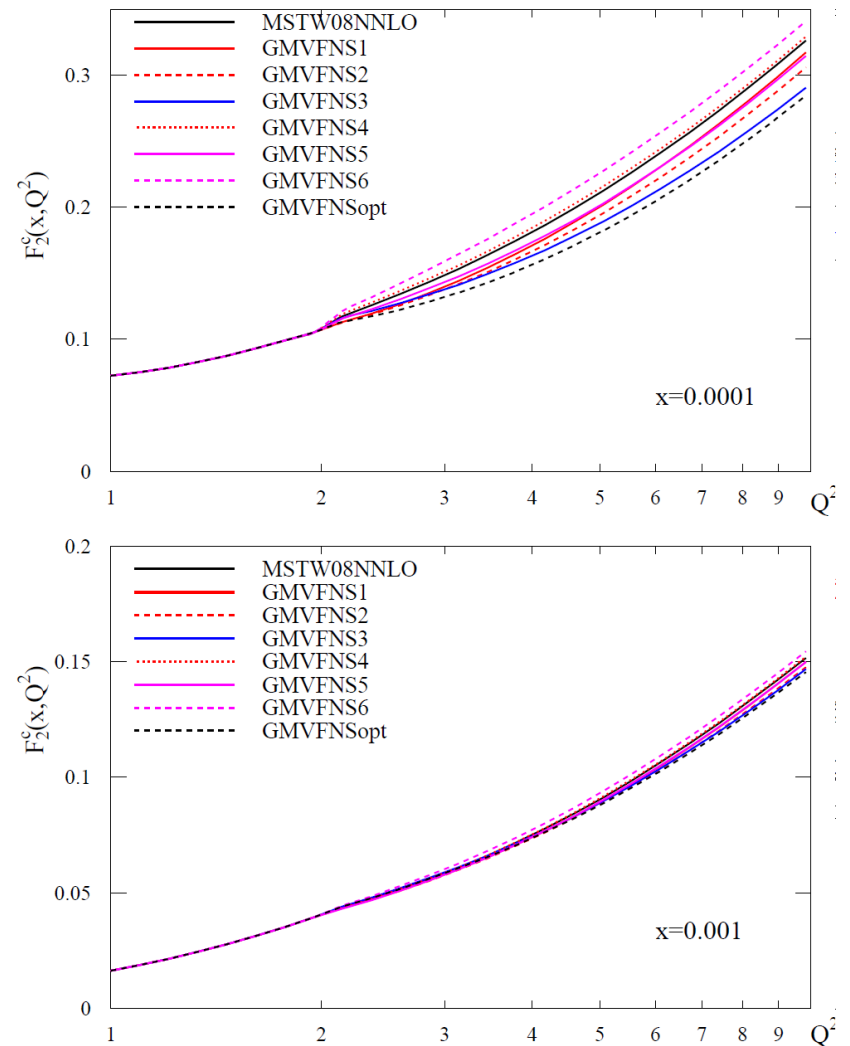
Optimal,  $a = 1, b = -2/3, c = 1$ , smooth behaviour.



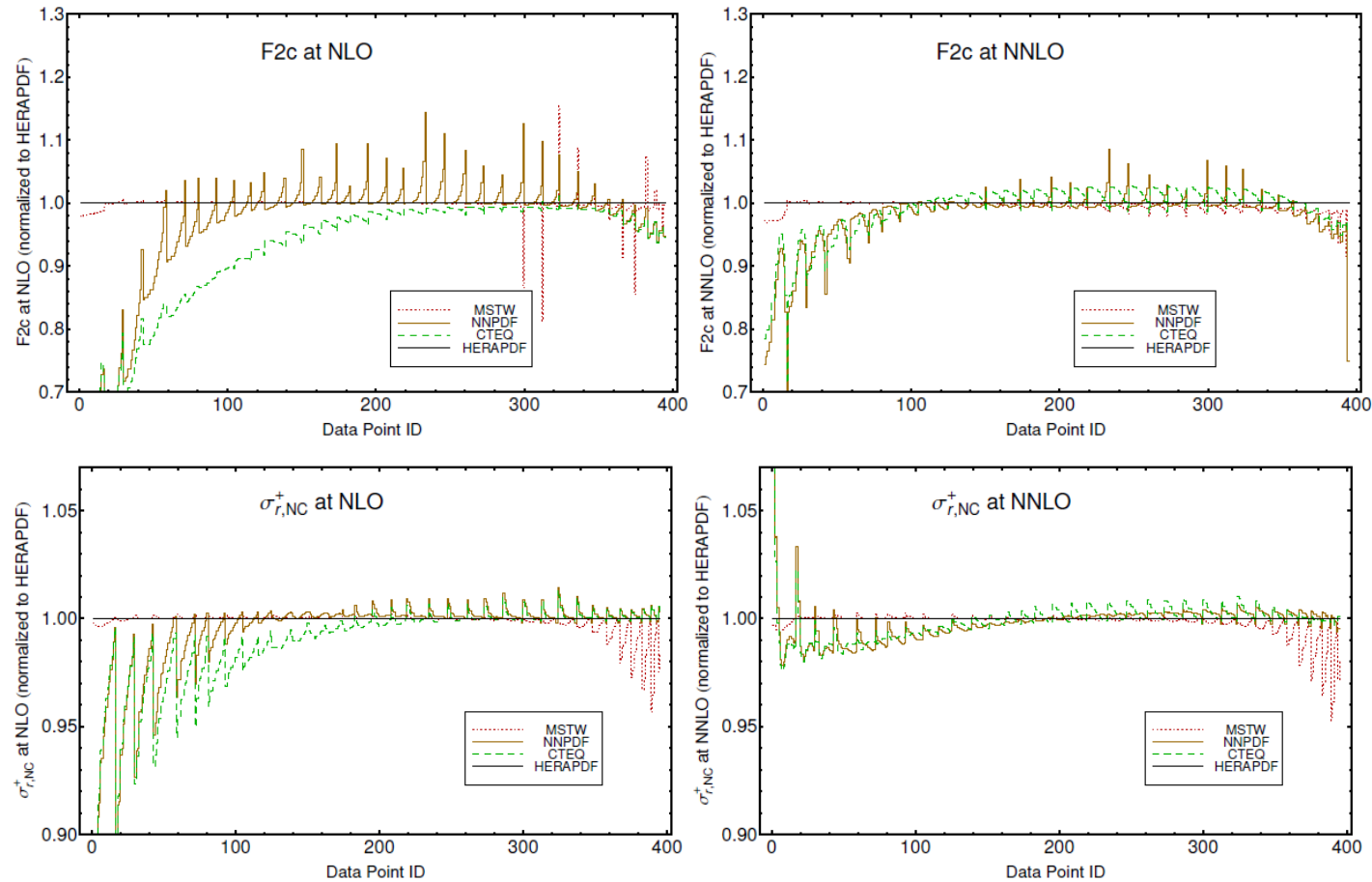
Variations in  $F_2^c(x, Q^2)$  near the transition point due to different choices of GM-VFNS at NNLO.

Very much reduced, almost zero variation until very small  $x$ .

Shows that NNLO evolution effects most important in this regime.



Also see convergence between groups in [Les Houches](#) benchmark study.



**NNLO TR** scheme larger at lowest  $Q^2$  due to use of  $\mathcal{O}(\alpha_s^3)$  coefficient function.



Not currently implemented for charged current cross sections – not too important for HERA data where zero mass for charm is a good approximation.

In principle available for NLO and for an old version of approximate NNLO.

Better to wait for the inclusion of the real NNLO corrections of Phys. Rev. Lett. 116 (2016), Berger *et al.*?

## MMHT uses of xFitter

Used for aid in fitting newly appearing LHC data.

Both data points, uncertainties and [applgrids](#) and [K-factors](#).

For example, most recent [ATLAS  \$W, Z\$](#)  data information obtained from <http://www.hepforge.org/archive/xfitter/1612.03016.tar.gz>.

(Aspire to check theory calculations independently).

Will definitely find it useful to use similar information provided for processes where initial state photon is important, e.g. recent [ATLAS high mass Drell Yan](#).

## Other possible exchange of ideas?

Best fit procedure - based on **Numerical Recipes**.

Uses **Levenberg-Marquardt** algorithm based on a combination of steepest decent and **Gauss-Newton** iteration. For parameters  $a_i$

$$a_{i+1} = a_i - (H + \lambda \text{diag}[H])_{ik}^{-1} \nabla_k \chi^2(a)$$

where  $H$  is the Hessian matrix and  $\lambda$  a constant.  $\lambda$  is varied to maximize convergence.

**Chebyshev polynomial** parameterisation. Not sure this requires any code.

Dynamical tolerance procedure (as alternative to  $\Delta \chi^2 = 1$  if constraints from a variety of data sets start to seem too constraining?)

## Conclusions

New **HERA II** combined data studied with context of **MMHT2014** PDFs. Fit quality good – better at **NNLO**. No very significant changes in PDFs or predictions. Slight reduction in uncertainties.

Predictions turn out to be good for many **LHC** data previously not included in the fit. Few changes in central values, but some data reduce uncertainties, mainly in strange and low- $x$  valence quarks.

Failure to fit **ATLAS 7 TeV** jet data at **NLO** - common with other groups. **NNLO** leads to little change. Poor fit driven by a small number of correlated systematic uncertainties. Much improved with model of reduced rapidity correlation.

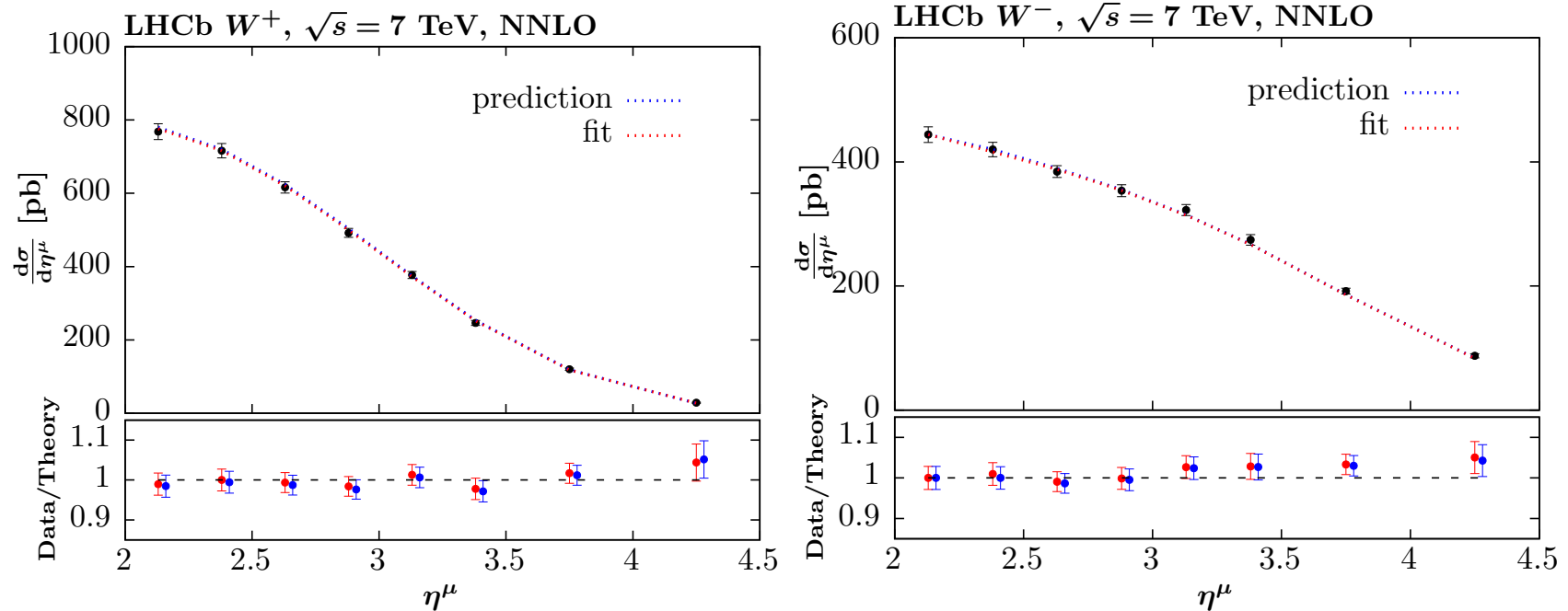
Prediction for new **ATLAS  $W, Z$**  data poor but much improved with new fit. Tensions with older data, particularly dimuon, and more-so **CMS 7 TeV** Drell Yan data. Increase in strange quark. Very compatible with most recent **CMS, LHCb** data in the fit. Improved ( $\bar{d} - \bar{u}$ ) helps.

Work on updated PDFs with **QED** corrections (**Nathvani**).

Variety of possible areas of reciprocity with **xFitter**.

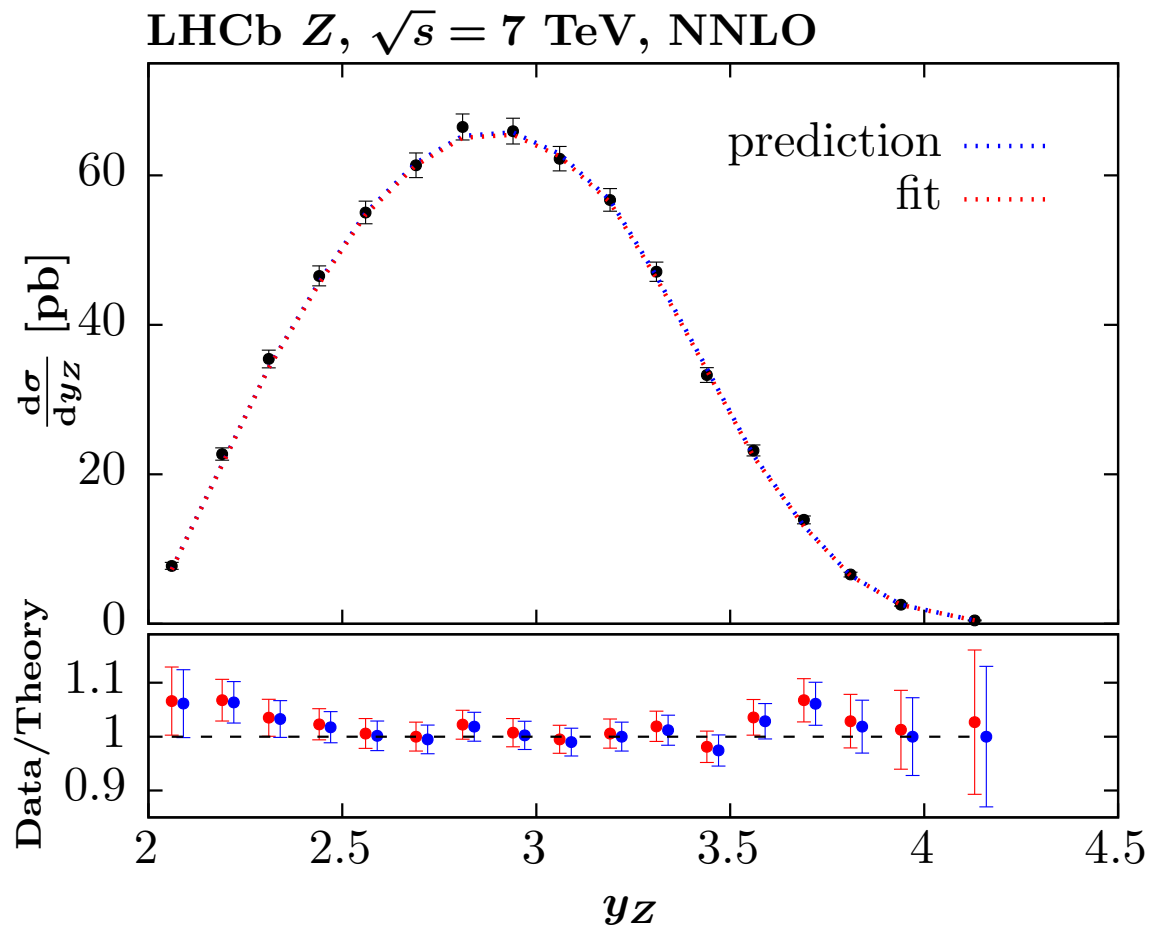
# Back -up

# New data on high rapidity $W$ production at LHCb at 7 TeV.



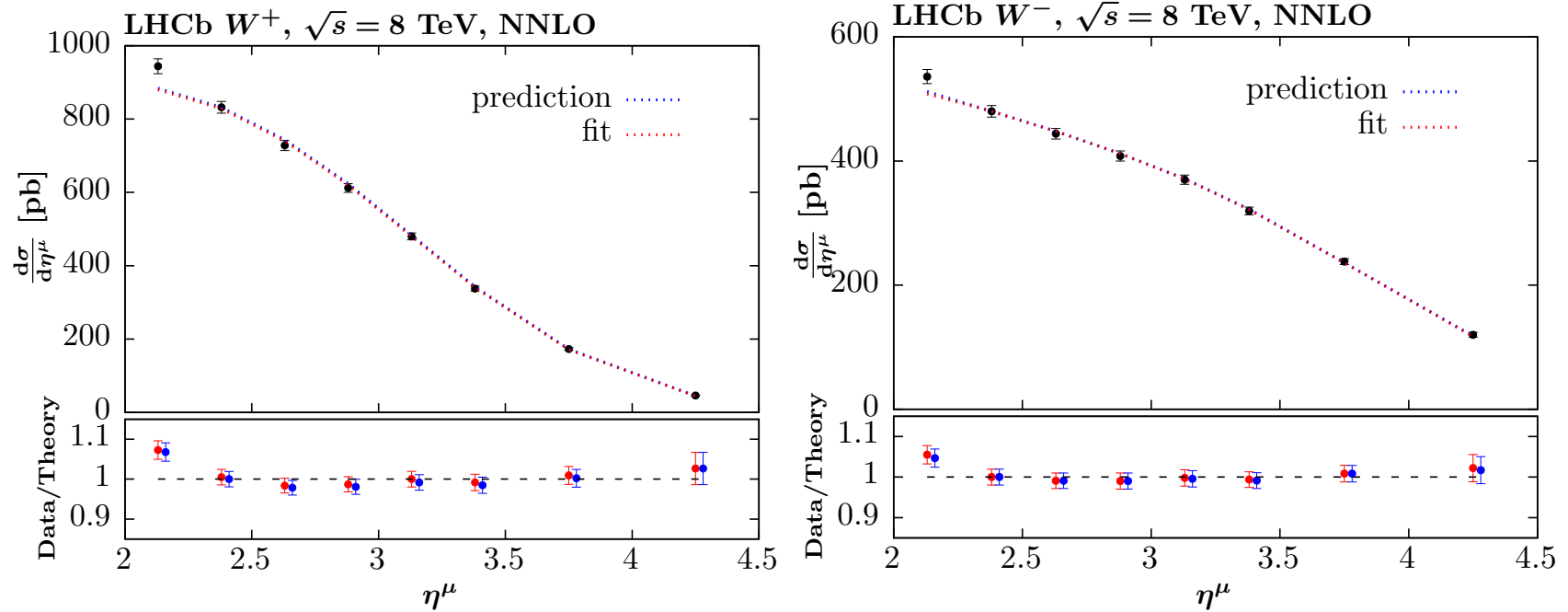
Generally perfectly good agreement using NNLO. Uncertainties added in quadrature on plot, but covariance matrix used in fit.

# New data on high rapidity $Z$ production at LHCb at 7 TeV.



Generally perfectly good agreement using NNLO. A little low at low  $y_Z$ .

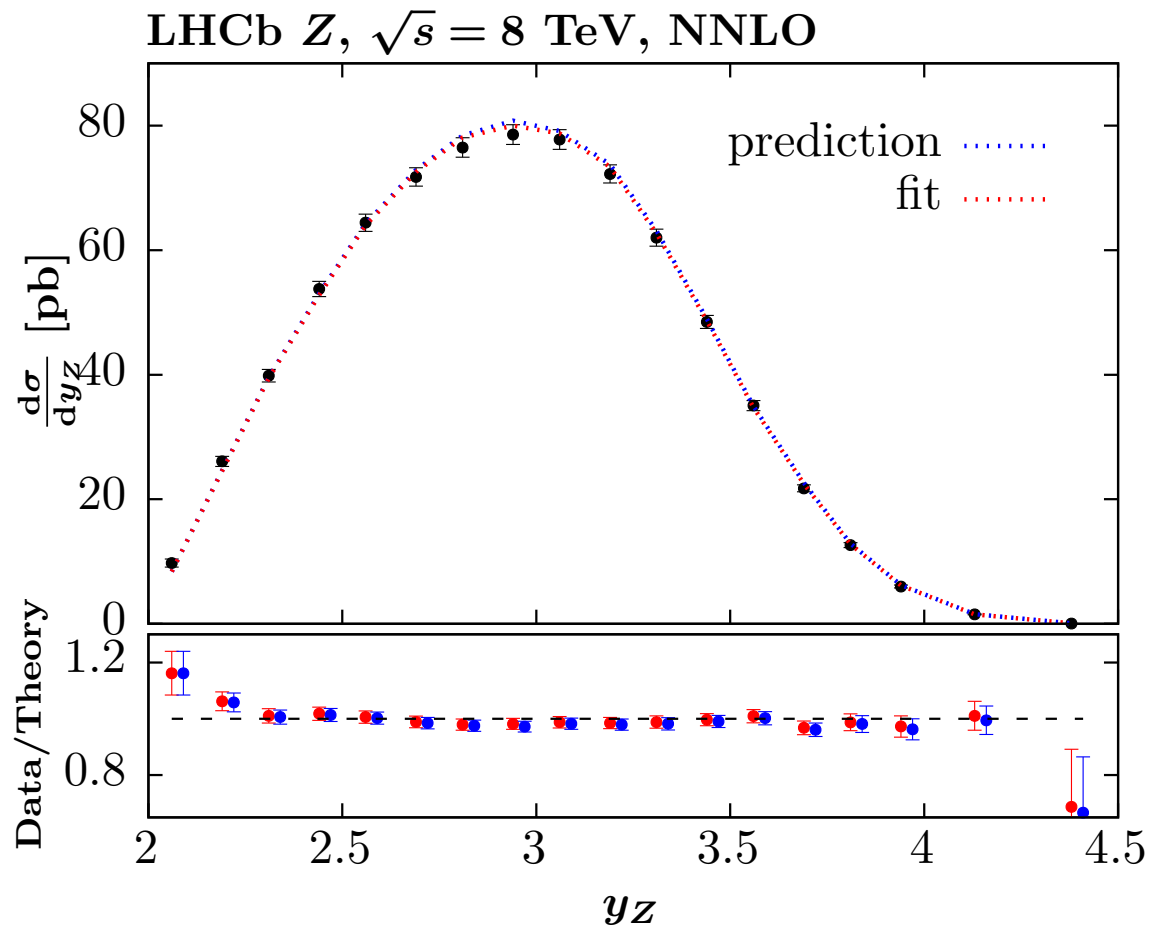
# New data on high rapidity $W$ production at LHCb at 8 TeV.



Good fit except at lowest  $\eta_\mu$  point in each case.

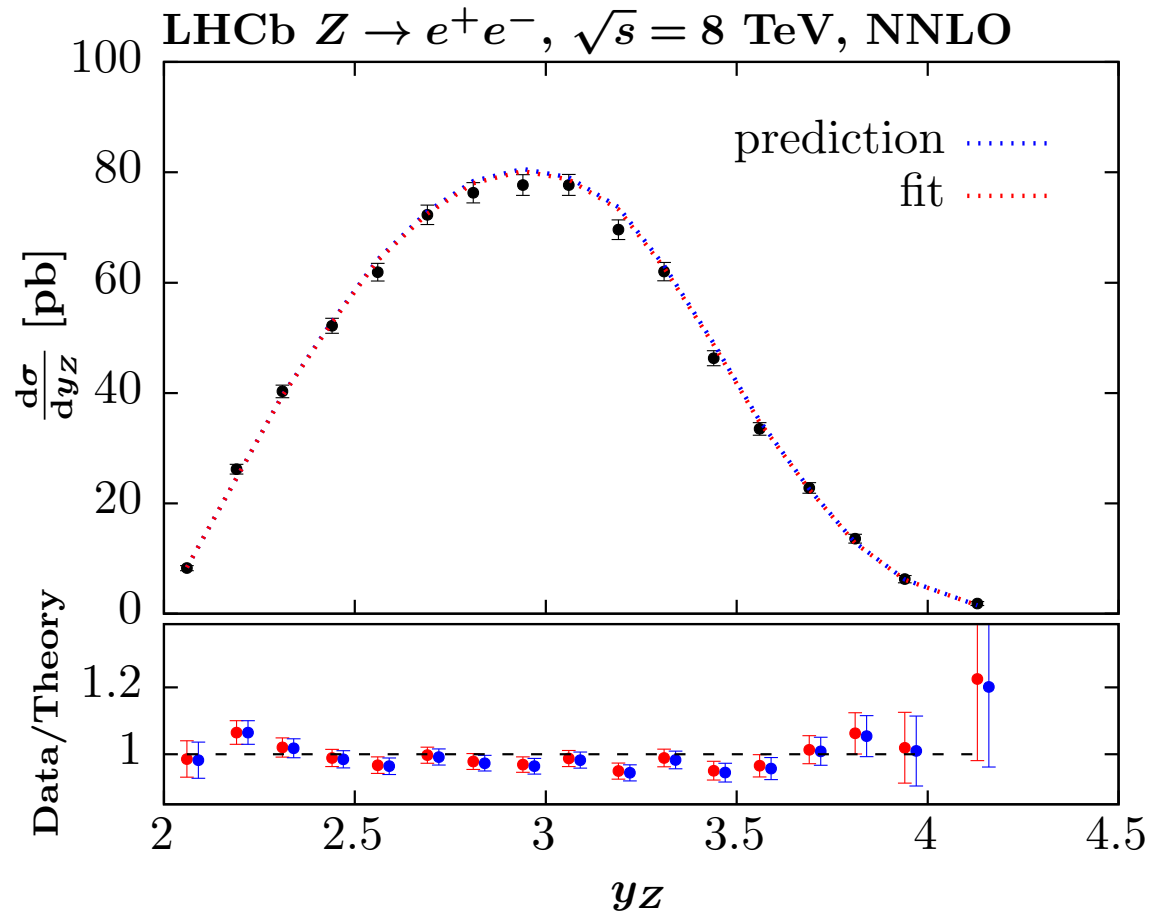


# New data on high rapidity $Z$ production at LHCb at 8 TeV.



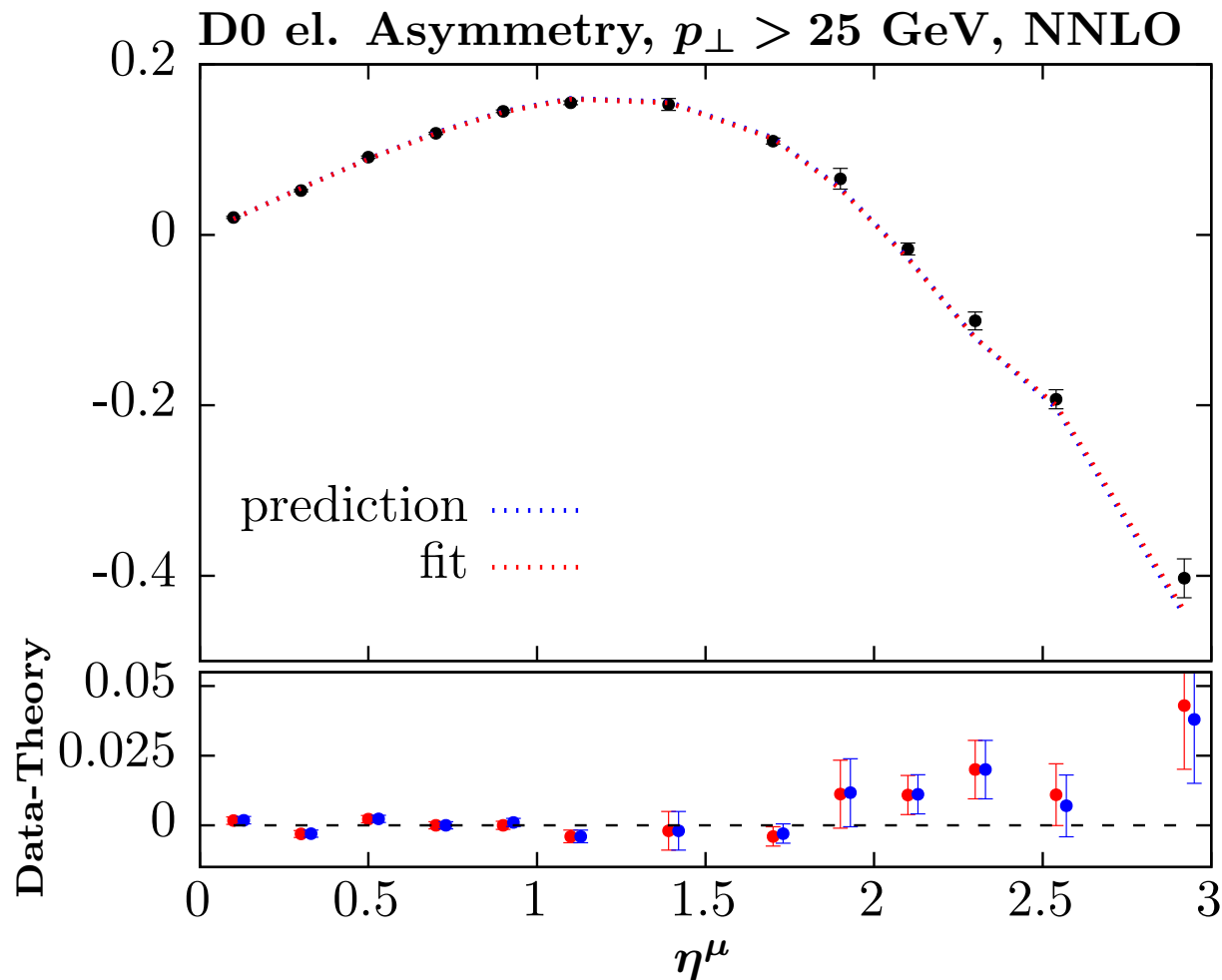
Same issue with lowest  $y_Z$  point. PDFs at moderate  $x$  for these points and well constrained by DIS data.

# New data on high rapidity $Z$ production at LHCb at 8 TeV with electrons.



No issue at lowest  $y_z$  with these data. Relatively large  $\chi^2$  only down to fluctuations.

## Good agreement with new D0 $e$ asymmetry data



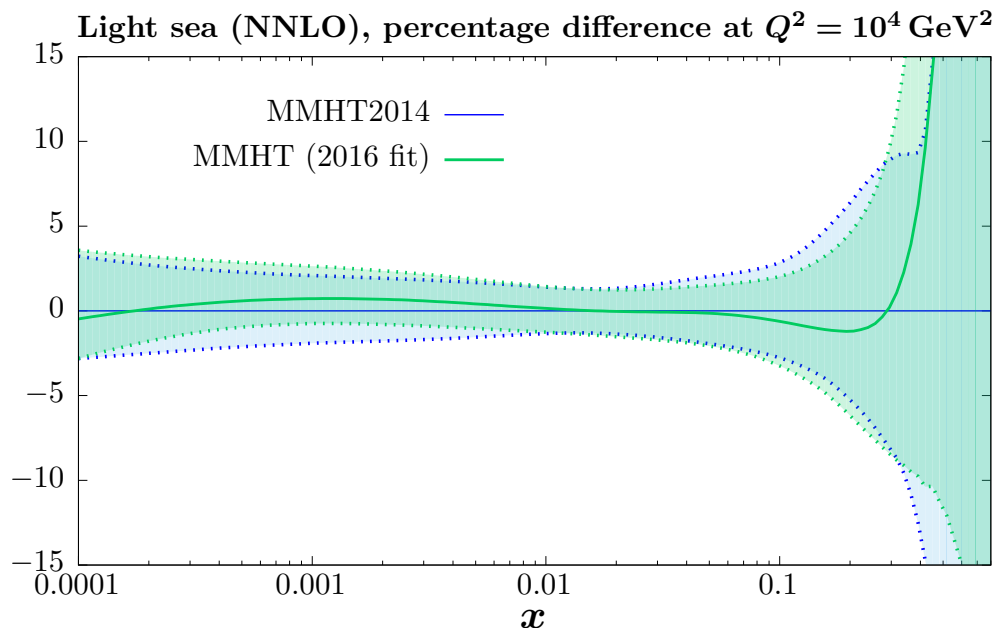
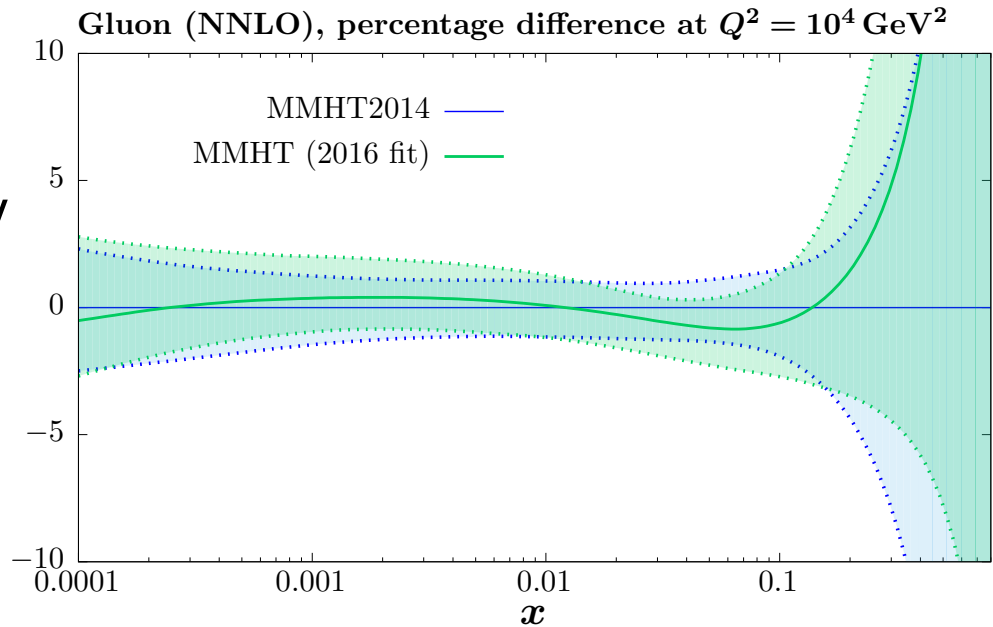
Slight undershooting at highest  $\eta_e$ . Implies slightly smaller down quark, but other data does not prefer this.

(Use the prescription for systematic uncertainties advocated in [Eur.Phys.J. C75 \(2015\) no.9, 458](#) for these and other **Tevatron** data.)

# Effect on PDFs

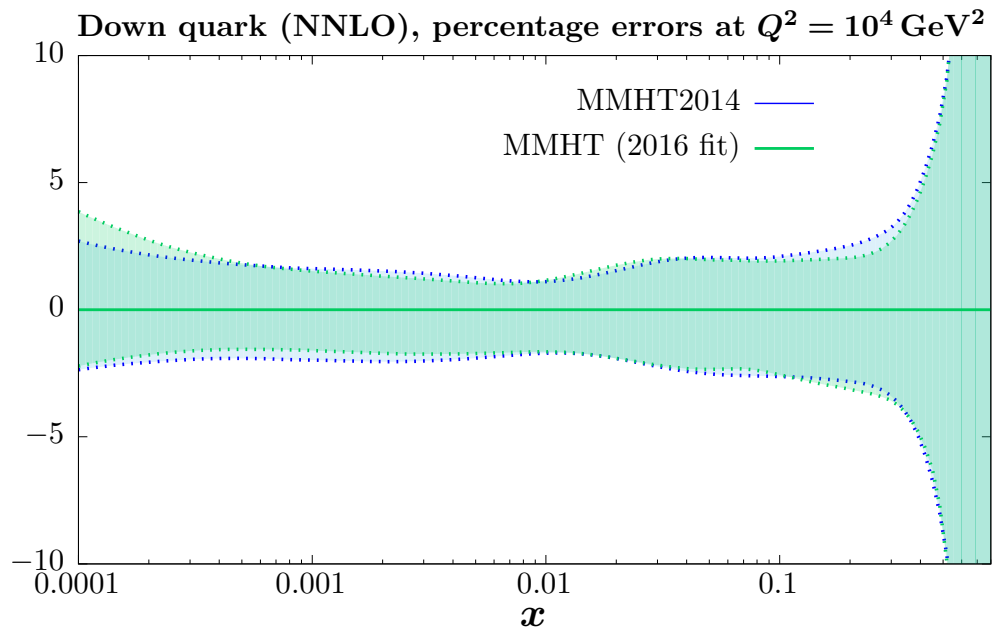
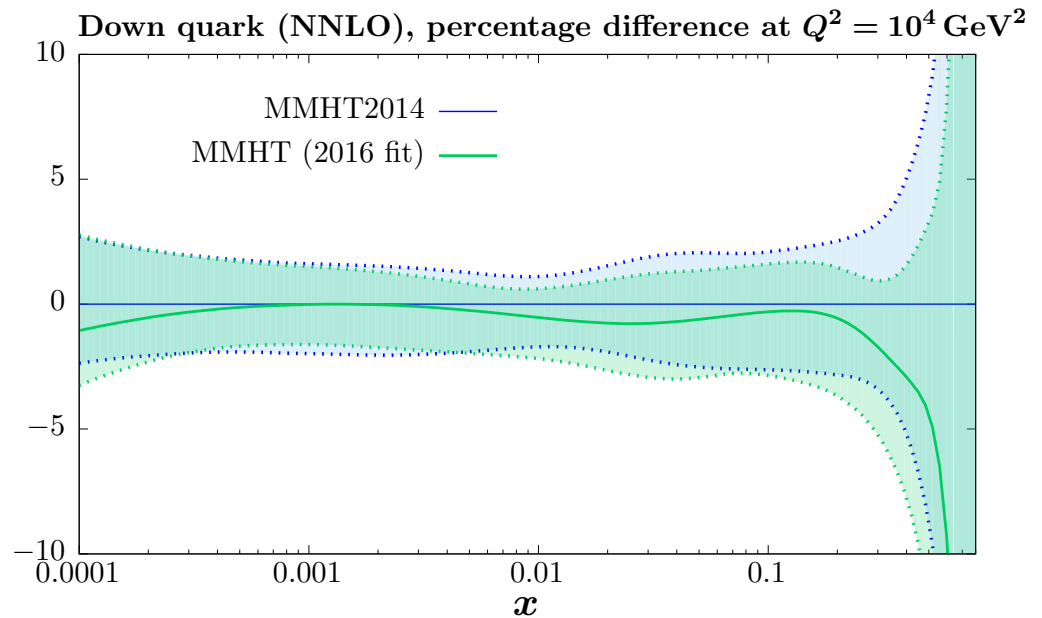
No significant change in gluon or light sea.

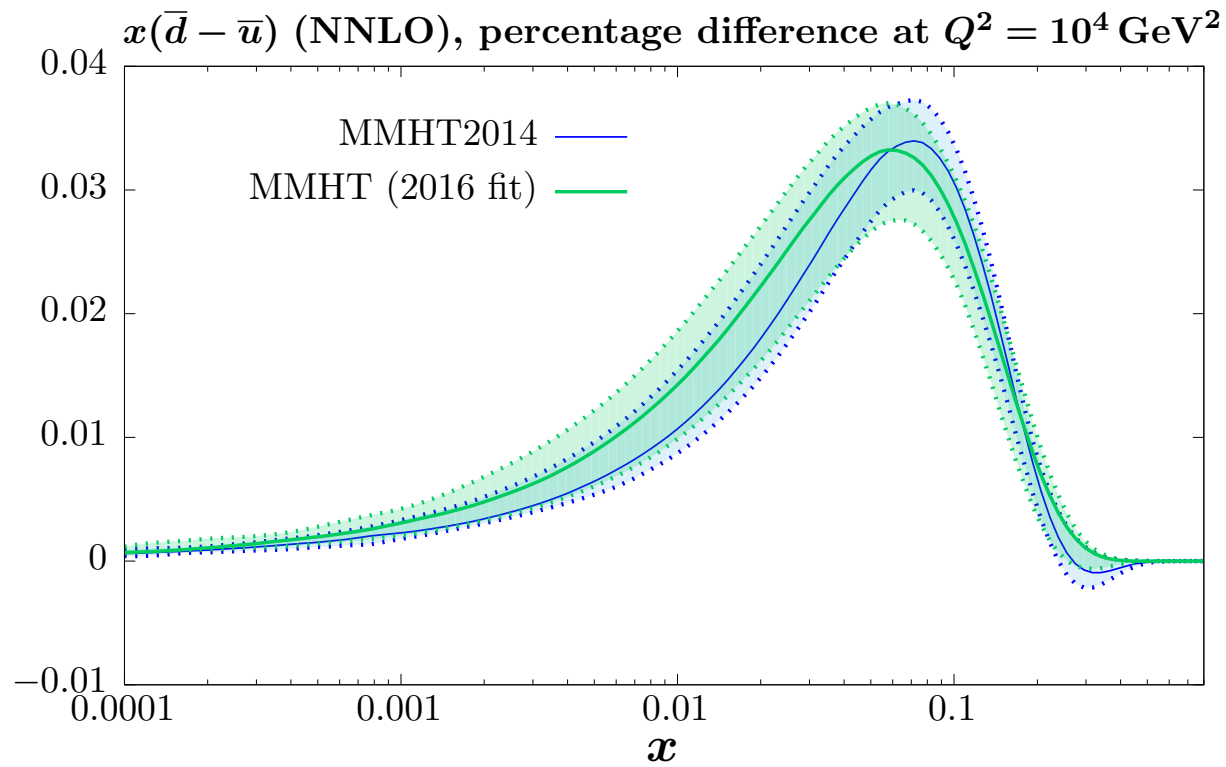
Small decrease in uncertainty in some small- $x$  regions due to new HERA data.



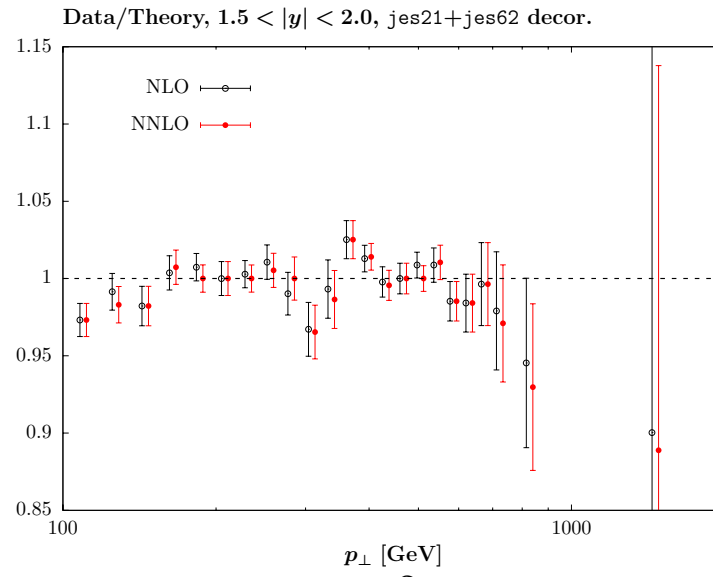
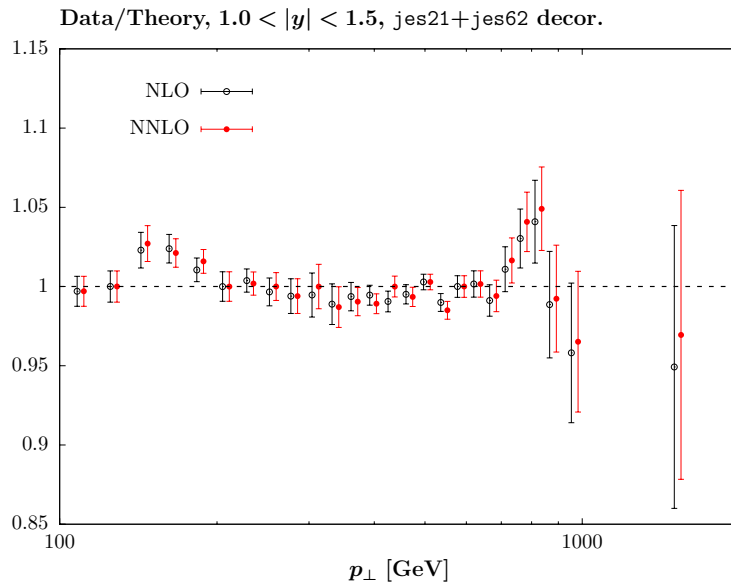
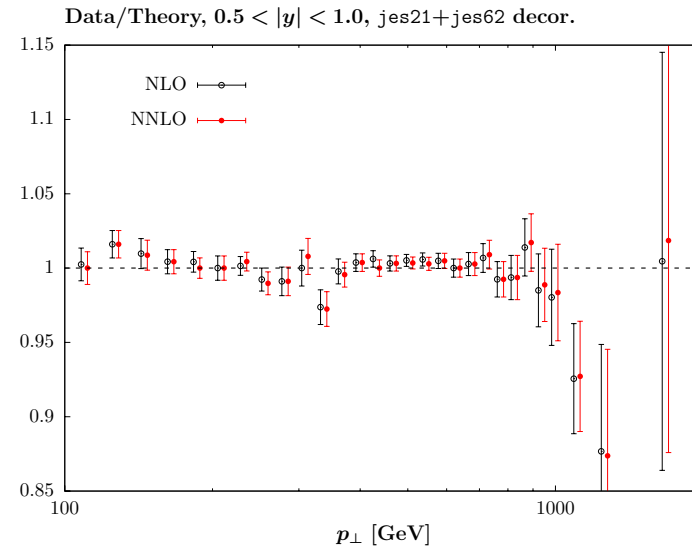
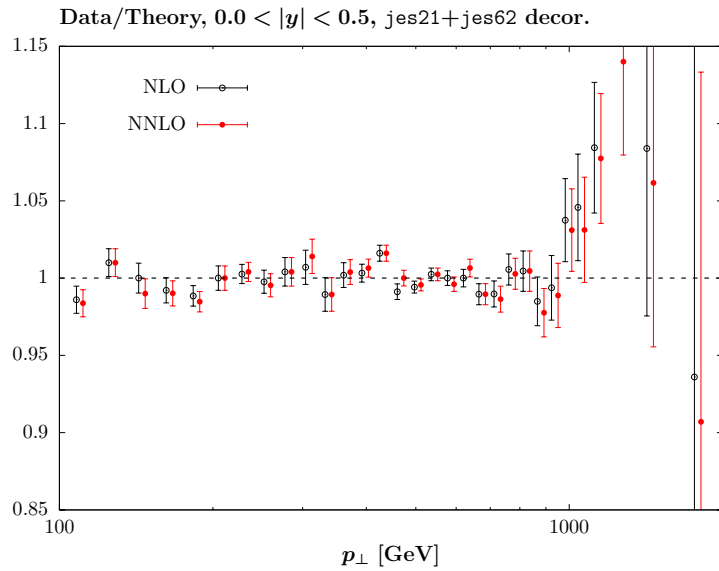
Small change in  $d$  at  $x \sim 0.01$  and some reduction in uncertainty.

Significant change in  $d$  at high  $x$  and some reduction in uncertainty for  $x \sim 0.2$ .

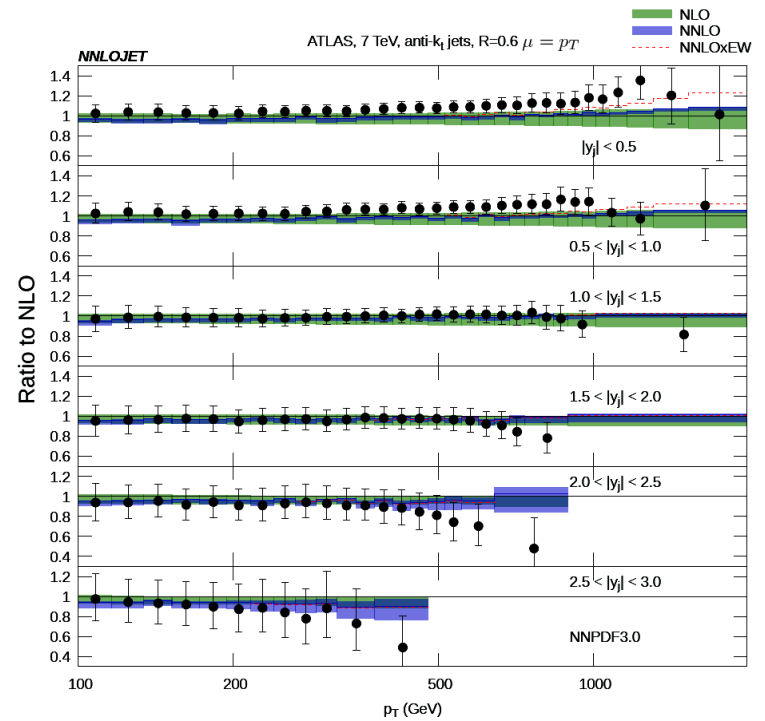
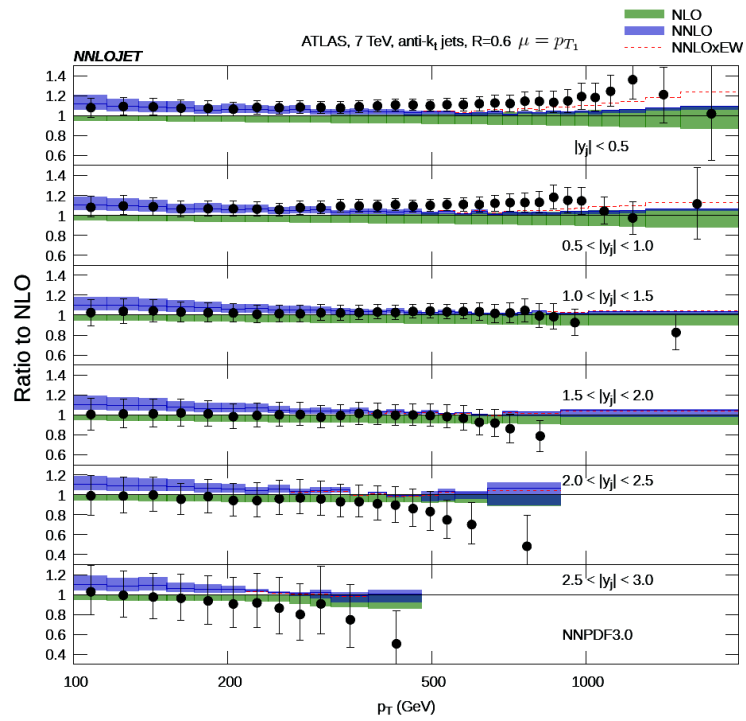




No major change in  $\bar{d} - \bar{u}$ , but even less inclination towards a change in sign at high  $x$  which was a feature of earlier sets.

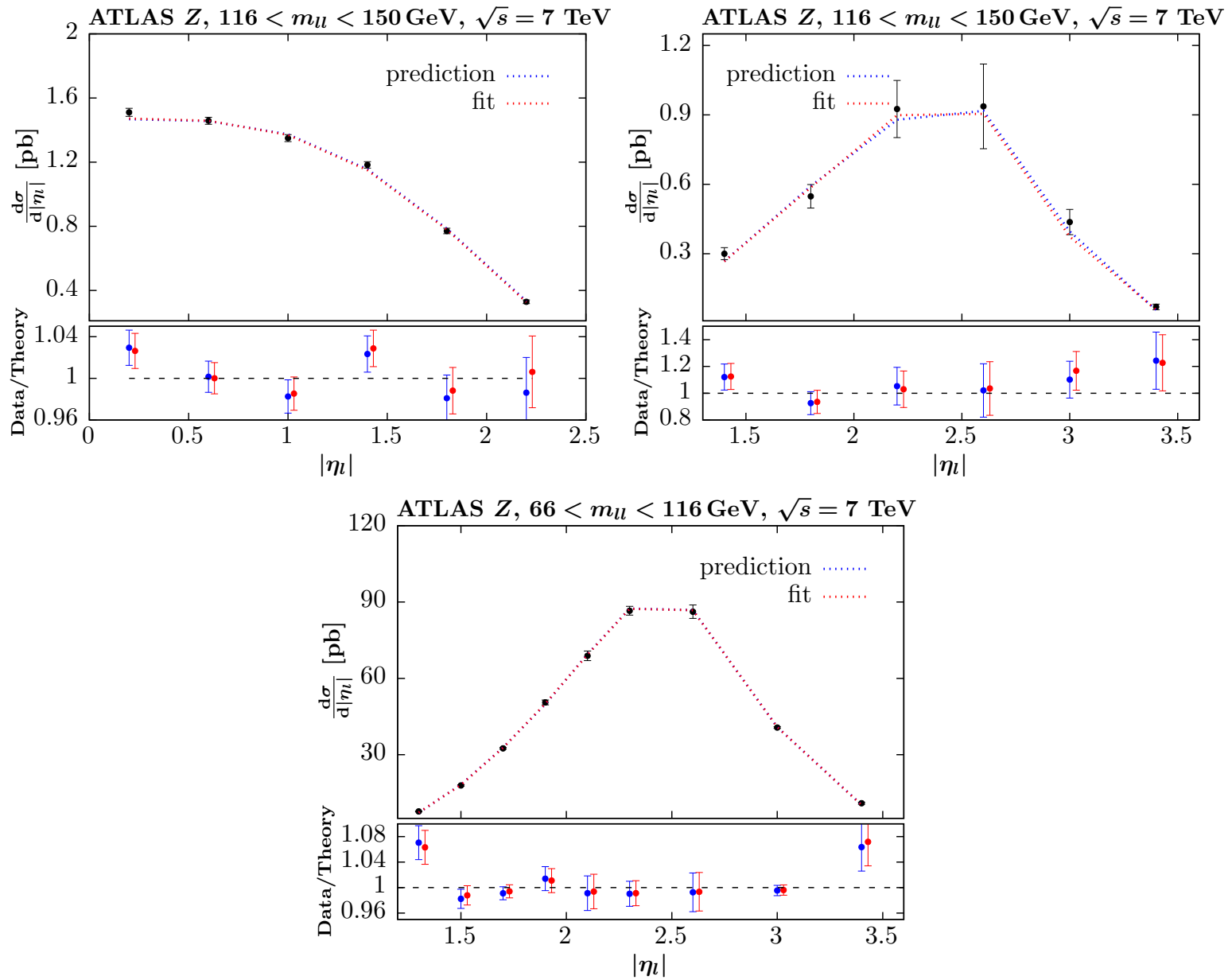


Fit quality at **NLO** and **NNLO** when jes21 and jes 62 decorrelated between  $|y|$  bin.



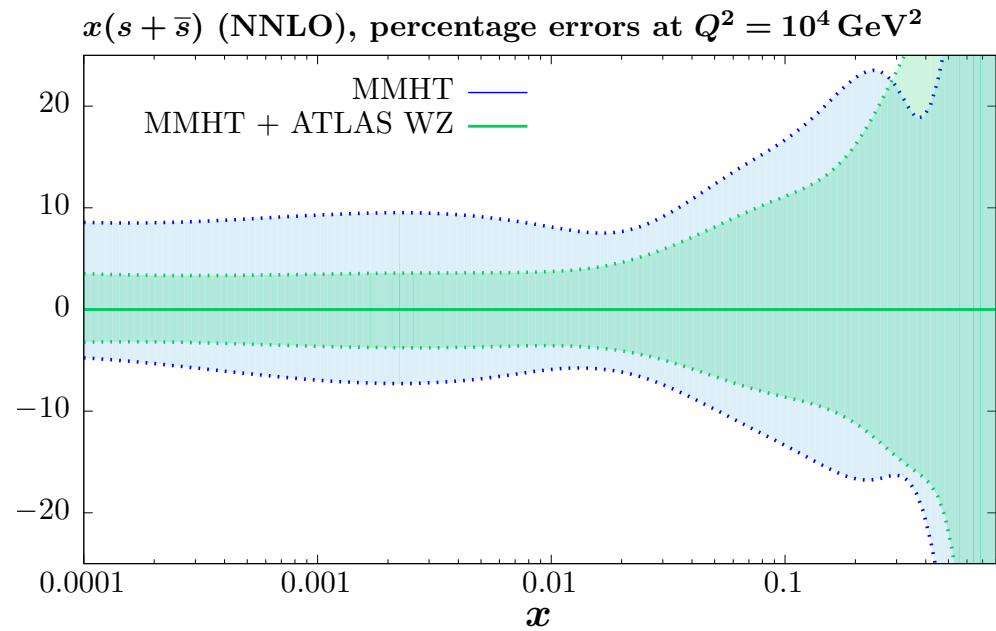
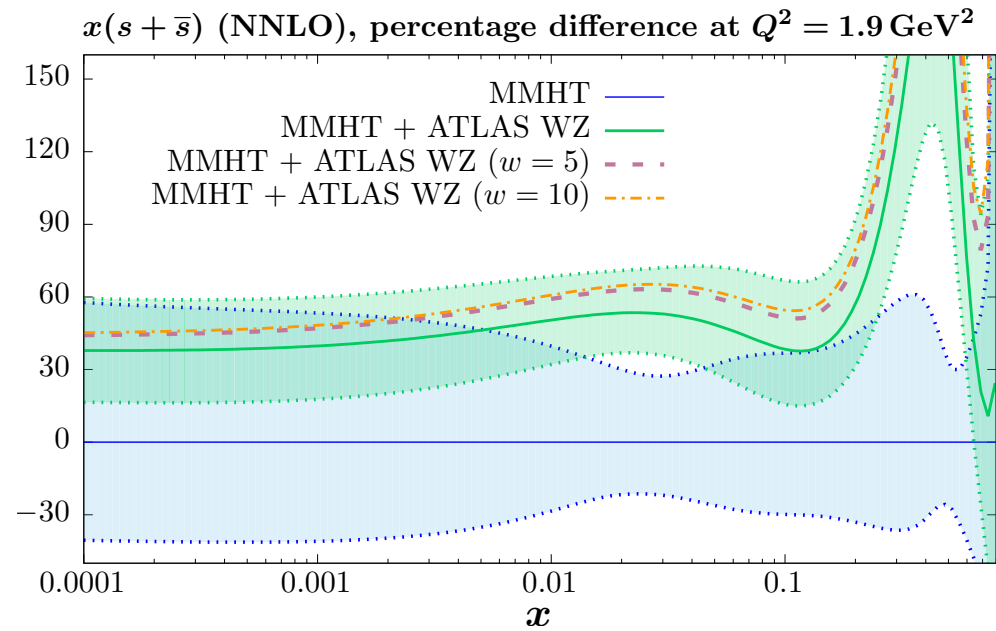


# Fit to new ATLAS $W, Z$ data



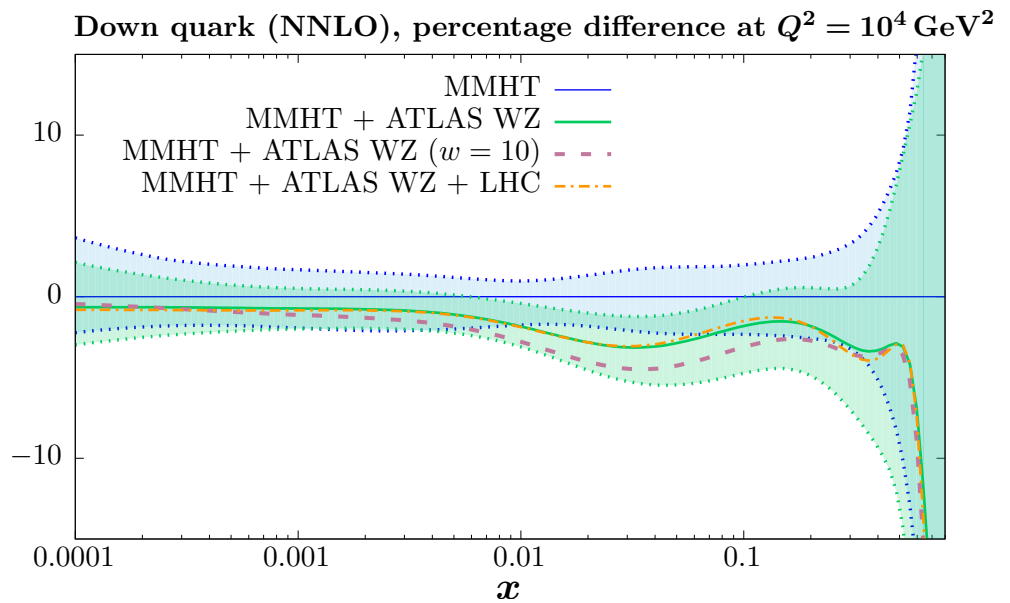
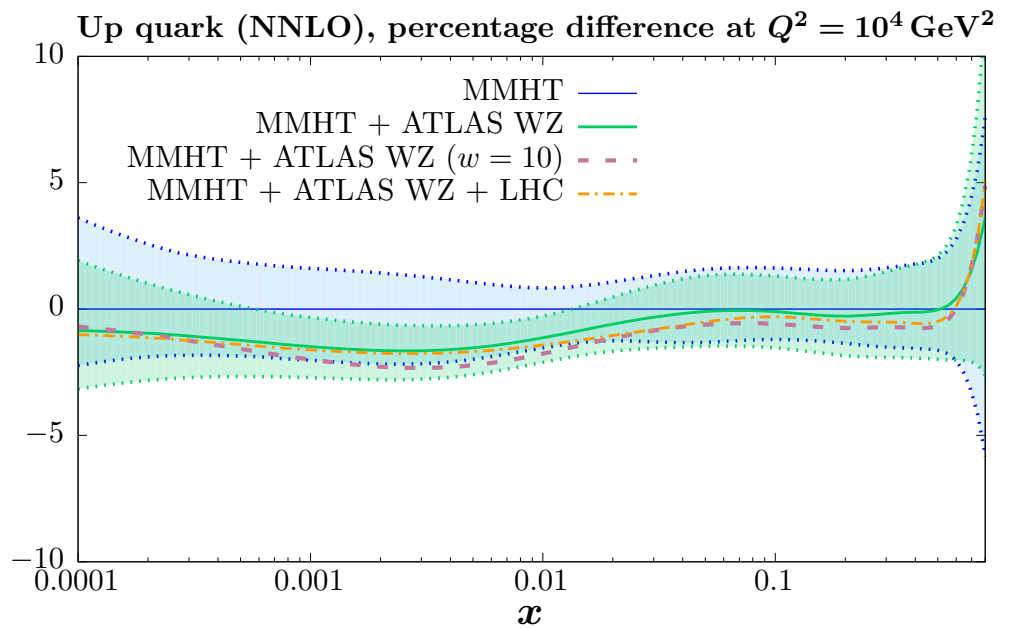
Strange quark at lower  $Q^2$  with addition of new ATLAS  $W, Z$  data.

Uncertainty reduction in strange.

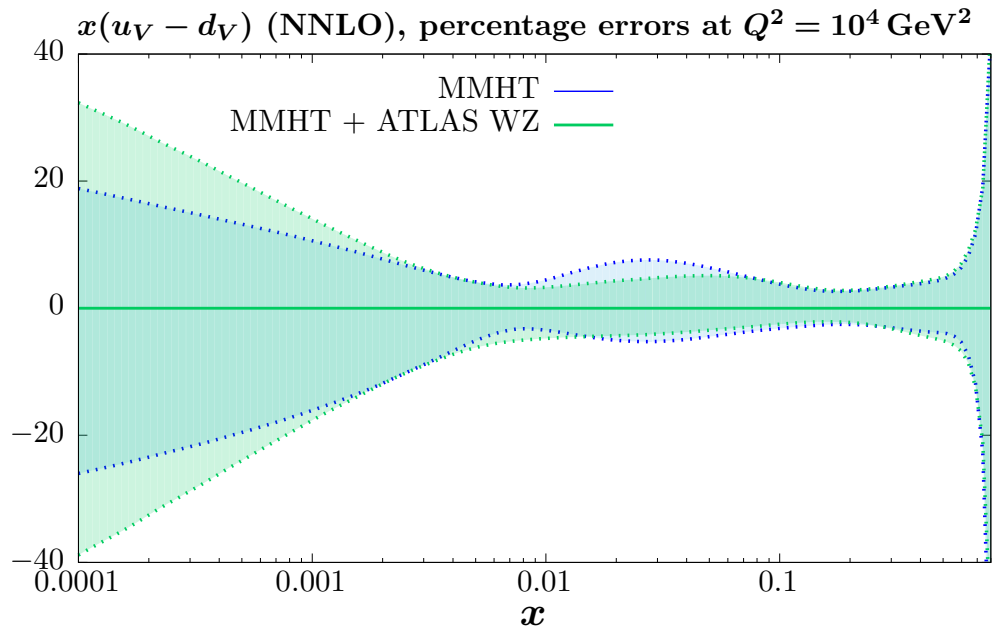


Slight decrease in low  $x$  up quark from increase in strange.

More pronounced effect in down quark.



Decrease in  $u_V - d_V$   
 uncertainty for  $0.01 < x < 0.1$ .



Improved constraint on  
 strange feeds into total  
 sea uncertainty.

